WG3: Summery Dusan Odstrcil and Fang Shen

1. Scientific Objective

(1)Provide global context for all CME events investigated by the ISEST team:

(2) Investigate processes of the CME initiation, heliospheric propagation, and CMEs interaction

(3) Develop tools to assist collaboration of numerical modelers, theoreticians, and observers

2. Scientific and Modeling Questions

- 2.1 Initiation of CMEs
- 2.2 Propagation and interaction of ICMEs
- 2.3 Impact on Geospace
- 2.4 Investigation of Mechanisms, Processes, and Forces (with WG2 Theory)
- 2.5 Forecasting the CME Arrival and Impact (with WG4 Campaign Events)

3. Future Plan and Action Items

- In the next 16 months (by the next ISEST workshop), , we will:
- (1) Make simulation of the ISEST Event Periods (with WG1) for all (ENLIL model), and for major CME events (COIN-TVD model)
- (2) Investigate Mechanisms, Processes,
- Forces, Energies, and Interactions (with WG2 Theory)
- (3)Forecast and validate the CME Arrival and Impact (with WG4 Campaign Events)
- (4)Provide standard outputs for observers (synthetic in situ plots and data, height-time plots and data of shock and leading edge of the driver)

4. Overview of the models used in ISEST

- (1) ENLIL (sumerian god of wind and storms) model: 1RS-21.5RS, WSA (Wang-Sheeley-Arge), transients: hydrodynamic ejecta (Cone or Rope geometry), heliosphere >21.5 Rs - 3D MHD model
- (2) COIN-TVD(Corona-Interplanetary Total Variation Diminishing): 1Rs-beyond 1AU: 3D MHD model, transients: magnetized plasma blob model
- (3) H3DMHD: 1RS-21.5RS, HAF (Hakamada-Akasofu-Fry) model >21.5 RS, 3D MHD model
- (4) SWMF (Space Weather Modeling Framework, BATSRUS): 1Rs-beyond 1AU: 3D MHD model, transient: analytic magnetic flux rope