

Modeling and Data Analysis Tools in support for Solar Orbiter and Parker Solar Probe science

R. Pinto (1), IRAP/CDPP and MADAWG teams

(1) IRAP / CNRS-OMP / U. Paul Sabatier, Toulouse, France

Connectivity Tool

- . past and future spacecraft – solar surface connectivity
- . Solar Orbiter, Parker Solar Probe, etc.

Accurate background solar wind simulations

- . physics based modeling, alternative to semi-empirical (WSA)
- . data driving, full set of background wind properties (speed, density, etc)
- . flexible setup: full 3D, plane-of-sky, orbit, etc.
- . a two step problem: corona + heliosphere
- . fit into/complement existing SWx tool ecosystem (cf. swe.ssa.esa.int, CCMC)
- . Interface with ConnectTool, PropagationTool, AMDA, etc.

MADAWG (Modeling and Data Analysis WorkGroup)

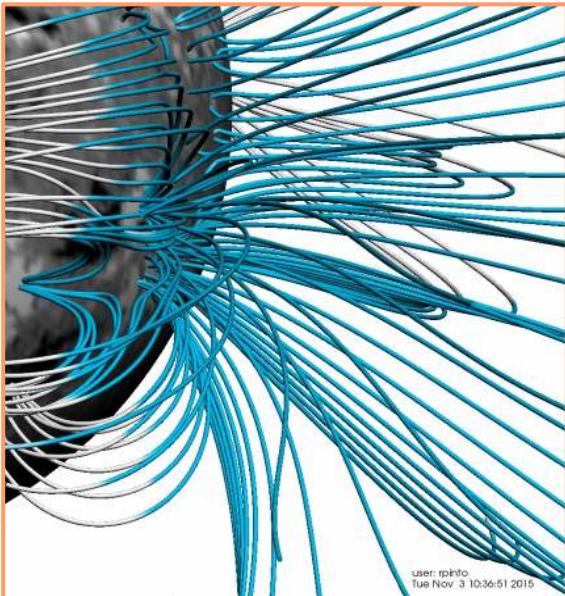
STORMS (Solar Terrestrial ObseRvations and Modeling Service)

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rui.pinto@irap.omp.eu



<http://www.helcats-fp7.eu/>
<http://stormsweb.irap.omp.eu>

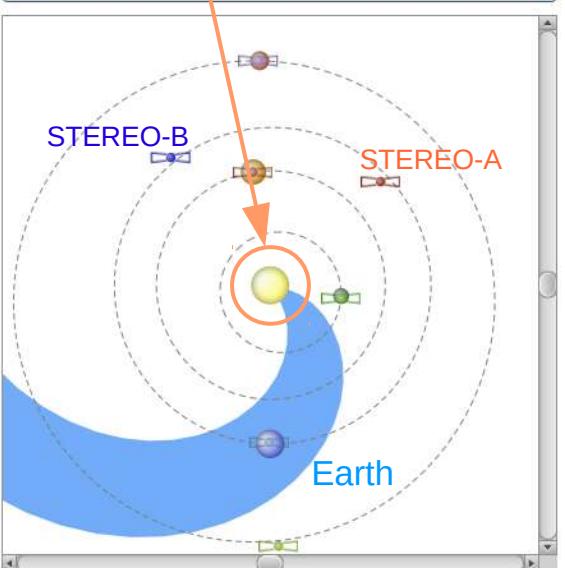
Connectivity Tool



- **Paths and delays** of propagation Sun – spacecraft
(wind flows, CME, shocks, energetic particles)



- **Use many models:** forecast redundancy
(but keep it simple!)



- **Planning and management of campaigns**
exploit synergies between multiple s/c.

- **Scientific exploitation**
past data: post-event analysis relating in-situ ↔ remote,

Connectivity Tool

Layout of the connectivity tool

1. Chose mode of operation

Post-event analysis

Forecast

2. Choose date or s/c position

3. Trace down magnetic field-line

i) Parker spiral or heliospheric model

use measured wind speed

use estimated wind speed

ii) Coronal field reconstruction

(PFSS, NLFFF, MHD, ...)

use HMI magnetograms

use ADAPT magnetograms

4. Find connectivity points at the surface

Take into account propagation delays (wind, particles)

Plot maps at slow/fast wind / SEP launch times + target time

Compare instrument FOV, estimate uncertainties

Connectivity Tool

Home Help Contacts

Prototype webpage <http://storms-connectsolo.irap.omp.eu/>

Solar Orbiter / Solar Probe Plus Connectivity Tool

Select date/time
at spacecraft

Date

2020-05-01 00:00

Select coronal
model

Coronal magnetic field

PFSS

Mode

SCIENCE

Select Science/
Forecasting Mode

Interplanetary magnetic field

PARKER

Select interplanetary
model

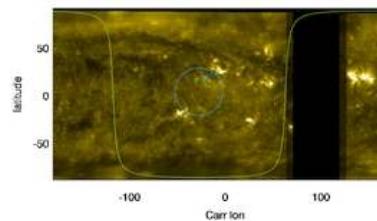
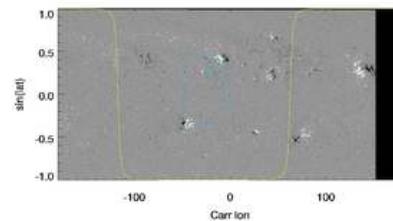
SEARCH

Time reference

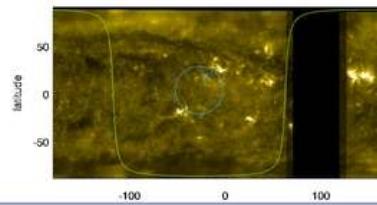
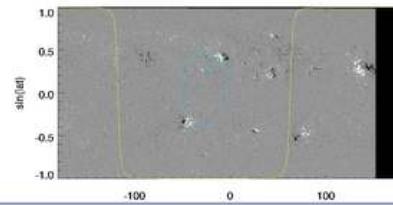
SDO/HMI

SDO/171

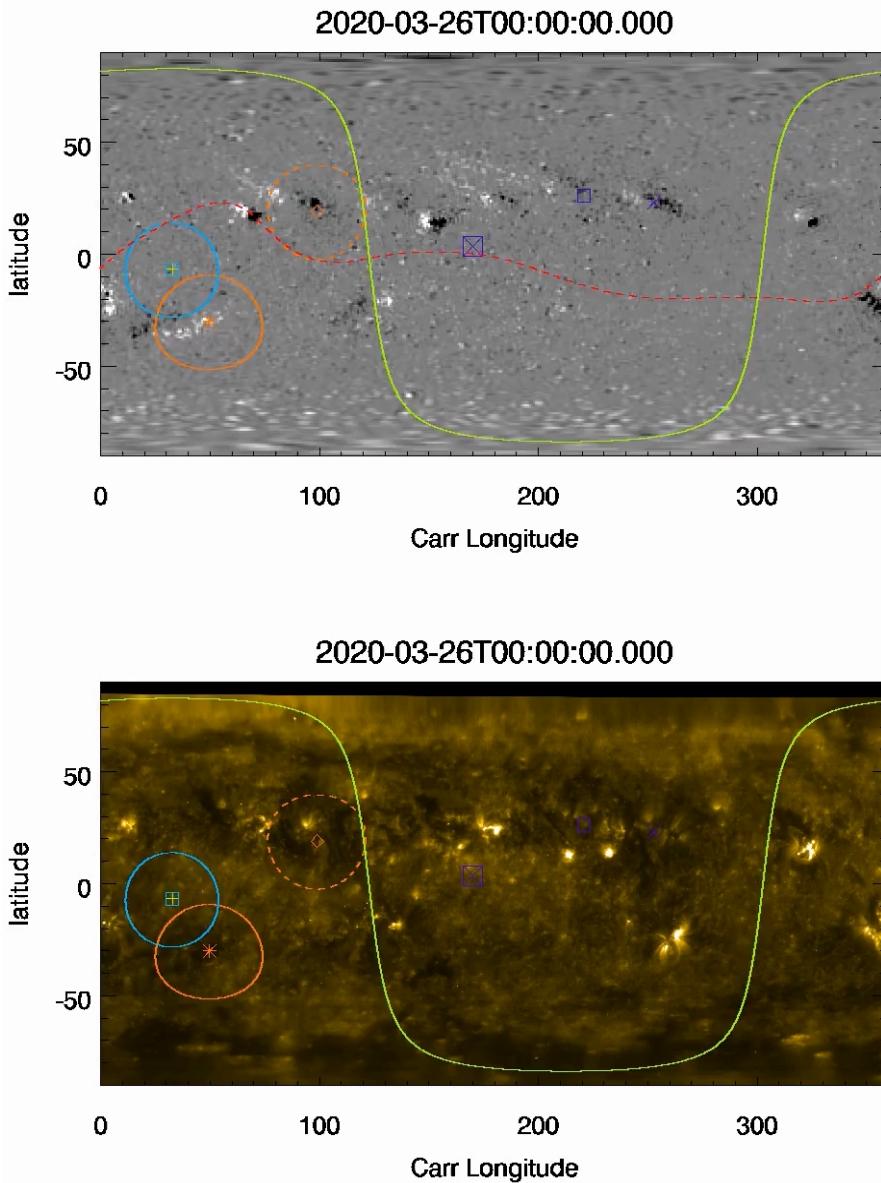
Carrington maps
at plasma impact ➤ Spacecraft
time



Carrington maps
at plasma release
time ➤ Sun



Solar Orbiter CONNECTIVITY TOOL



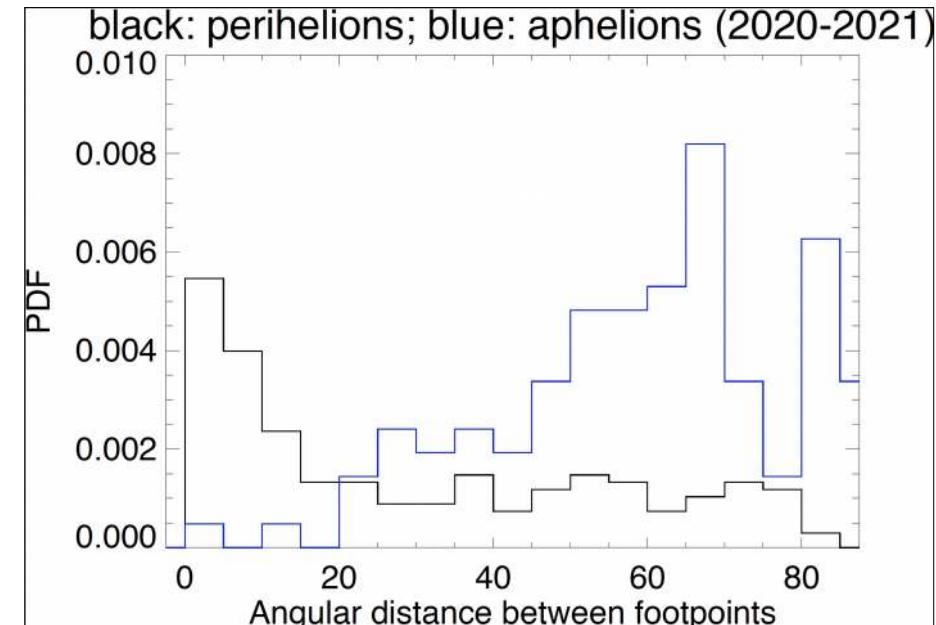
HMI magnetograms (time-evolved) EIT 171 Å

Limb as seen from Solar Orbiter
Neutral line

EUI FOV centered on Solar Orbiter position
EUI FOV centered on Solar Orbiter footpoints
(slow wind and fast wind)

Position and footpoints of Parker Solar Probe

Solar Orbiter footprint separation



Connectivity maps

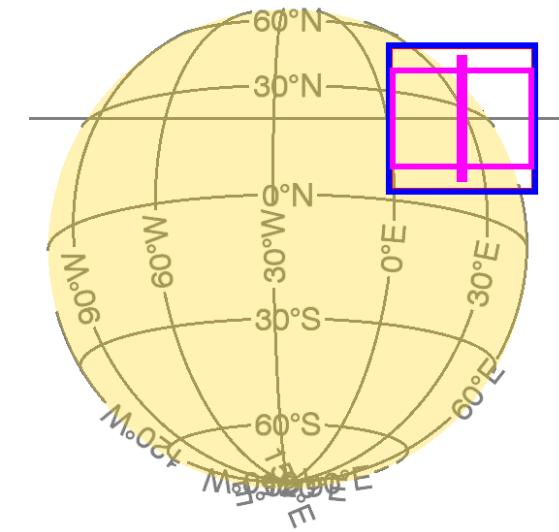
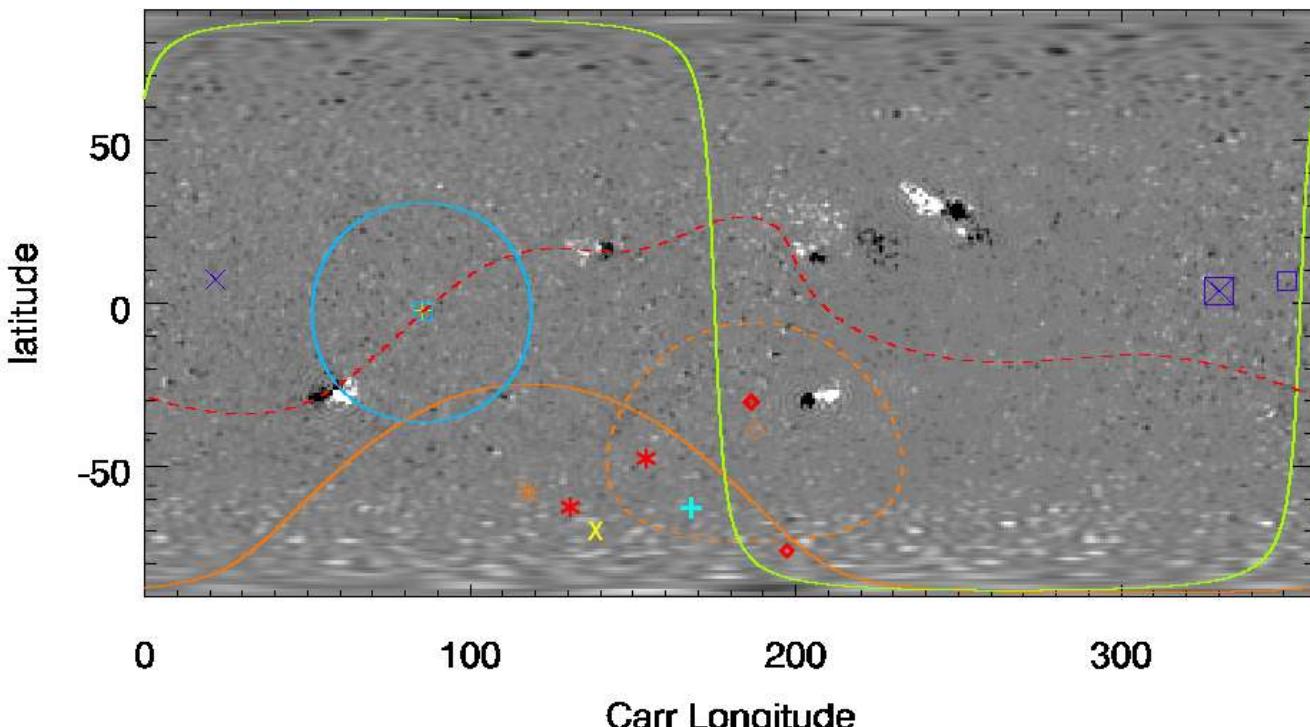
Forecasting: testing different approaches

Corona:

ADAPT+PFSS, ADAPT+NLFFF (maybe)

Heliosphere:

PARKER, EUHFORIA, ENLIL



End data product:

probability distribution of connectivity
(positions + sizes of distributions vs. FOV)

Souces of scatter :

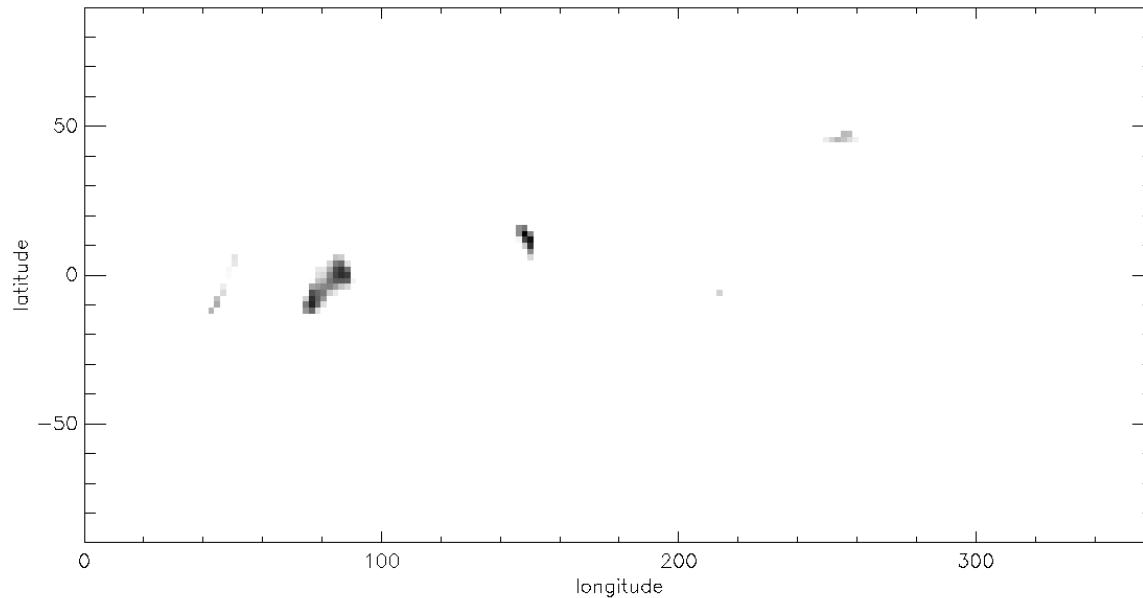
- different magnetogram sources
- magnetic field extrapolation methods
- assumed/forecasted wind profile

Tasks :

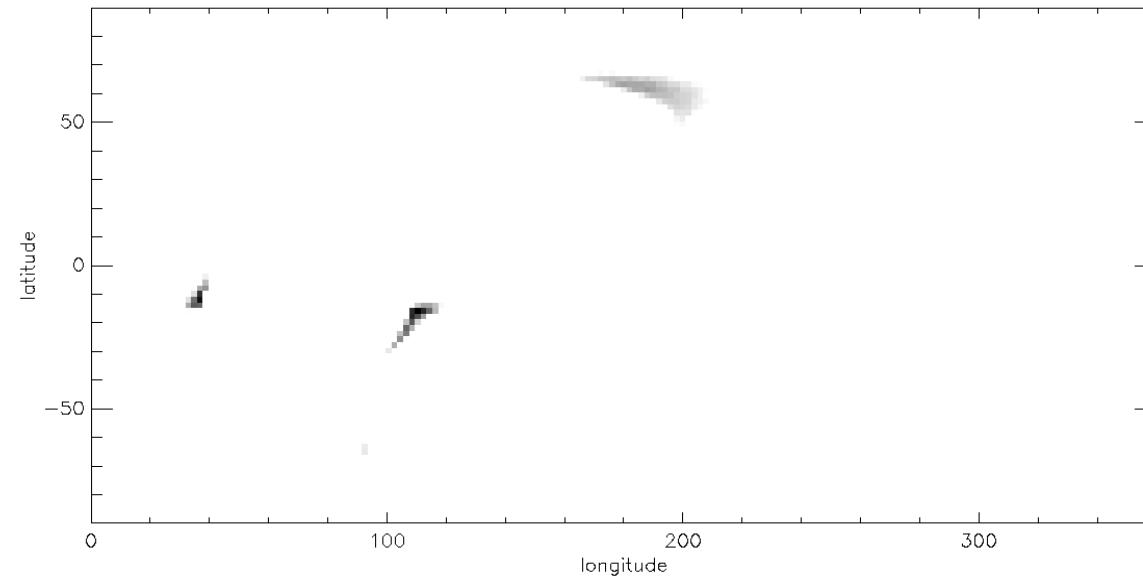
- cross-calibrate different magnetograms
(cf. Riley et al, 2014, Linker et al 2017)
- integrate fresh SolO/PHI data into the modelling
- synoptic → synchronic maps (e.g, Henney, et al)
- integrate other models from the community

Connectivity: mapping the uncertainty

Carrington map, conn. prob. density at surface



Carrington map, conn. prob. density at surface

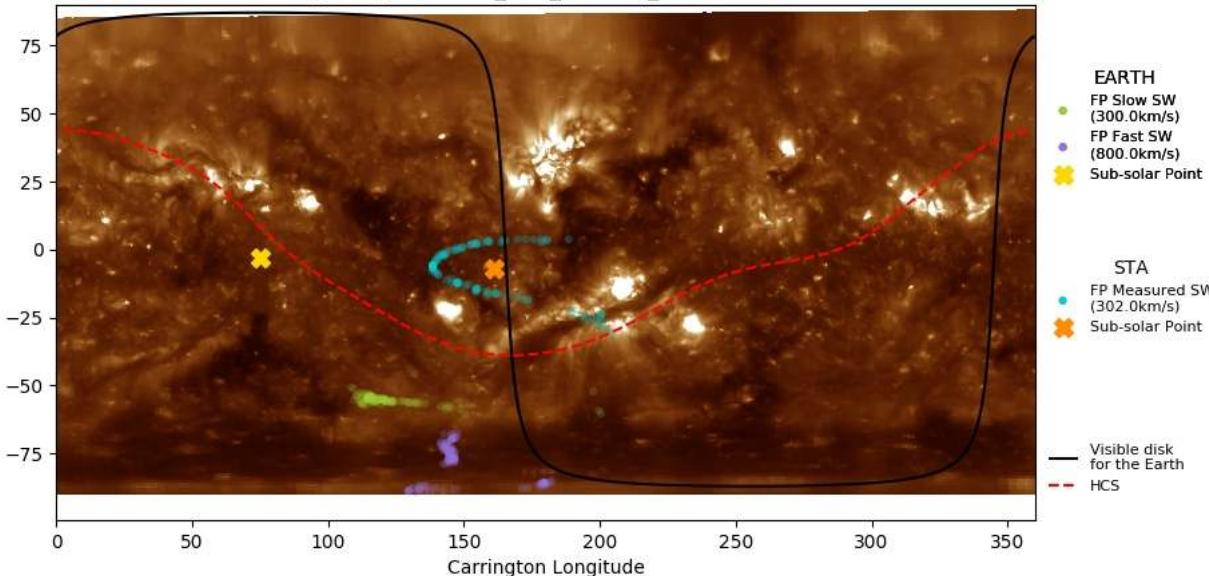


Method :

- 1) run an *ensemble* of connectivity estimations
different models, different parameters
- 2) assign φ and θ uncertainty at R_{ss} for each estimation
typically, a bi-gaussian with $\sigma_\varphi > \sigma_\theta$; $P_{ss}^i(t, \varphi, \theta)$
- 3) map probability density function down to $1 R_s$
take mag. field topology into account,
respect magnetic path bifurcations,
- 4) combine all into one probability density function
keep a minimal amount of discriminant info
- 5) pick reduced set of coordinates for science ops.

Connectivity: mapping the uncertainty

2011-01-01T00:00:00.000 CR2105
Magnetogram: wso_wso_CR2105_20101224.txt



Method :

1) run an *ensemble* of connectivity estimations
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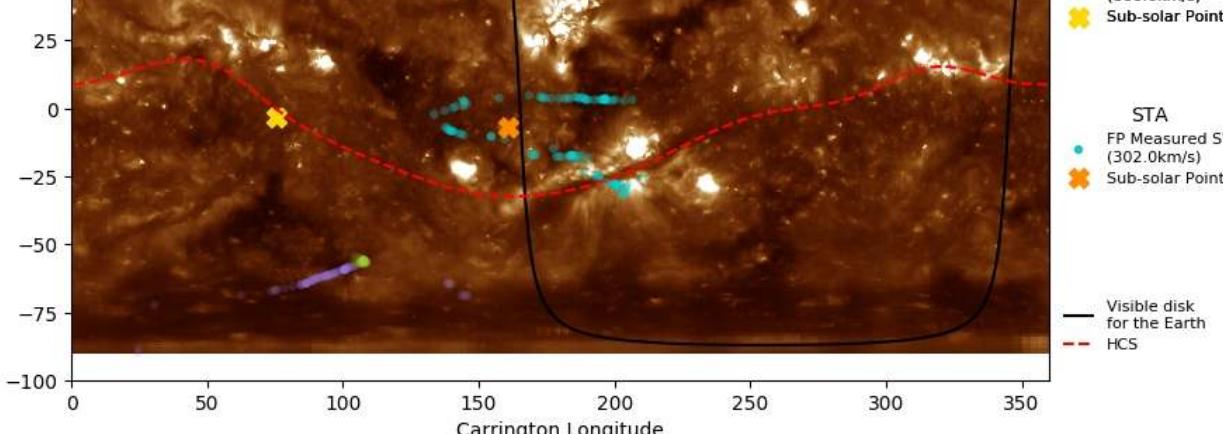
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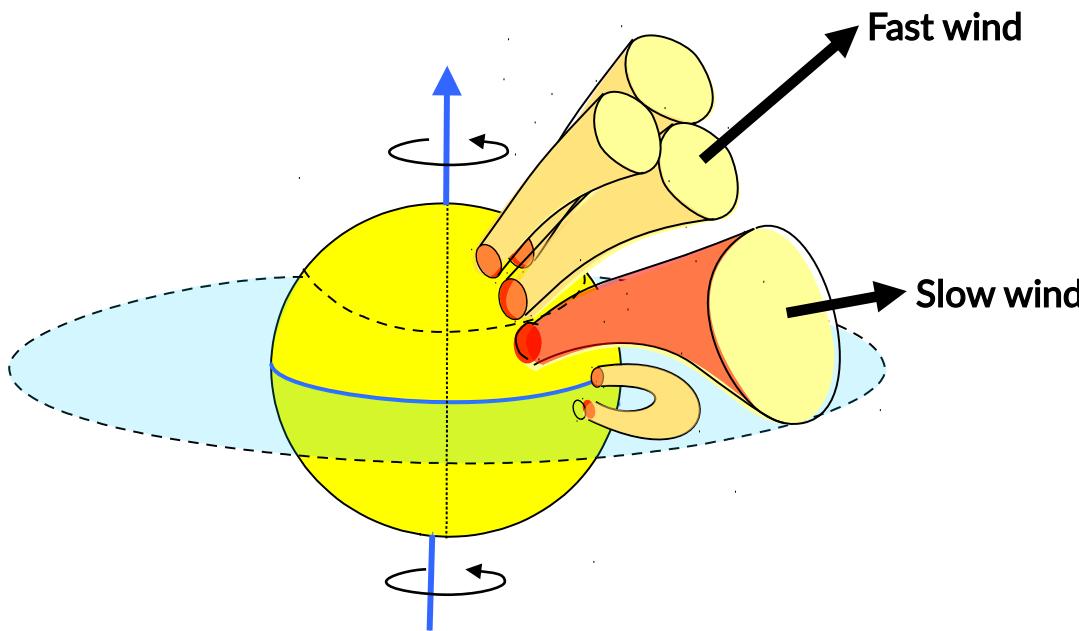
5) pick reduced set of coordinates for science ops.

2010-12-27T08:09:45.037 CR2105
Magnetogram: nso_gong_CR2105_20101224.fits



Estimating the solar wind speed distribution

Semi-empirical scaling laws



WSA (Wang-Sheely-Arge)

$$V_{wind} = 265 + \frac{1.5}{(1 + f_{ss})^{1/3}} \times \left[5.8 - 1.6 \exp \left[1 - \frac{\theta_b^3}{7.5^3} \right] \right]^{3.5} \text{ km s}^{-1}$$

f_{ss} : total flux-tube expansion ratio (Wang, 1995; Velli 2013)

θ_b : distance to coronal hole boundary

Coronal hole area – wind speed relation

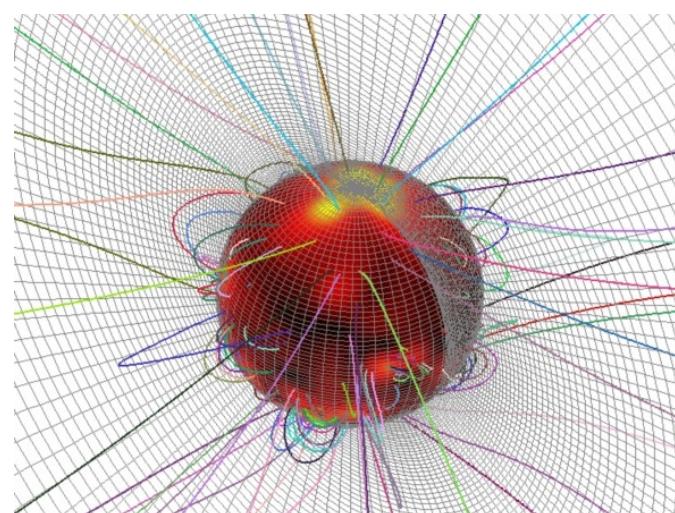
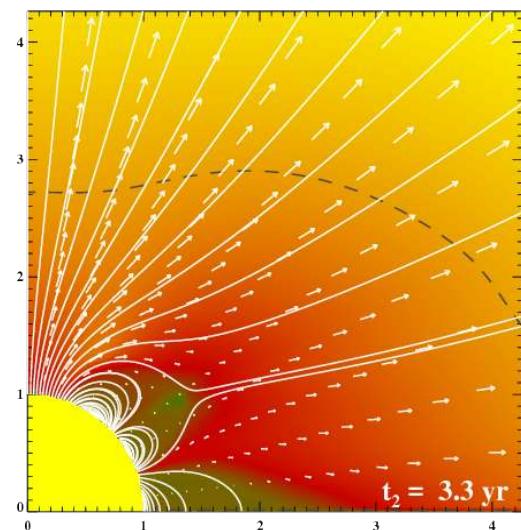
$$V = (80 \pm 2)A + 426 \pm 5.$$

V: wind speed at 1 AU

A: total CH area

(Nolte et al 1976, Vršnak, et al 2007, Tokumaru et al 2016)

Global MHD corona and wind models



More physics, but higher complexity

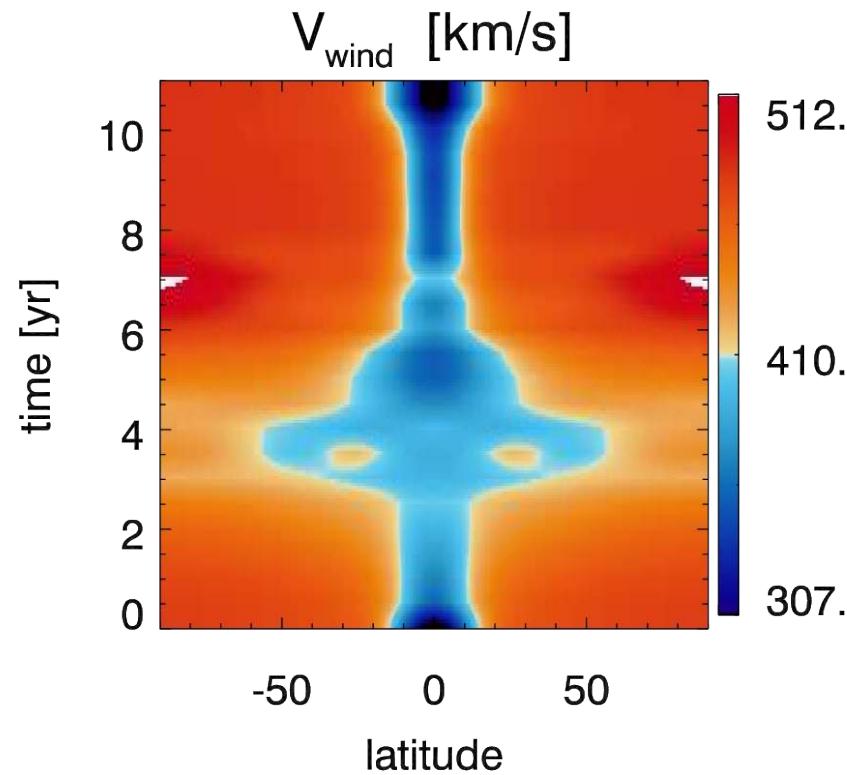
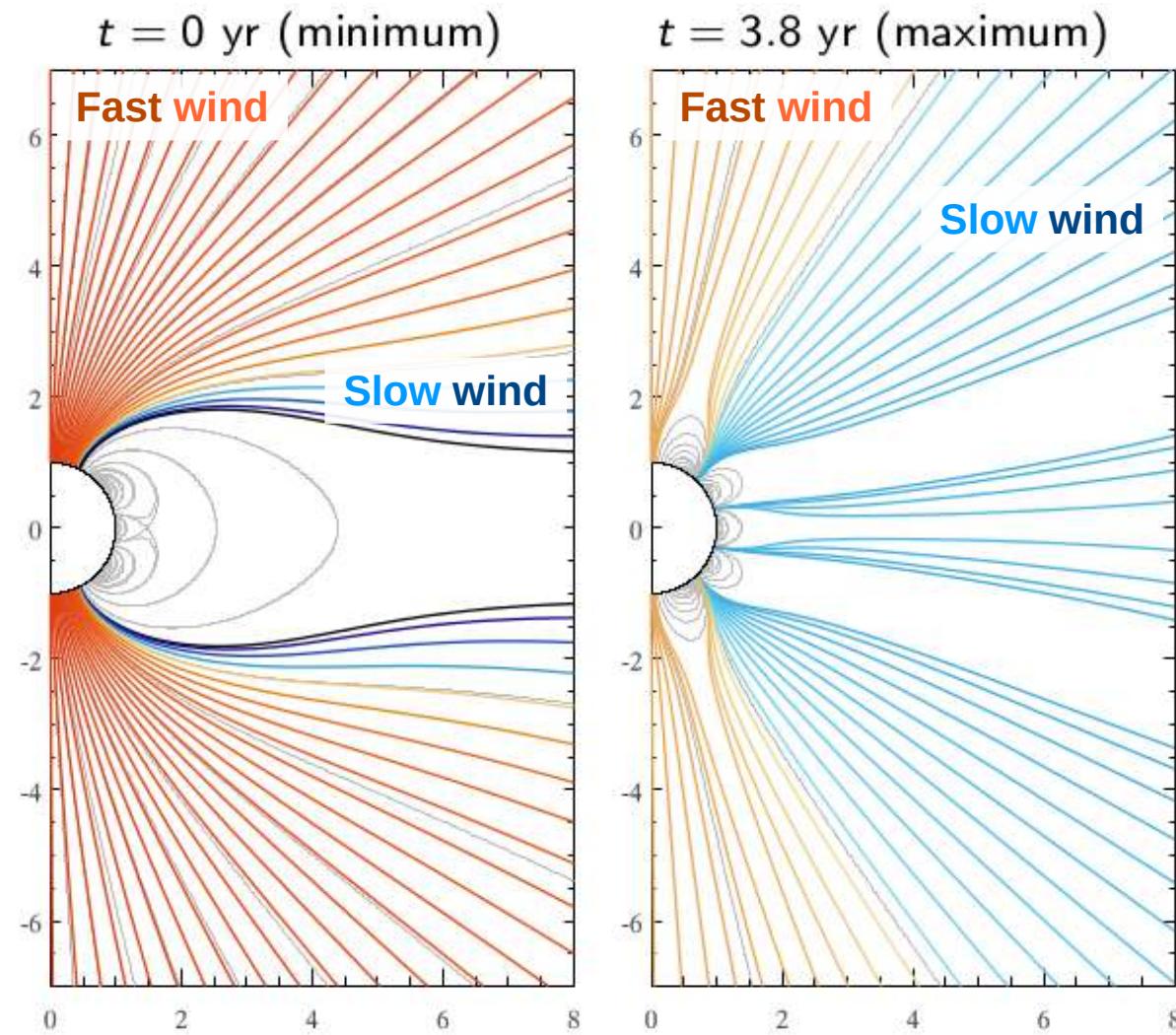
Full magnetic geometry
(but other simplifications)

CPU-heavy
(especially in the corona)

Mikić et al (2011), Gressl et al (2014),
Pinto et al (2011, 2013),
van der Holst et al (2015), +

Solar wind speed and flux-tube geometry

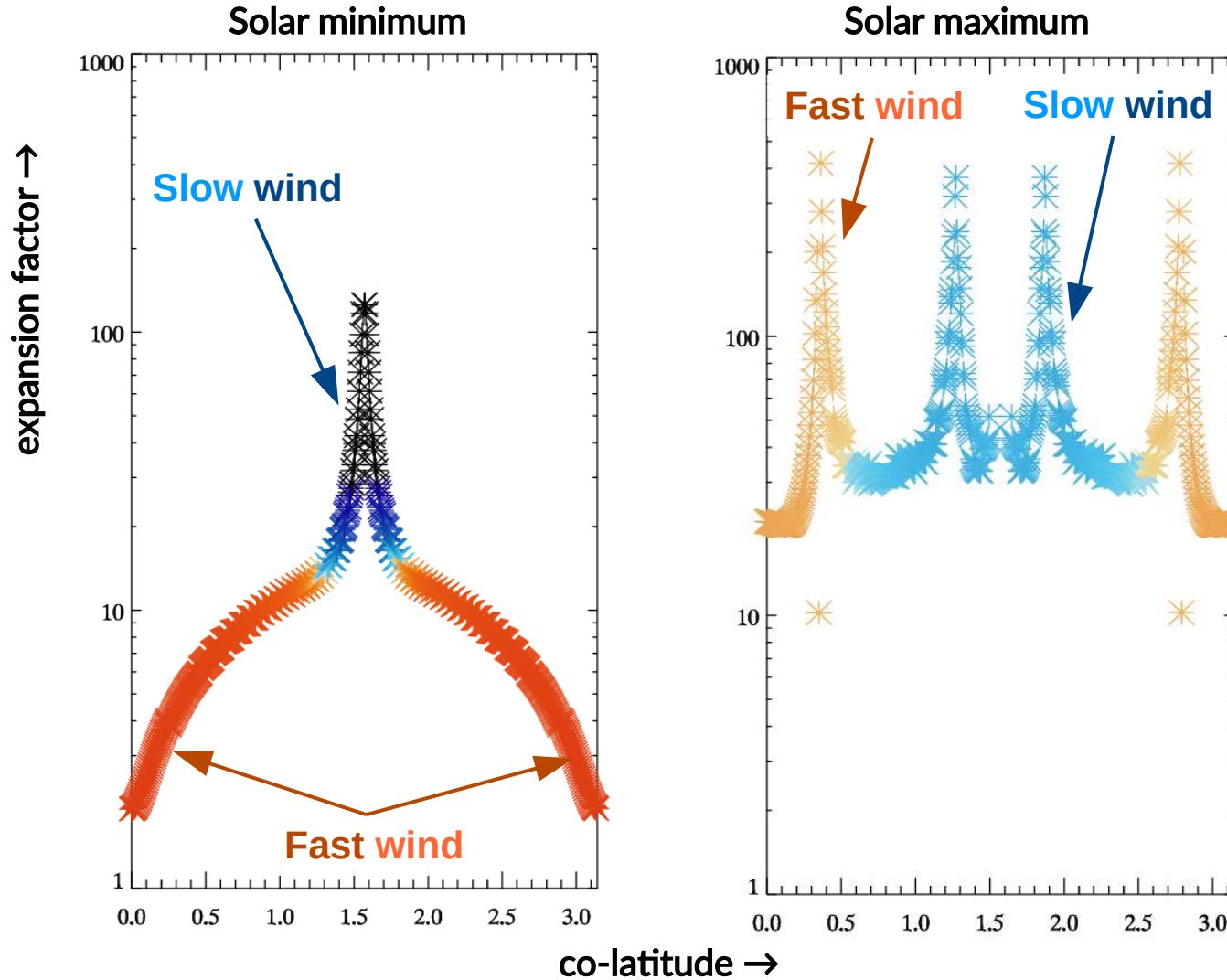
Global MHD simulations of the solar dynamo, corona and solar wind (11 yr cycle)



Wind speed distribution
↔
Magnetic geometry

Solar wind speed and flux-tube geometry

Expansion factor and wind speed



Flux-tube geometry:

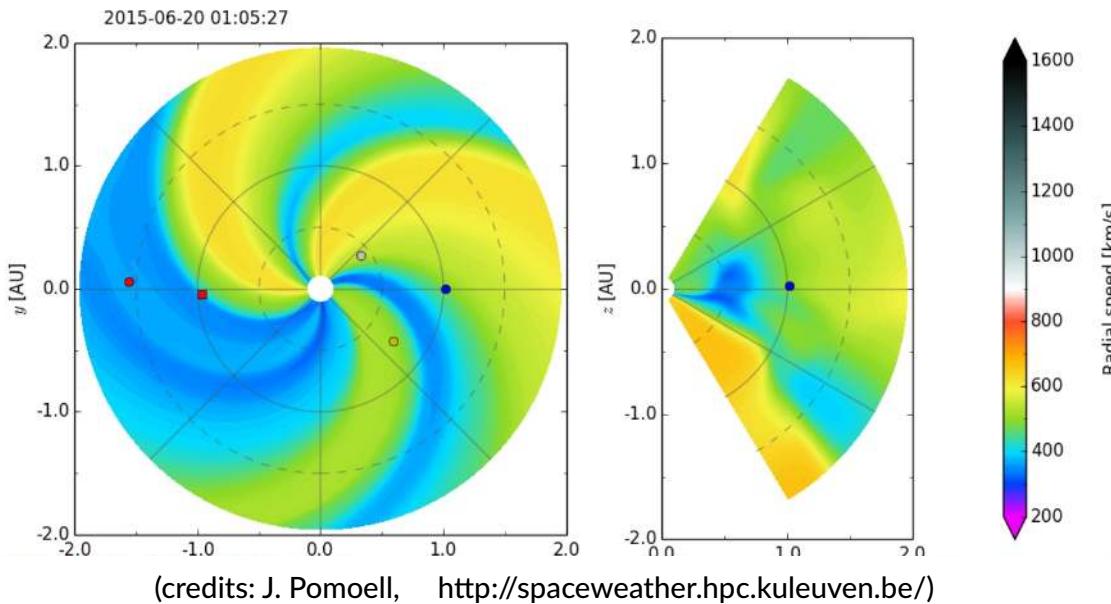
- expansion profiles
- fieldline inclination

$$\partial_s v = \frac{v}{1 - M^2} \left[\frac{GM_{\odot} m_p}{2r^2 k_b T} \cos(\alpha) - \frac{1}{A} \partial_s A \right]$$

α : inclination angle
 A : flux-tube cross-section

Solar wind throughout the heliosphere

VSWMC - EUFORIA



1) Semi-empirical model of the corona

- PFSS/SCS: geometry of B
- WSA: map of V_{wind} at $21.5 R_{\text{sun}}$ ($=0.1 \text{ AU}$)
- Empirical scalings $n-V_{\text{wind}}$ and $T-n$

$$\rho = \rho_{\text{fsw}} (v_{\text{fsw}} / v)^2$$
$$T = T_{\text{fsw}} (\rho_{\text{fsw}} / \rho)$$

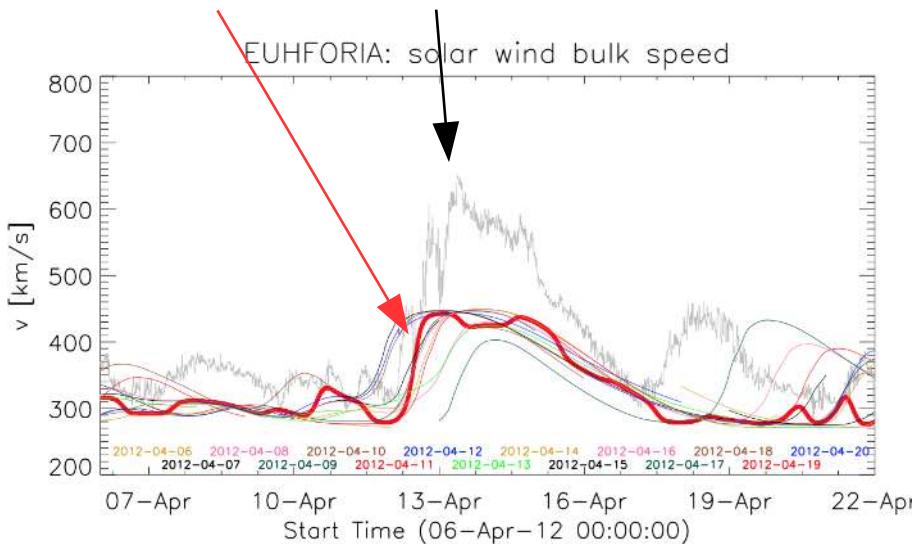
→ but there is now an alternative!

2) Physical (MHD) heliosphere

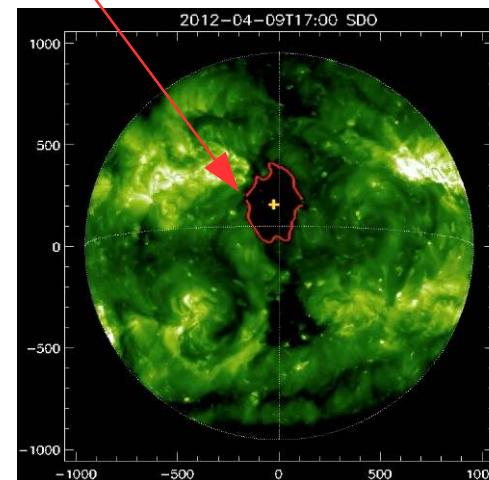
- EUFORIA

propagate super-critical wind into IP medium
rotation: formation of SIR

EUFORIA vs. ACE (forecasting of a HSS)



Boundary of most likely source CH



Heliospheric propagation of solar wind data

HELIOPROPA (helipropa.irap.omp.eu)

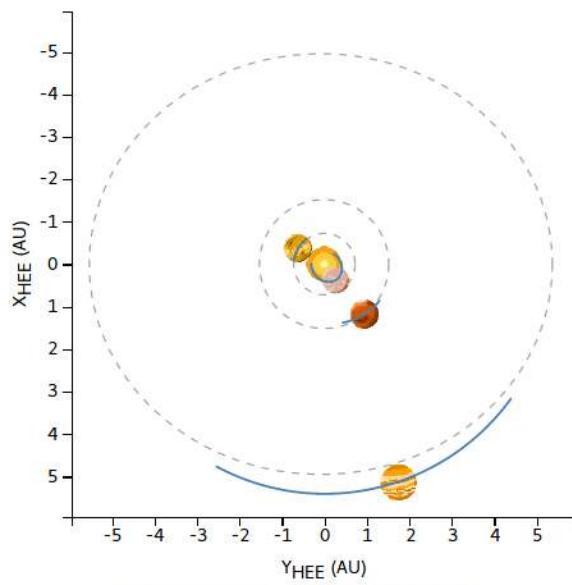


Forward + backward propagation of in-situ data
(to planets / other bodies at the ecliptic)

SW1D model

1D MHD heliospheric propagation model (Tao et al 2005)

Short-term forecast window (for outer planets)



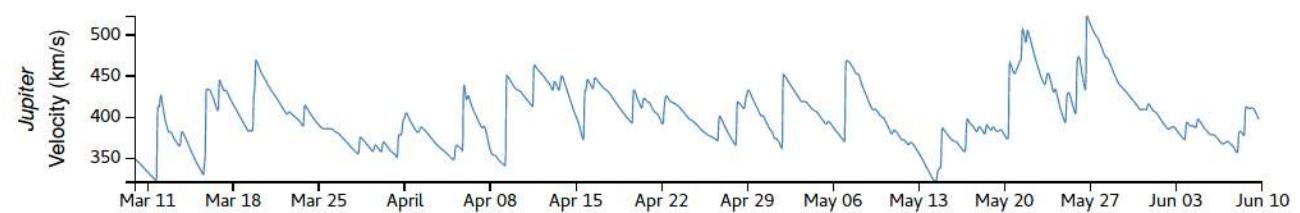
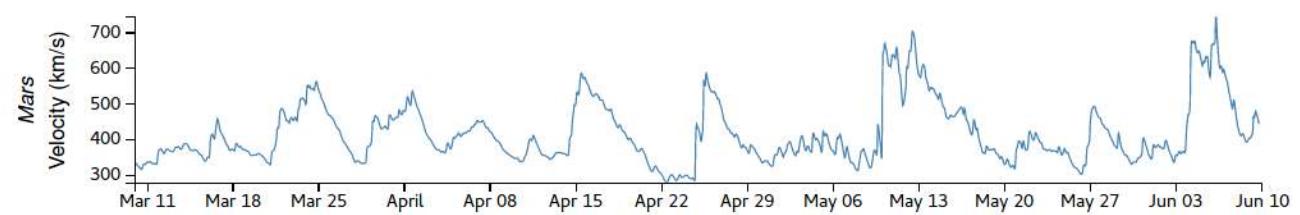
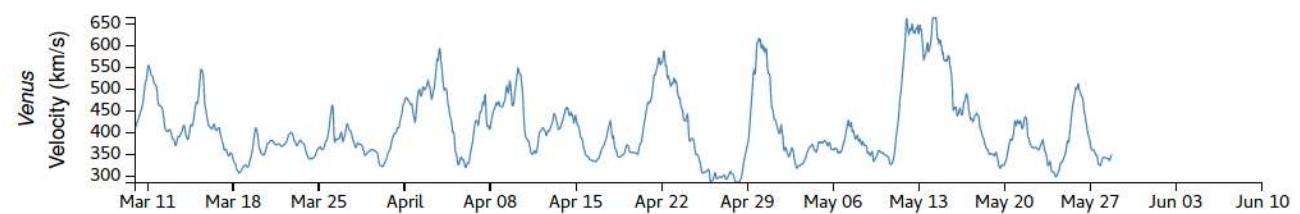
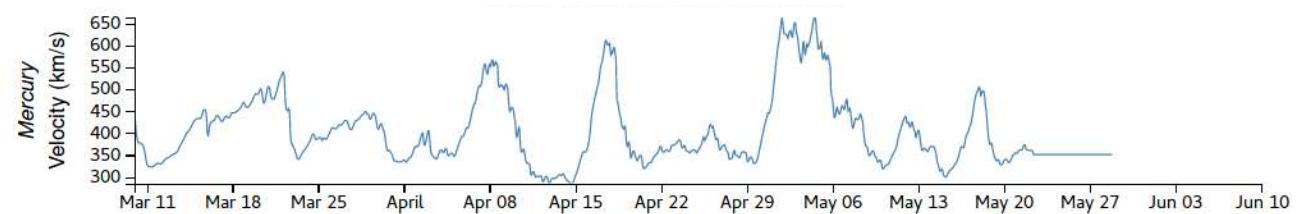
ZOOM IN

ZOOM OUT

DOWNLOAD

SAMP

ALERT



Solar wind modeling challenges

What we need now

- **Global physics-based modeling**
- **Other quantities**
(density, dyn pressure, phase speeds)
- **Diagnostics**
(synthetic imagery and in-situ)
- **Full surface to heliosphere**
(propagation to 1 AU, Earth, planets, multiple s/c positions)
- **Add minimal amount of complexity**
(robustness, quick computation)
- **Real-time modeling**
(current full 3D MHD models are very CPU intensive)

SWiFT / MULTI-VP Data-driven solar wind model

Sun / surface observations

(magnetograms: HMI, WSO, ADAPT)

Coronal field reconstruction

(PFSS, NLFFF, etc)

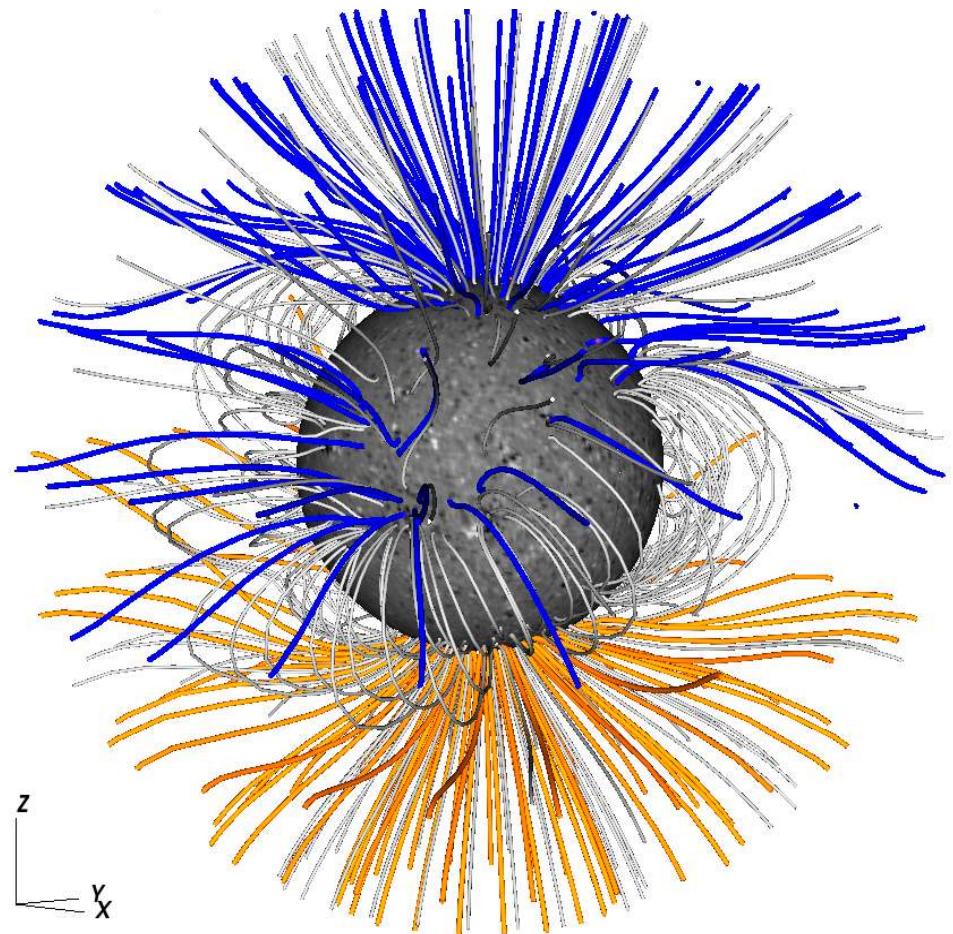
MULTI-VP

Heliospheric propagation

(ENLIL, EUHFORIA)

Earth / interplanetary medium

(S/C data, heliospheric imaging)



PFSS field lines: **positive** / **negative** polarity

SWiFT / MULTI-VP Data-driven solar wind model

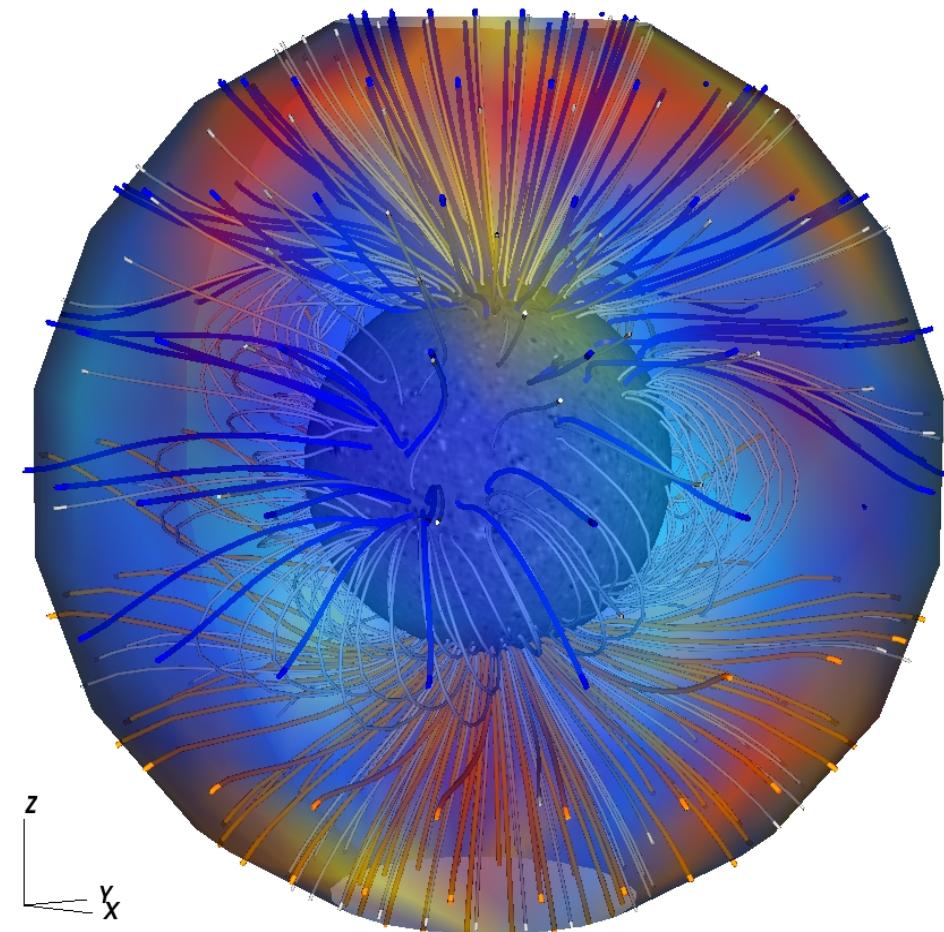
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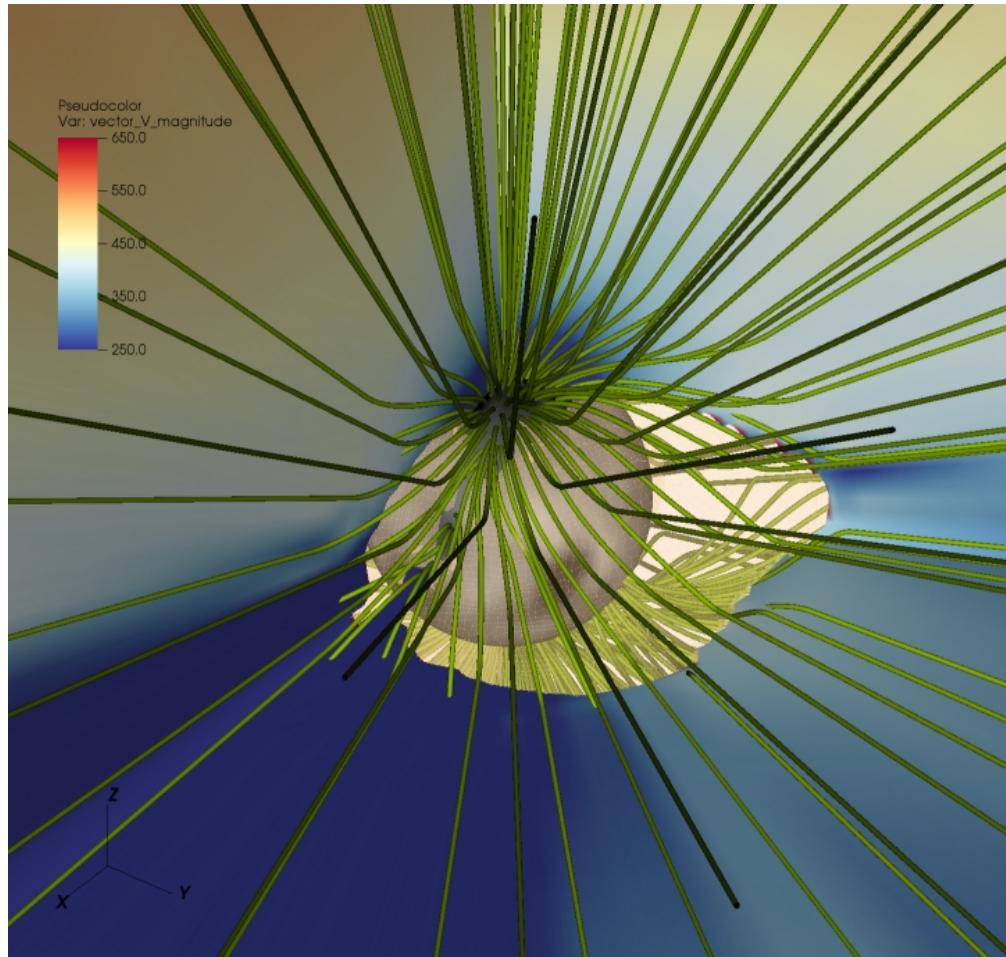
PFSS field lines: **positive** / **negative** polarity
Wind speed: **300** / **700** km/s

SWiFT framework pipeline

MULTI-VP Data-driven solar wind model

CR 2055

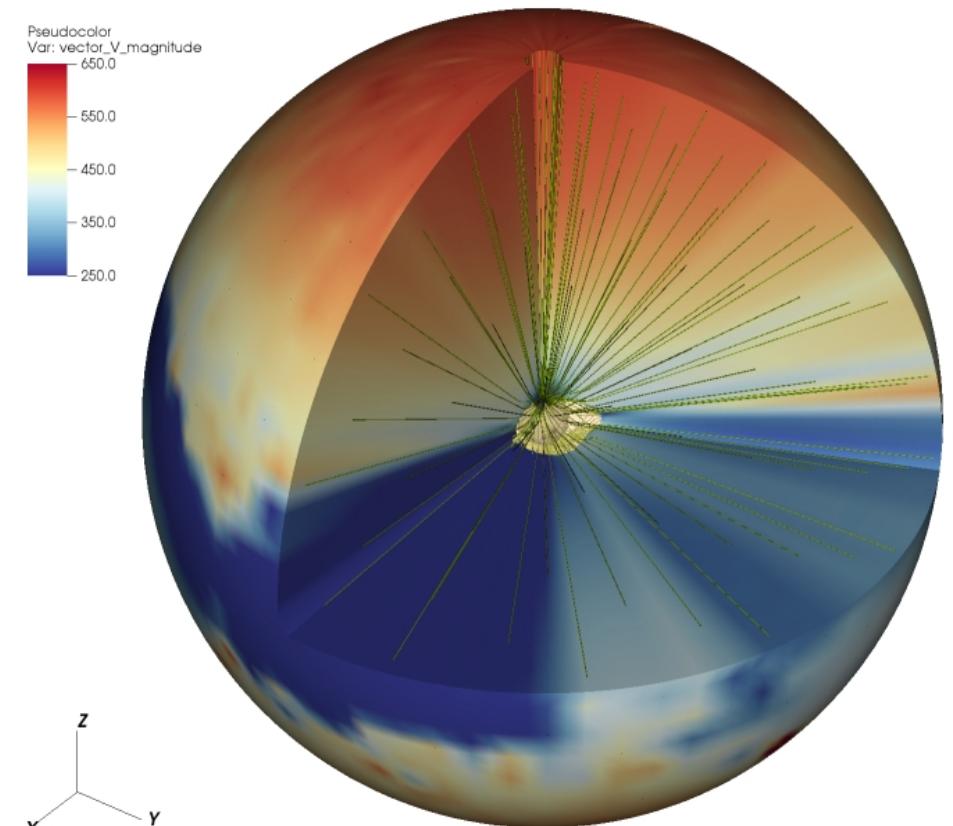
Solar wind speed



Low corona (close-up view)

Open magnetic fieldlines ("coronal holes")
Streamer / coronal hole boundaries

Fast wind
Slow wind



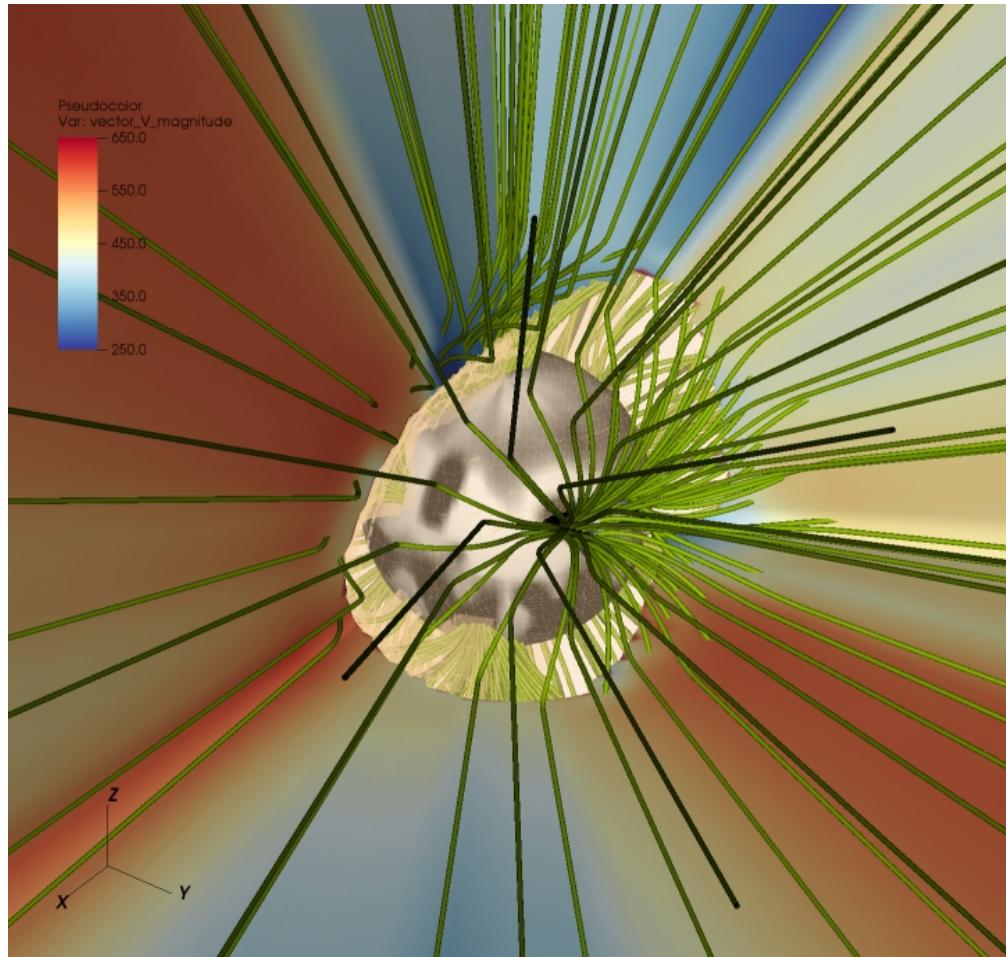
High corona ($1 - 15 R_{\text{sun}}$)

Pinto, Rouillard, ApJ (2017)

MULTI-VP Data-driven solar wind model

CR 2132

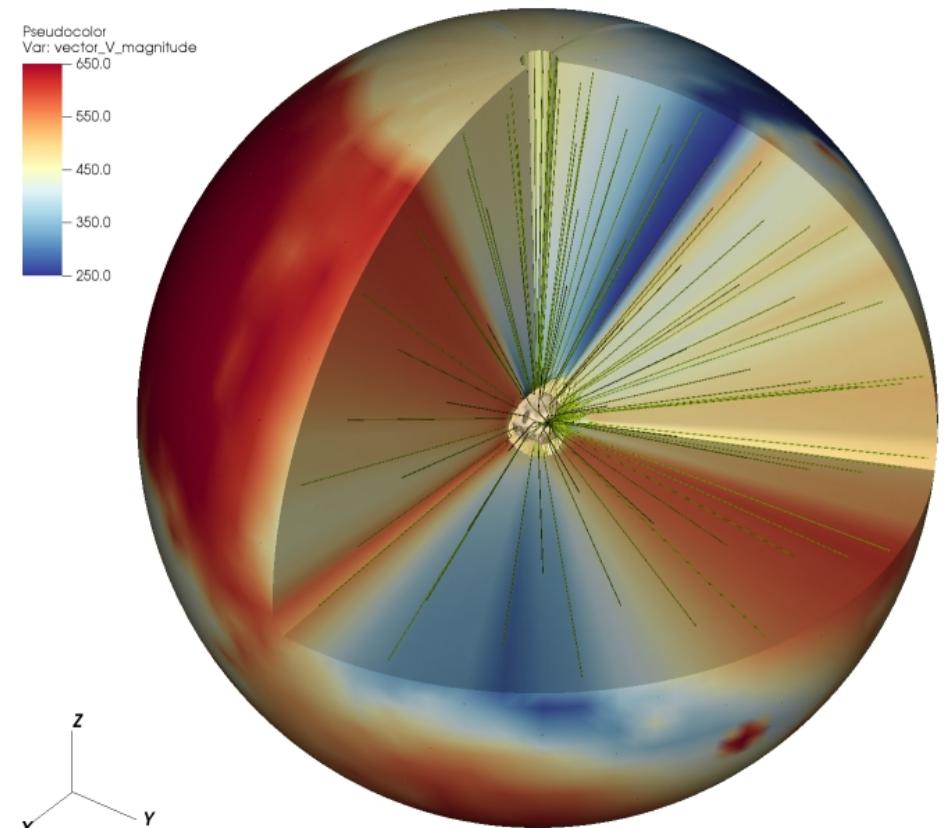
Solar wind speed



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Slow wind

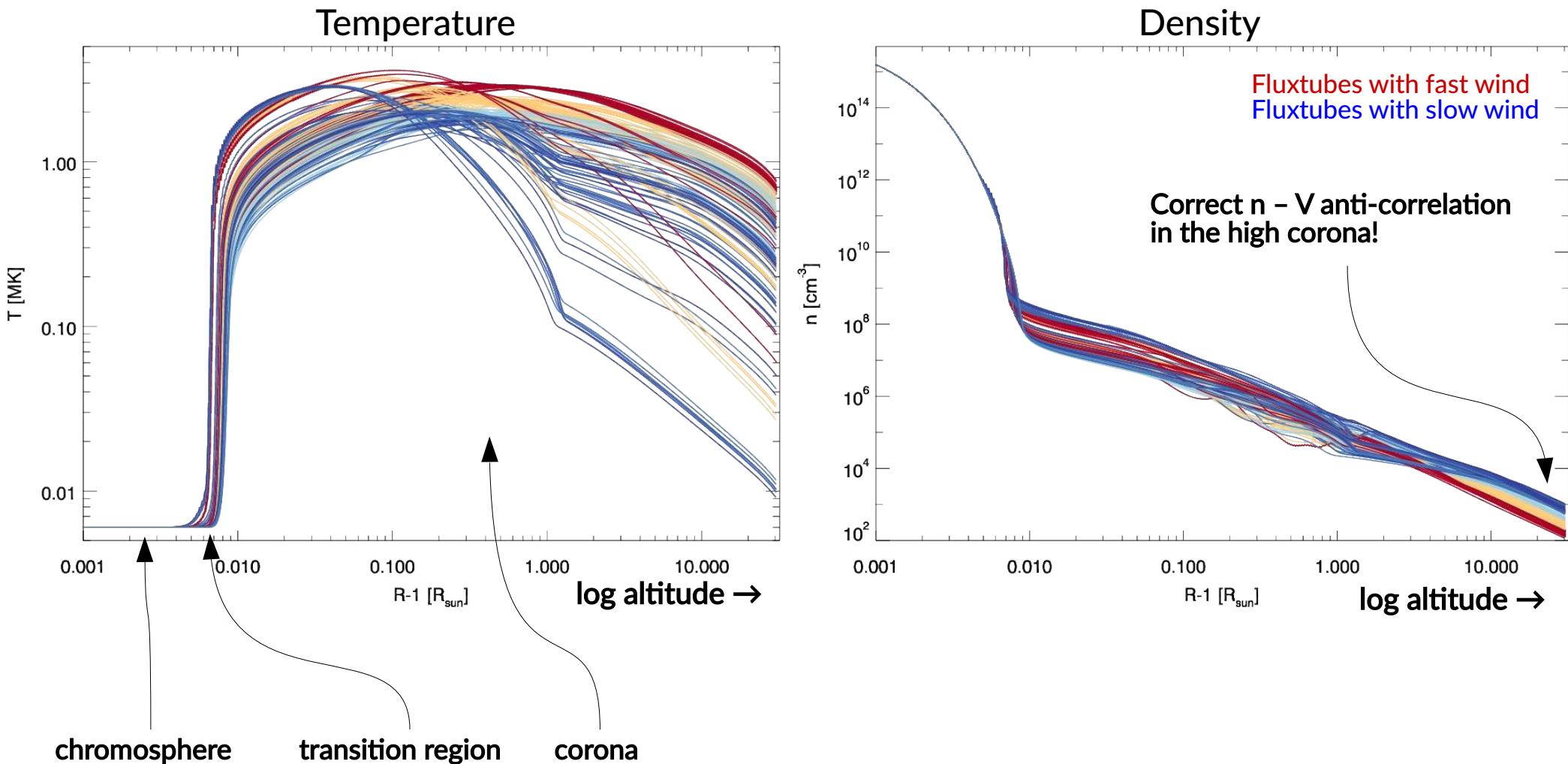


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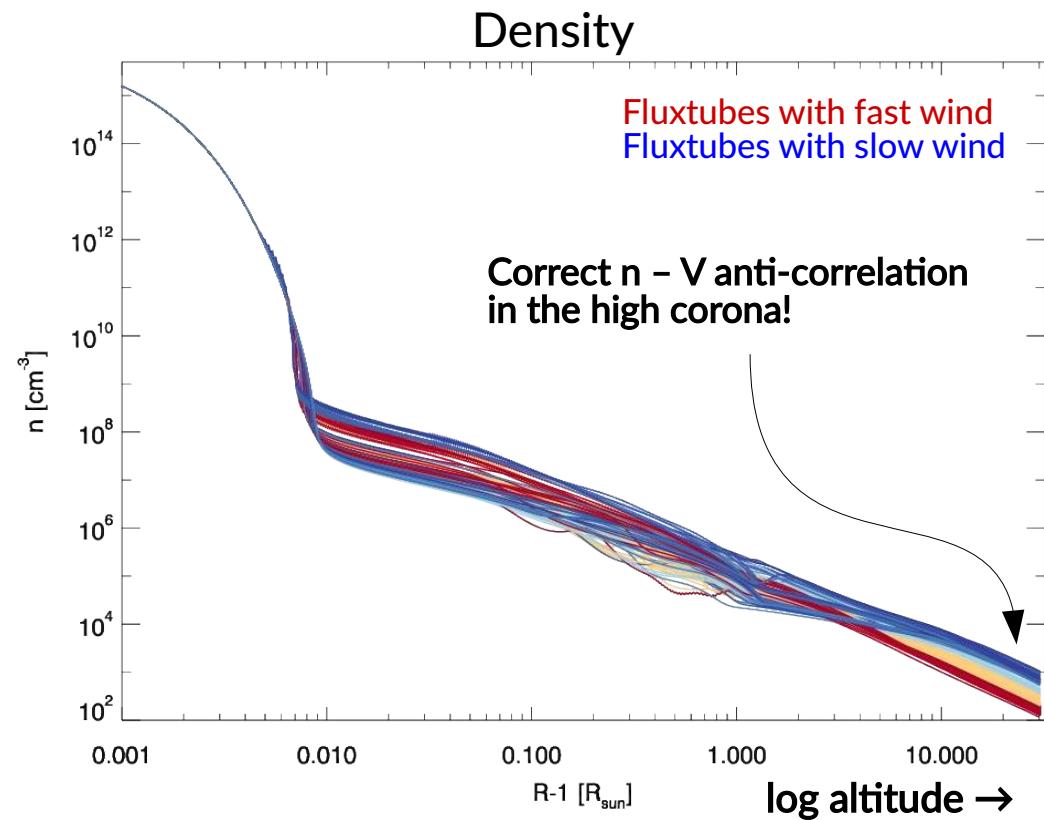
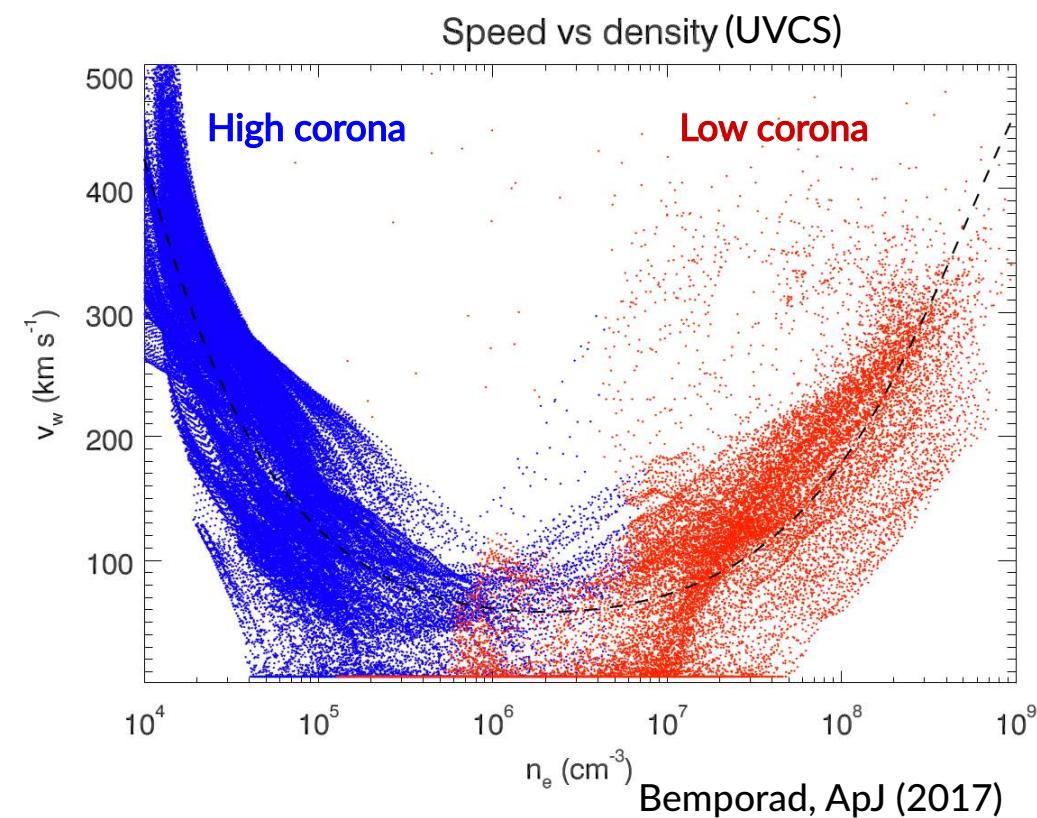
MULTI-VP Data-driven solar wind model

From the surface to the corona



MULTI-VP Data-driven solar wind model

From the surface to the corona



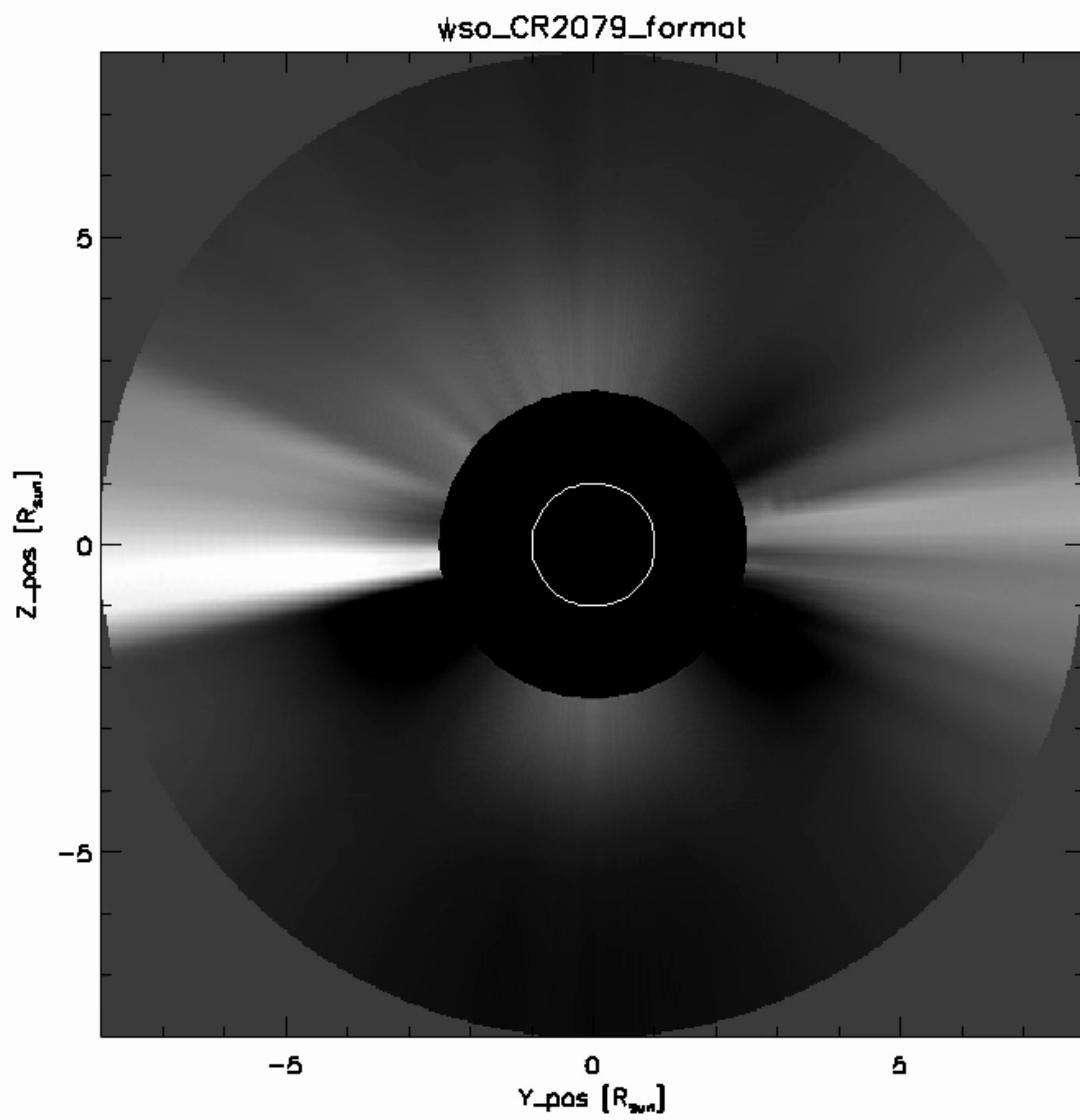
MULTI-VP Data-driven solar wind model

Synthetic images of the corona

CR 2079 – 2080

MULTI-VP + FORWARD

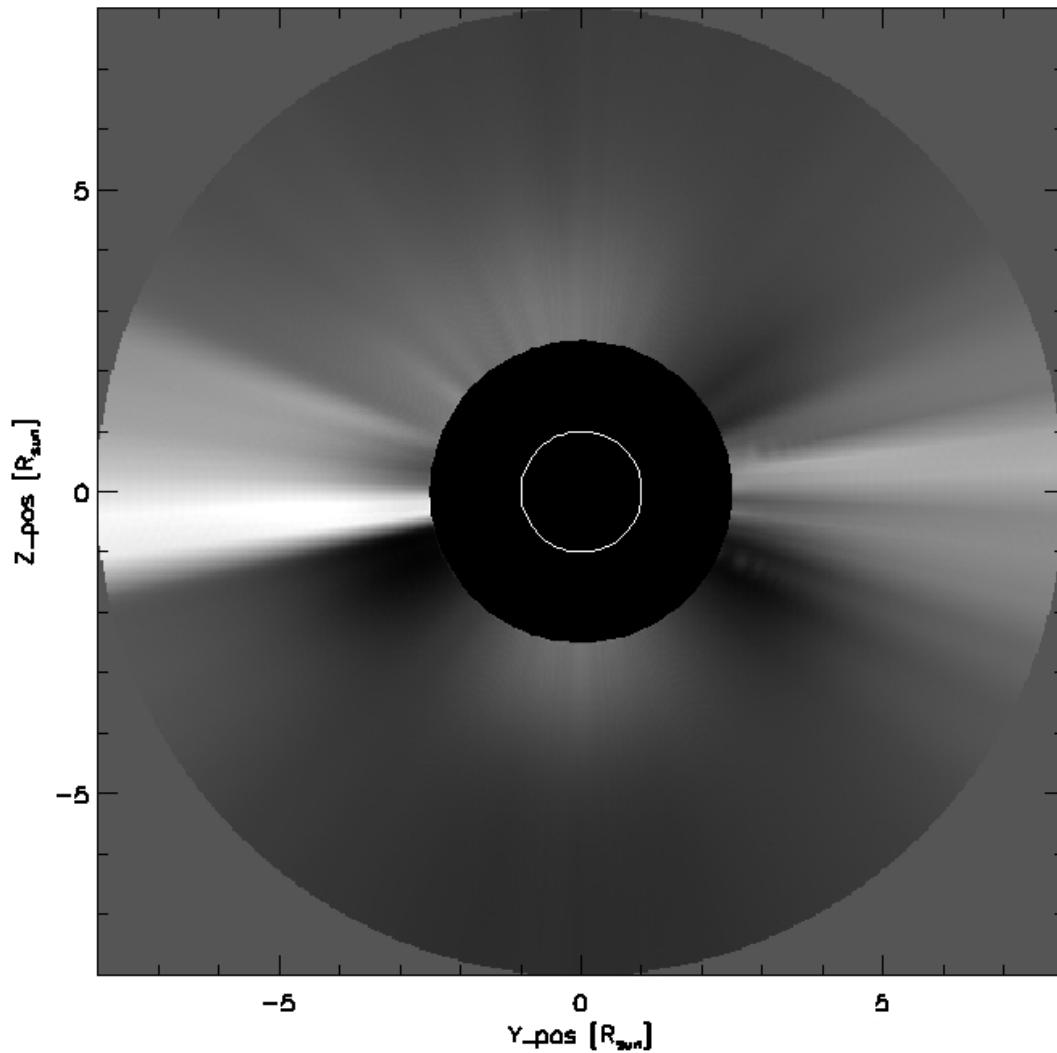
(NRGF-filtered, ~C2 FoV)



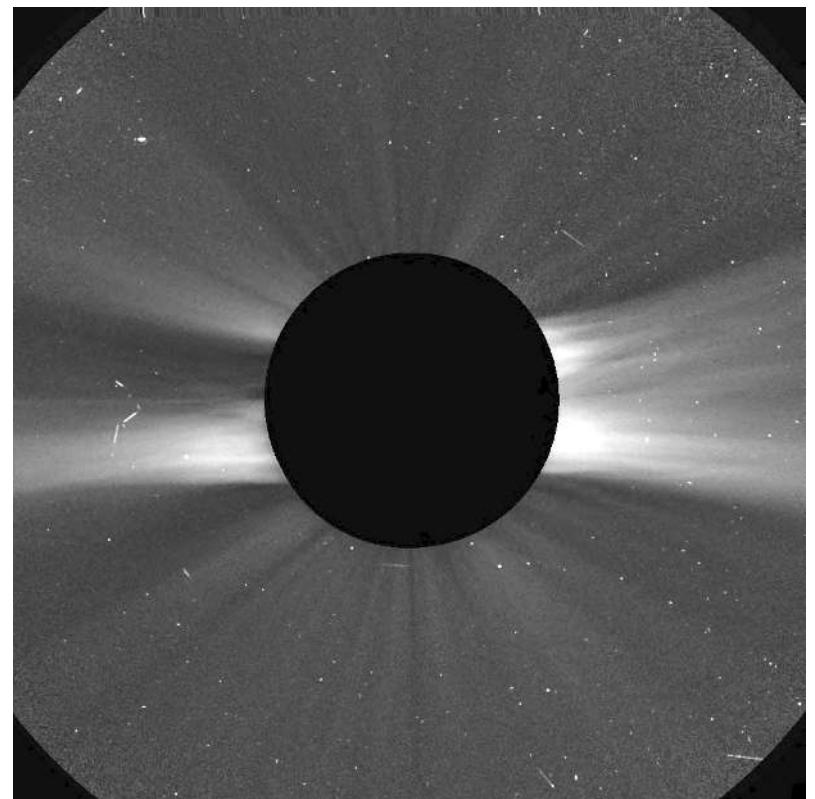
MULTI-VP Data-driven solar wind model

Synthetic images of the corona

CR 2079 (Earth, mid-CR, MULTI-VP NRGF-filtered)



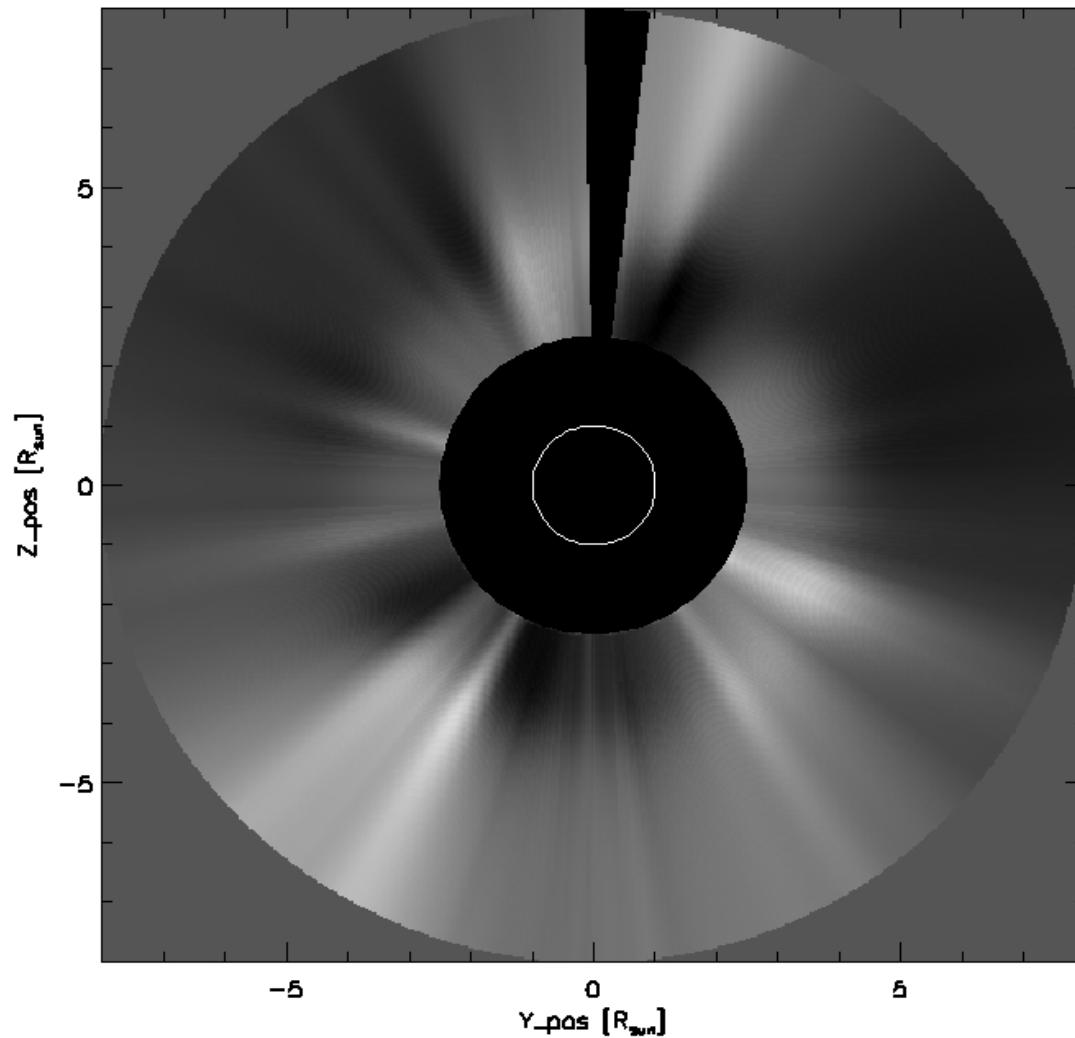
CR 2079 (L1, mid-CR, LASCO C2)



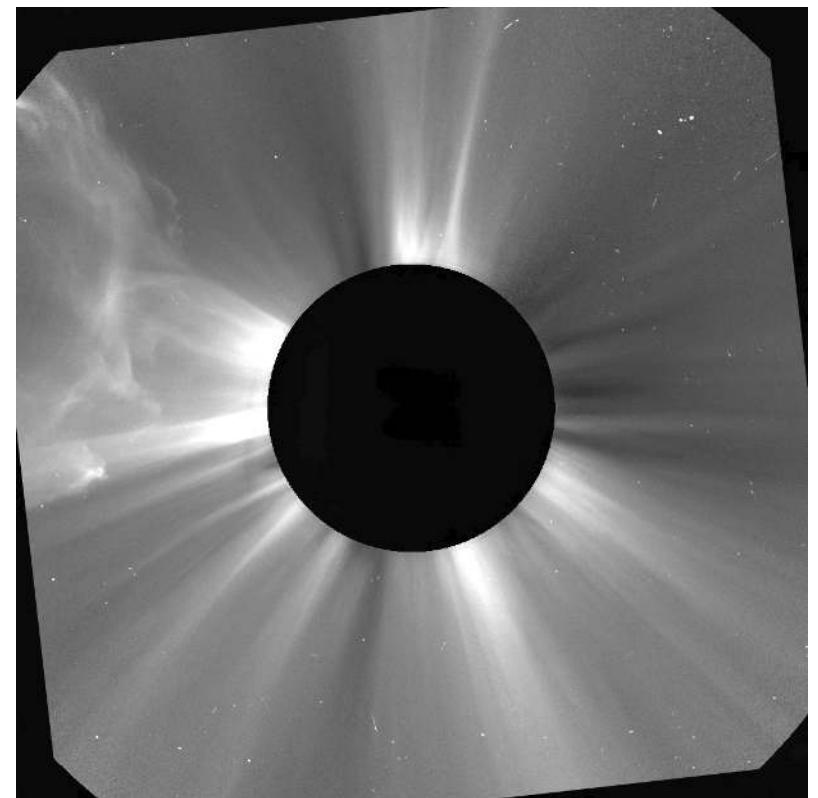
MULTI-VP Data-driven solar wind model

Synthetic images of the corona

CR 2136 (Earth, mid-CR, MULTI-VP NRGF-filtered)



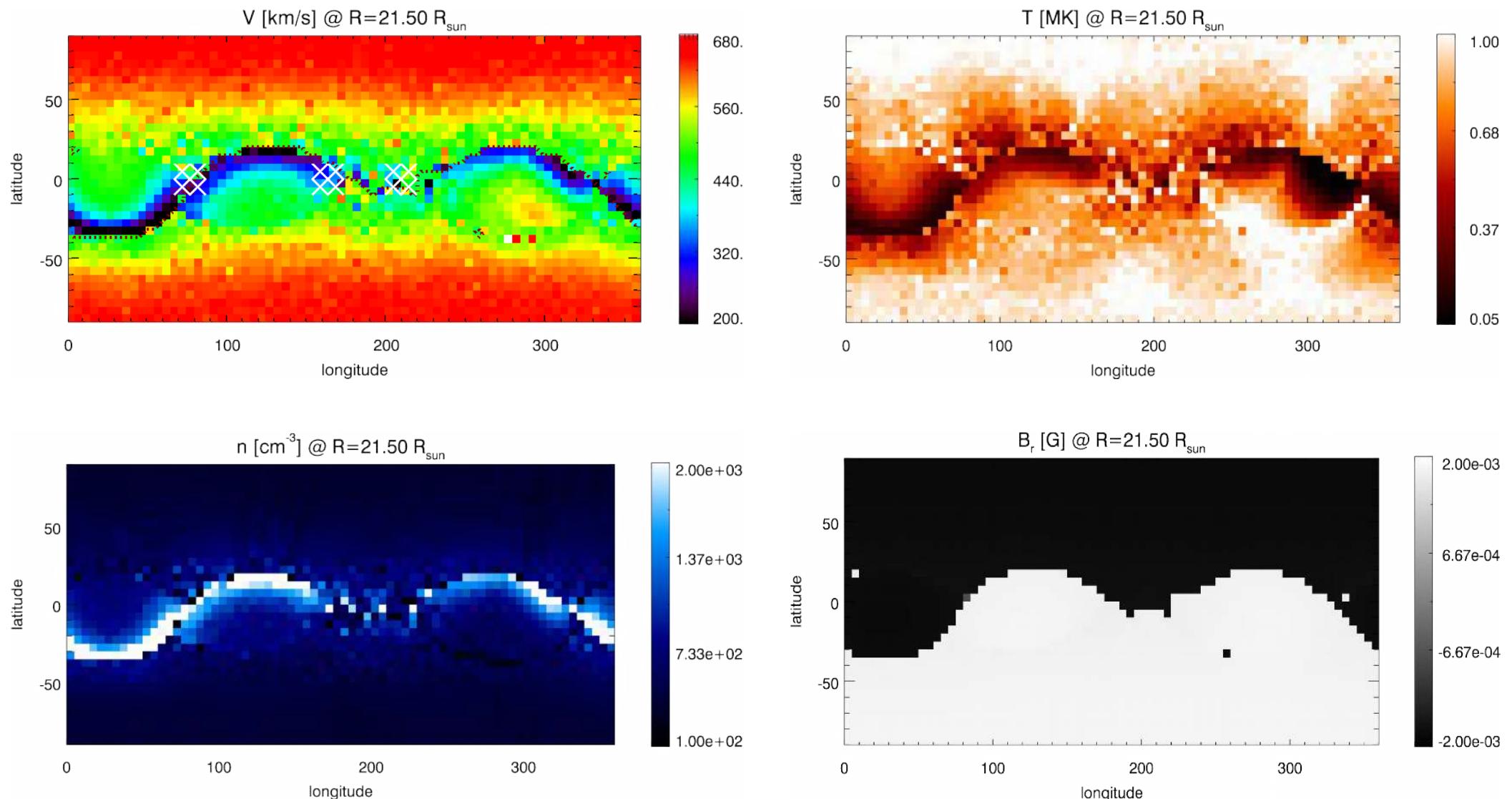
CR 2136 (L1, mid-CR, LASCO C2)



MULTI-VP Data-driven solar wind model

Solar wind maps

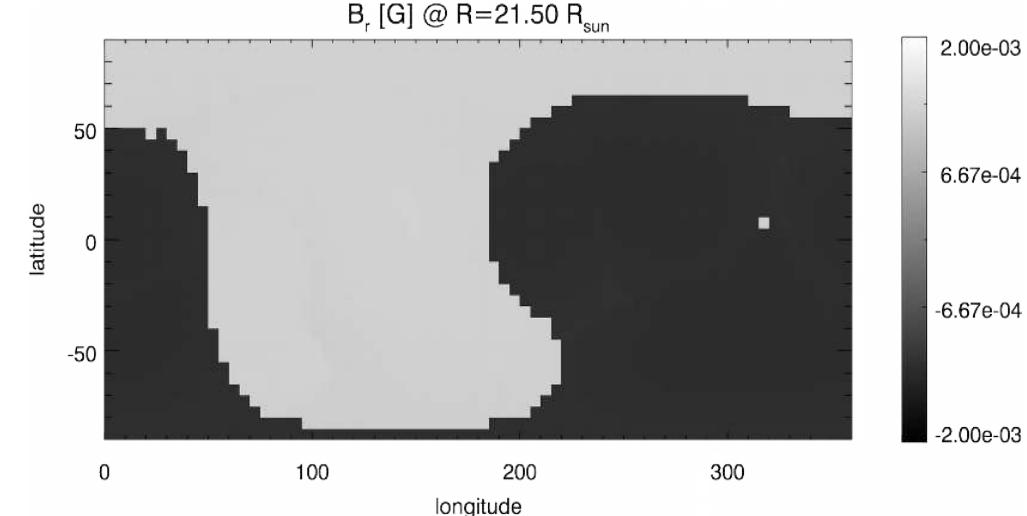
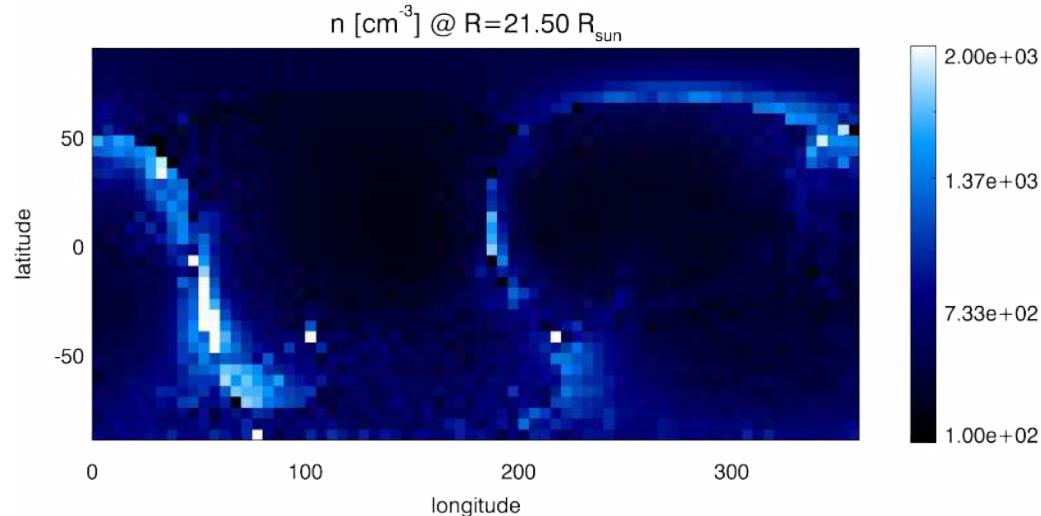
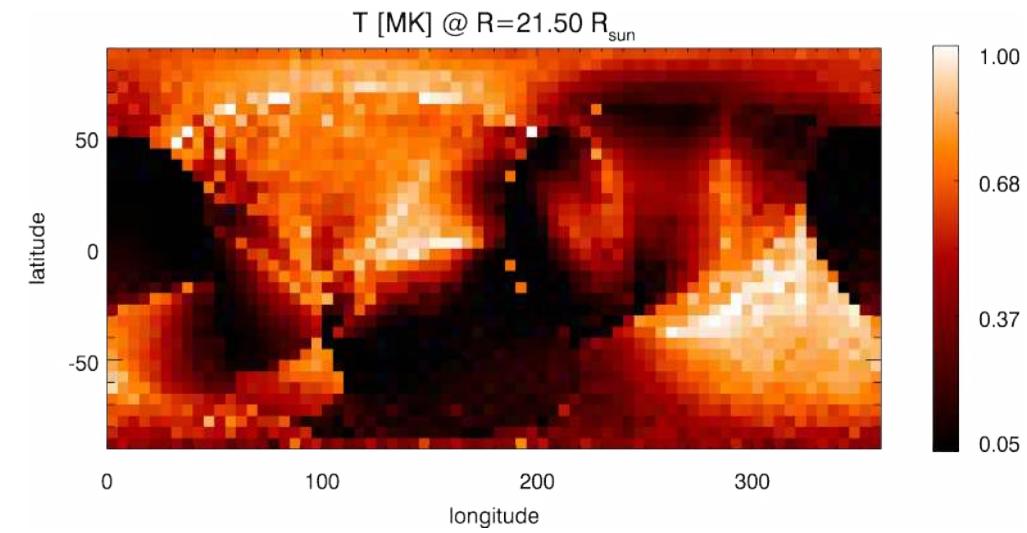
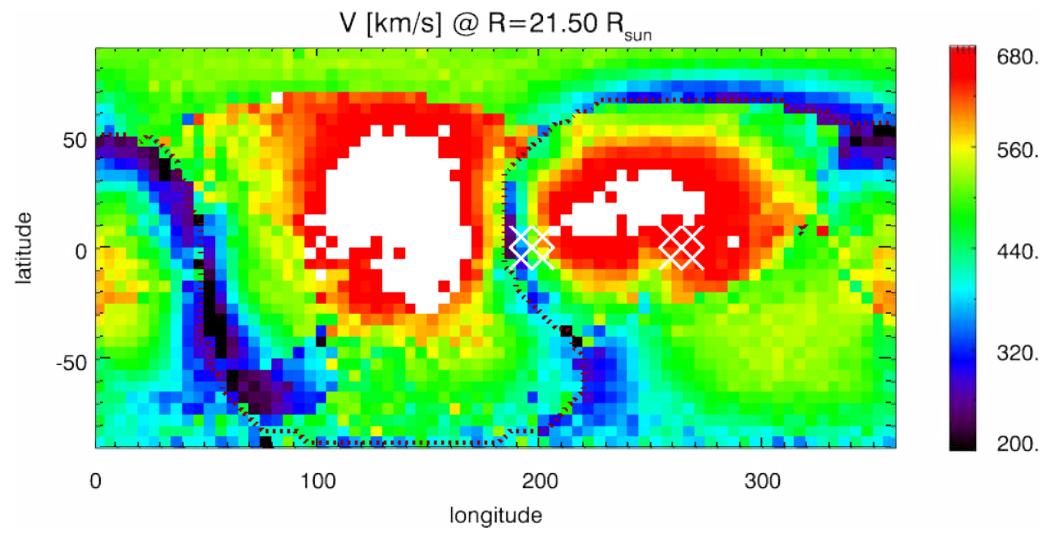
CR 2056 (2008, minimum)



MULTI-VP Data-driven solar wind model

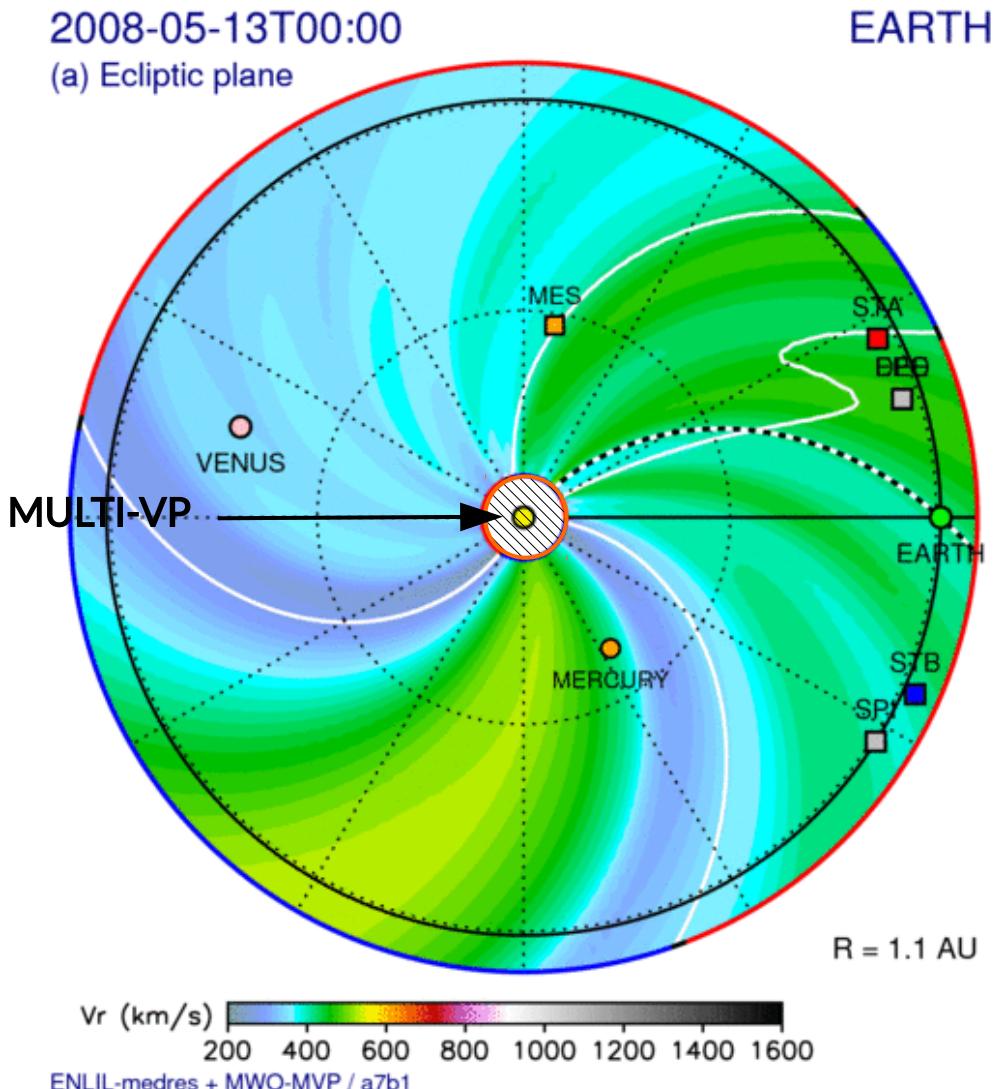
Solar wind maps

CR 2136 (2013, maximum)



Predicting the solar wind conditions at 1 AU

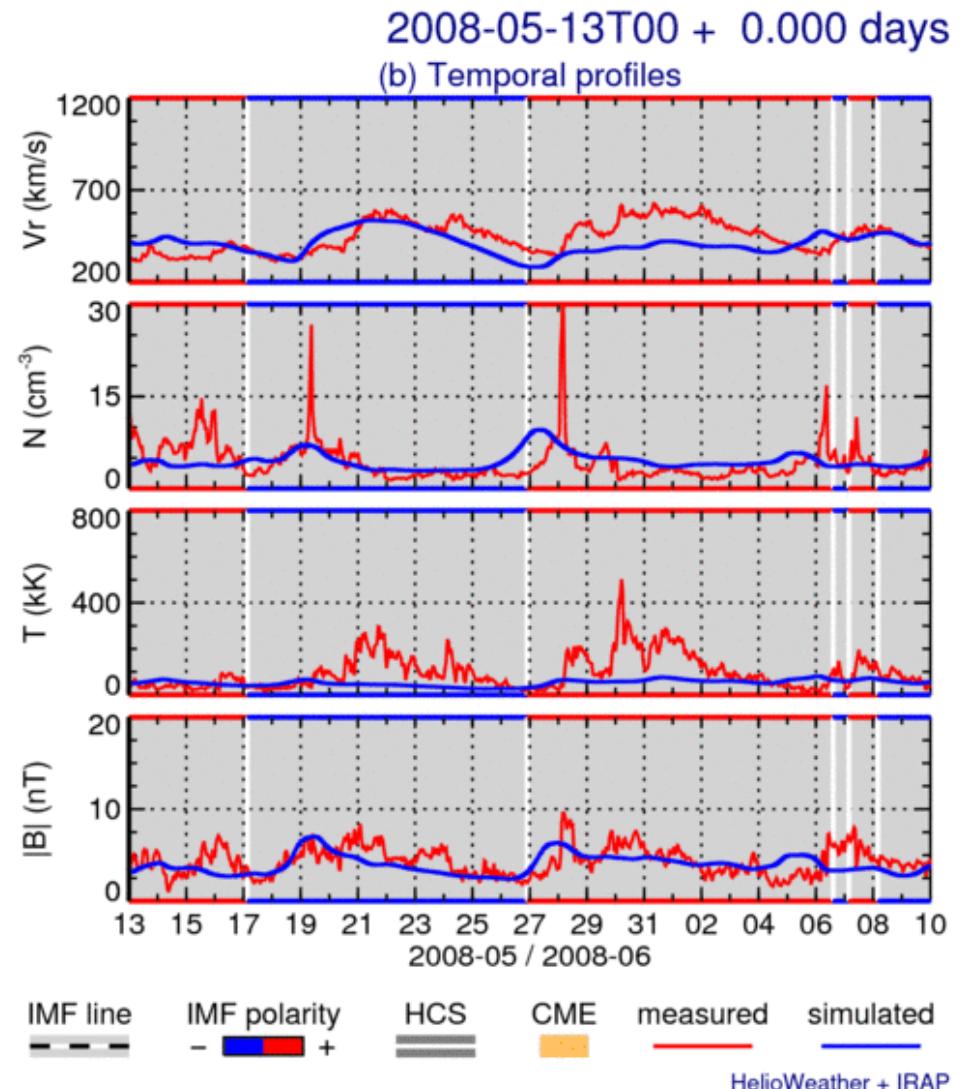
Interplanetary medium, in-situ data



MULTI-VP maps
at 21.5 Rsun



ENLIL



(Pinto, Rouillard, Odsctrill, Mays, in prep)



<http://www.helcats-fp7.eu/>
<http://stormsweb.irap.omp.eu>

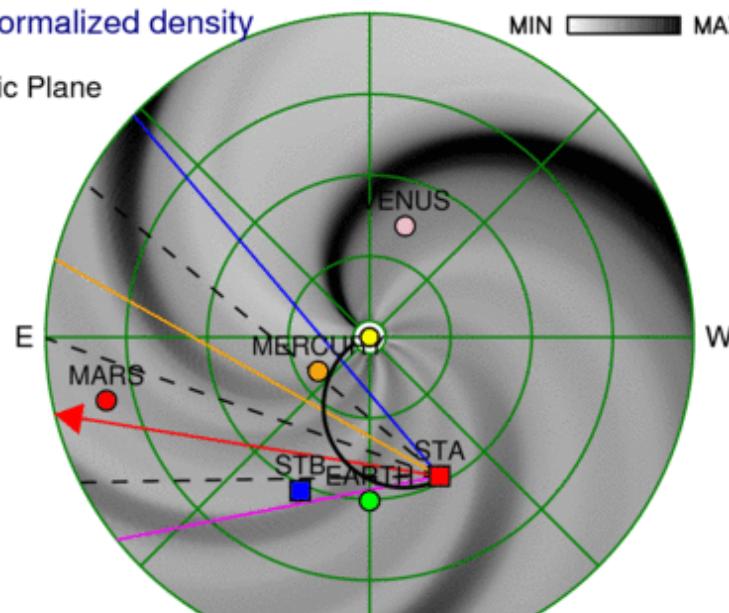
HELCATS catalogue: <https://stormsweb.irap.omp.eu/doku.php?id=windmaptable>

Predicting the solar wind conditions: J-maps

2008-05-13T00:00

(a) Normalized density

Ecliptic Plane



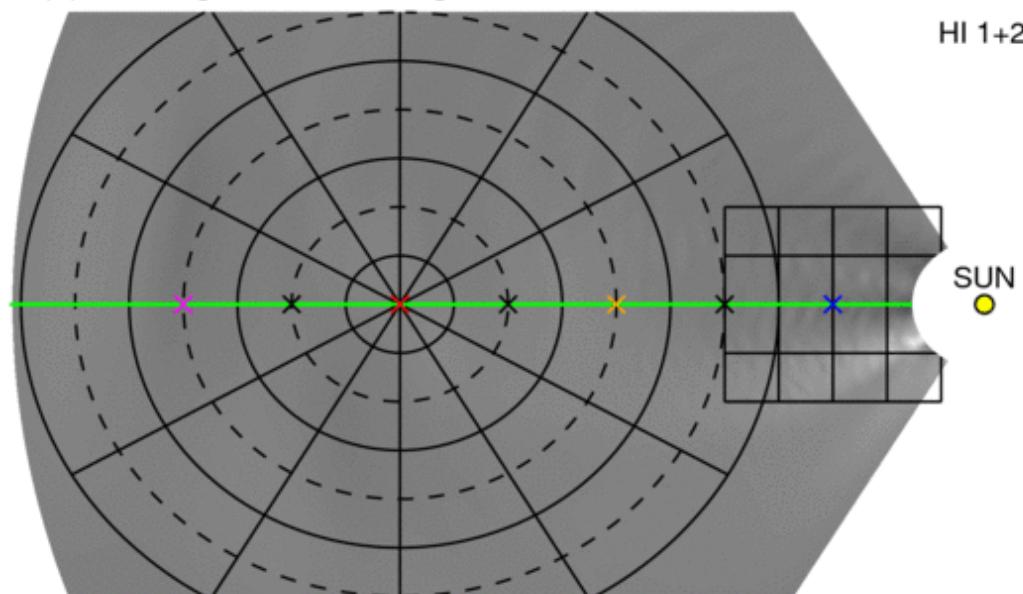
STEREO-A

(b) Running-difference image

MIN ■ MAX

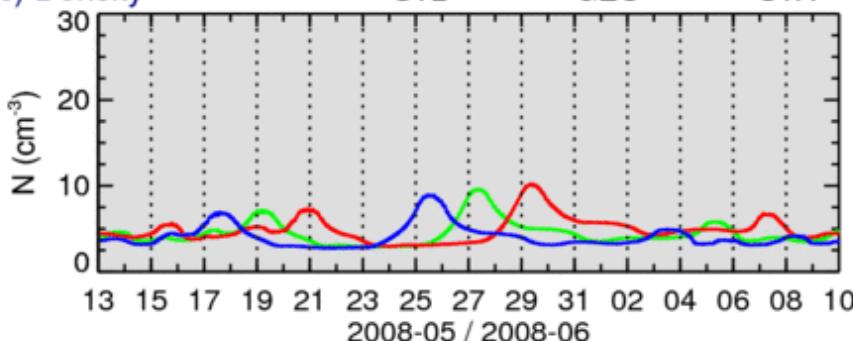
HI 1+2

2008-05-13T00 + 0.000 days



(c) Density

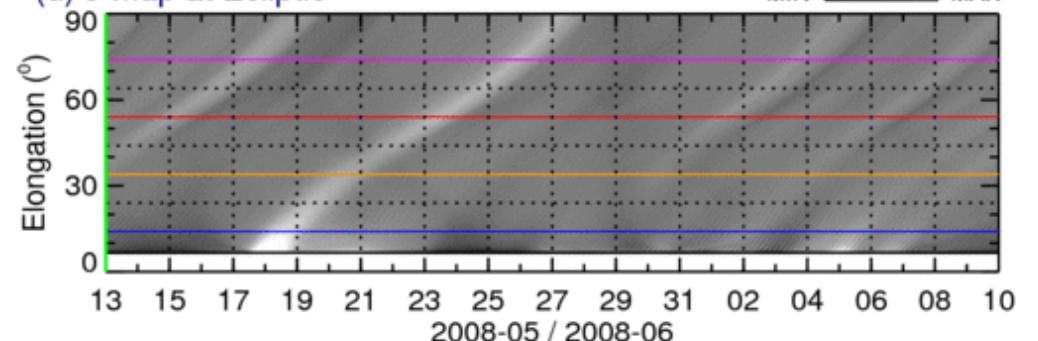
— STB — GEO — STA



ENLIL-medres + MWO-MVP / a7b1

(d) J-map at Ecliptic

MIN ■ MAX



2008-05 / 2008-06

HelioWeather + IRAP

(Pinto, Rouillard, Odsstrill, Mays, et al)



<http://www.helcats-fp7.eu/>
<http://stormsweb.irap.omp.eu>

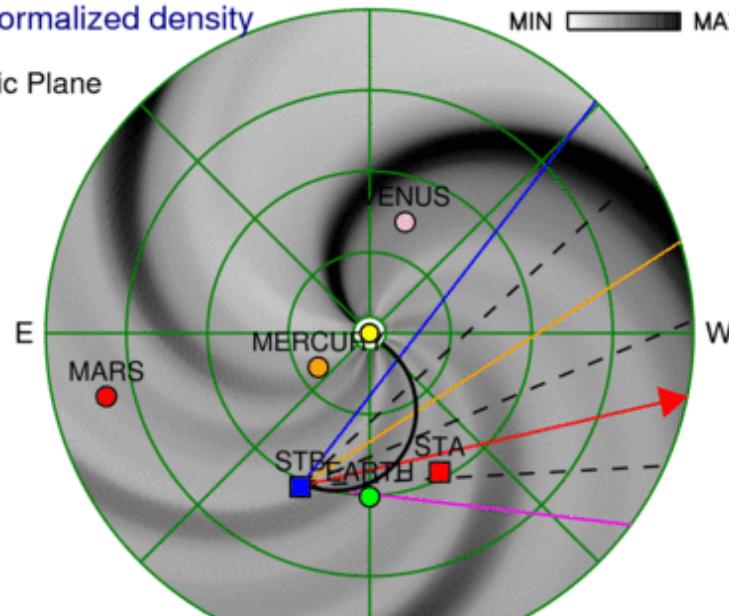
HELCATS catalogue: <https://stormsweb.irap.omp.eu/doku.php?id=windmaptable>

Predicting the solar wind conditions: J-maps

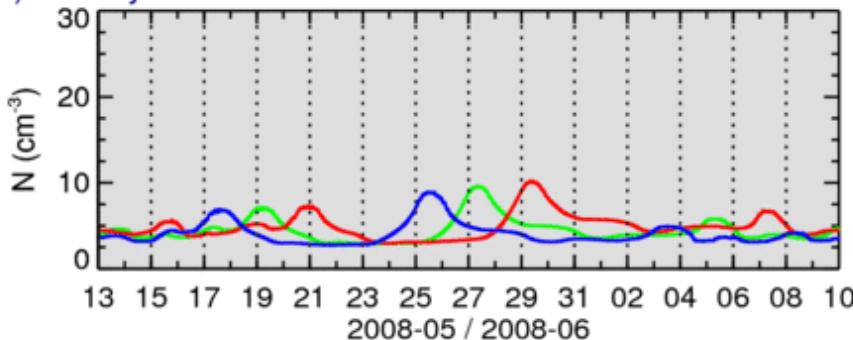
2008-05-13T00:00

(a) Normalized density

Ecliptic Plane



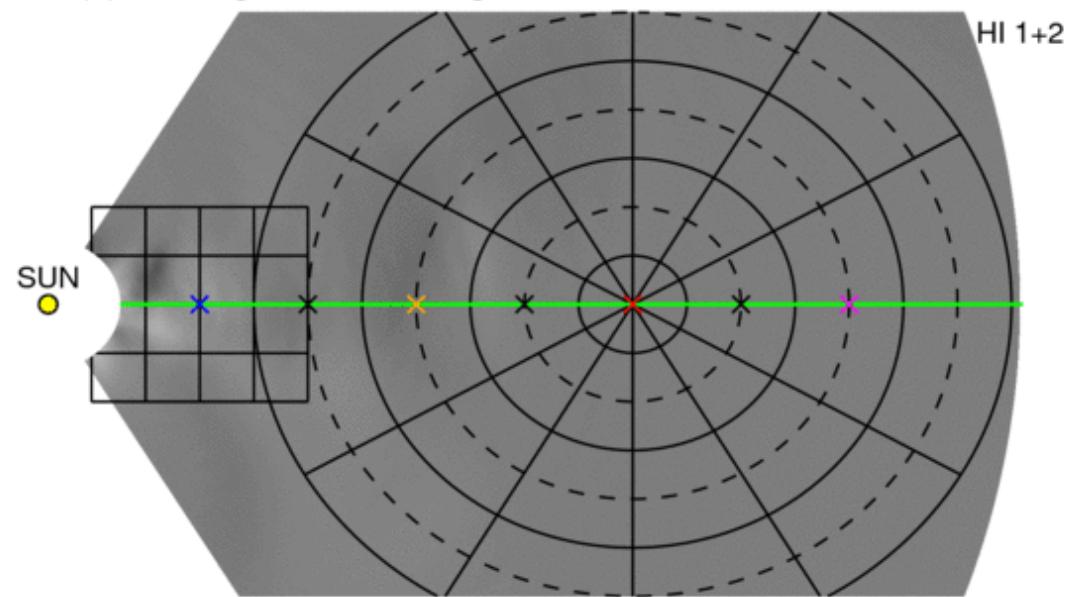
(c) Density



ENLIL-medres + MWO-MVP / a7b1

STEREO-B

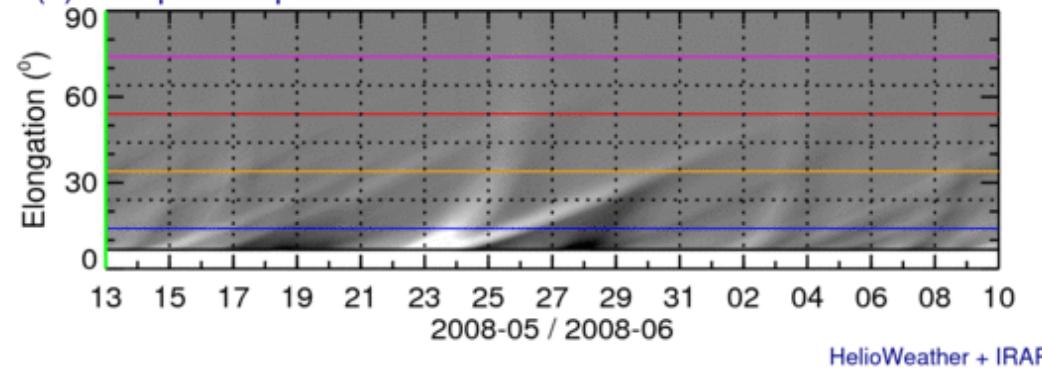
(b) Running-difference image



2008-05-13T00 + 0.000 days

MIN — MAX

(d) J-map at Ecliptic



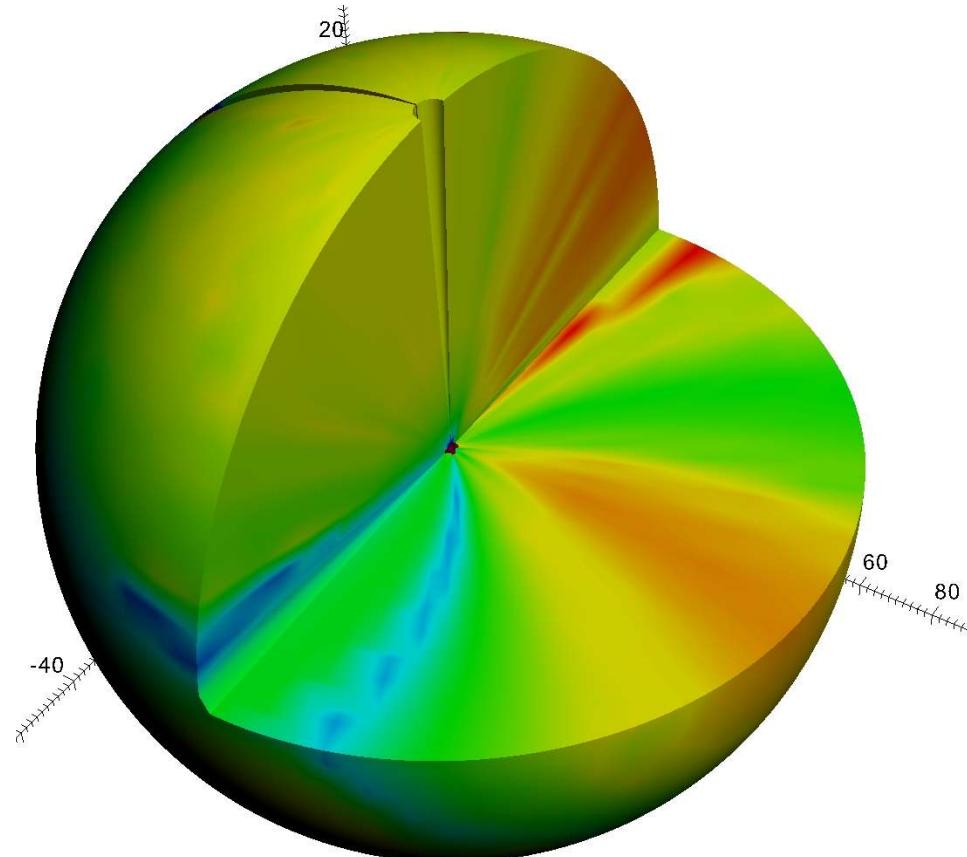
(Pinto, Rouillard, Odsstrill, Mays, et al)



WIP: SWiFT with 1D heliospheric propagation

1D Heliospheric SW propagation

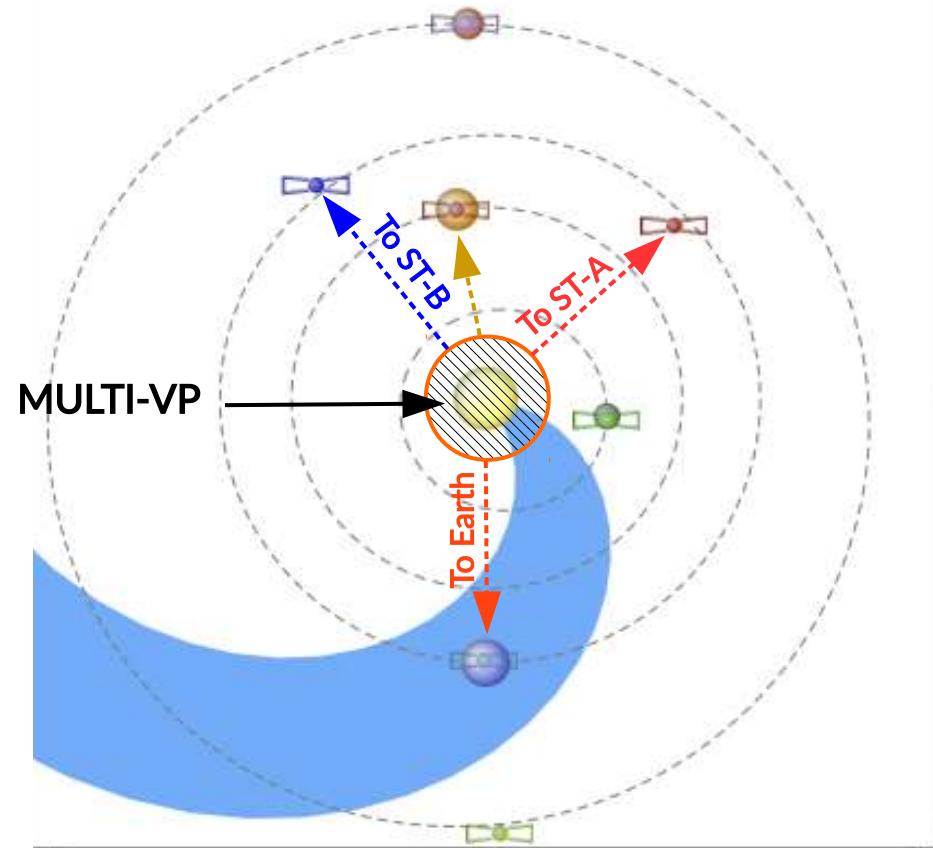
Multiple wind propagation paths, global view **1 to 90 R_{sun}**



Parker Solar Probe

perihelions: ~36 to 9 R_{sun}
radial scan regions (quasi-corotation) covered

Sun to spacecraft propagation paths (up to 1 AU ++)



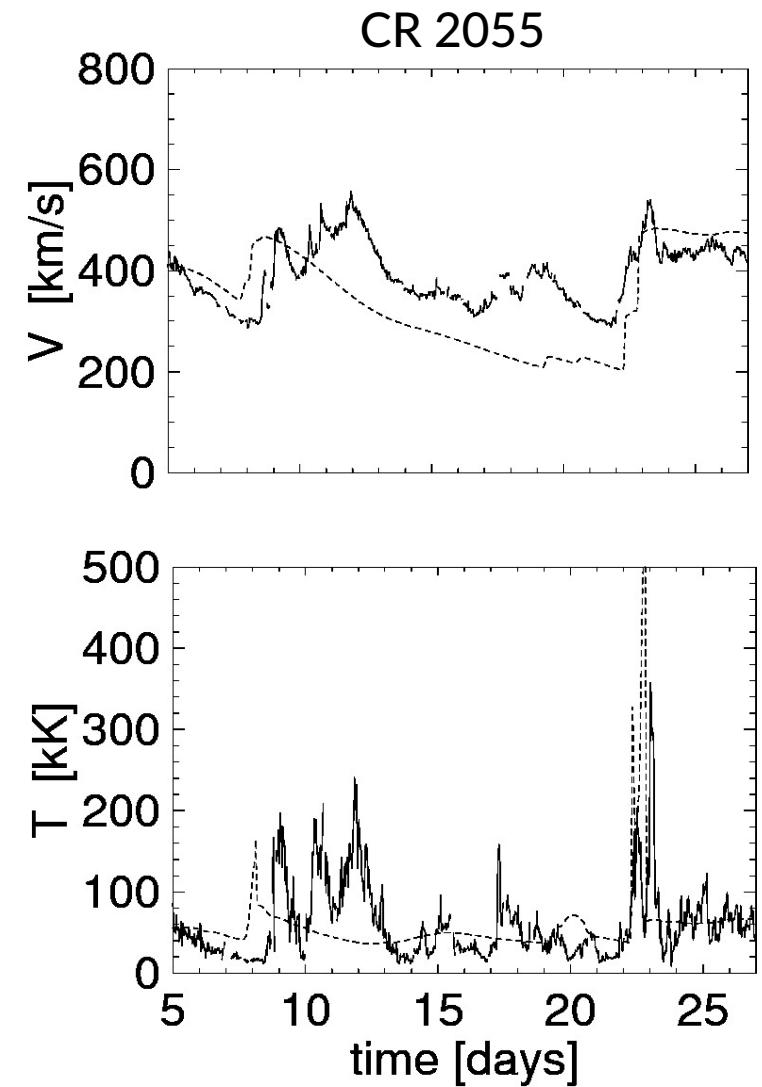
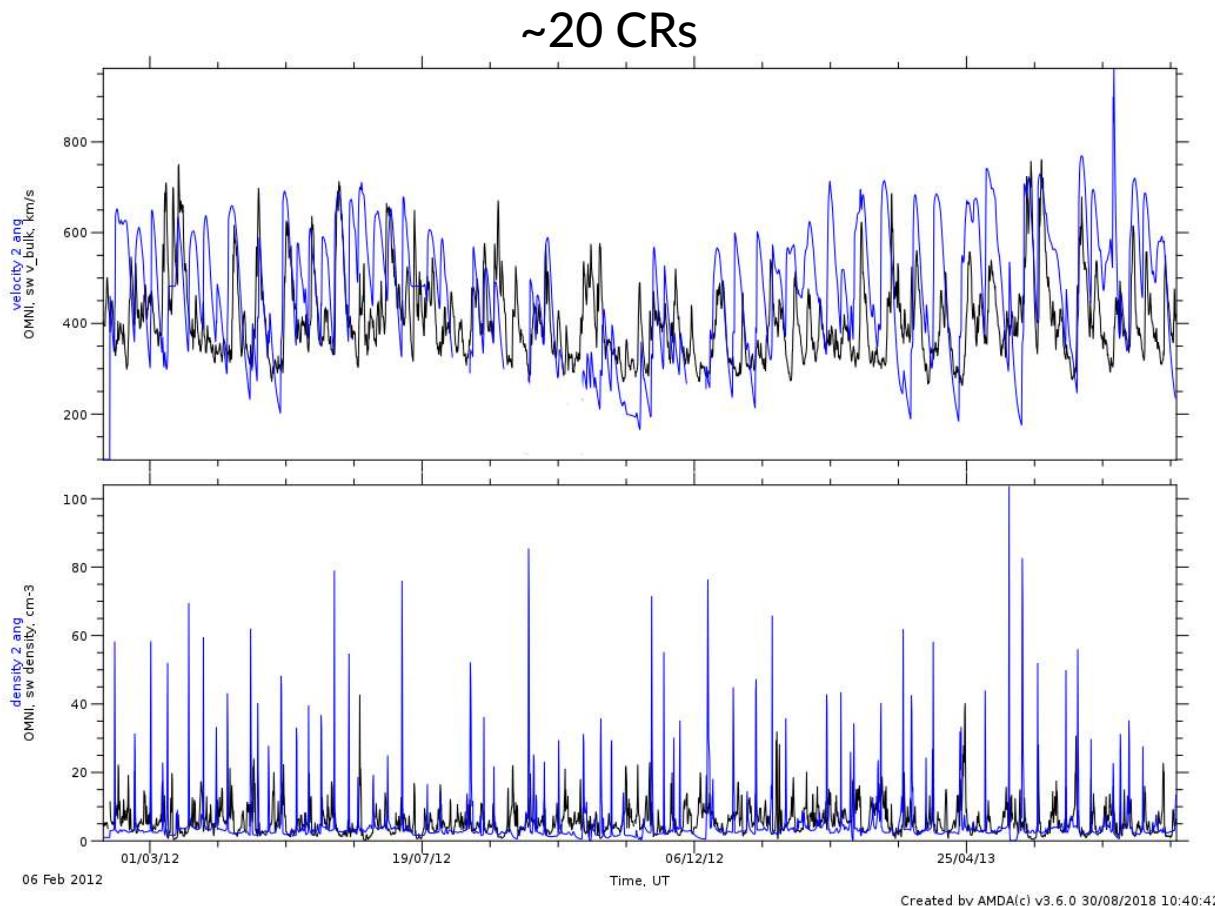
propagationtool.cdpp.eu
(propagation method from Tao et al, 2005)

Forward propagation from output of **MULTI-VP**
(arbitrary radial paths, any latitude or longitude)

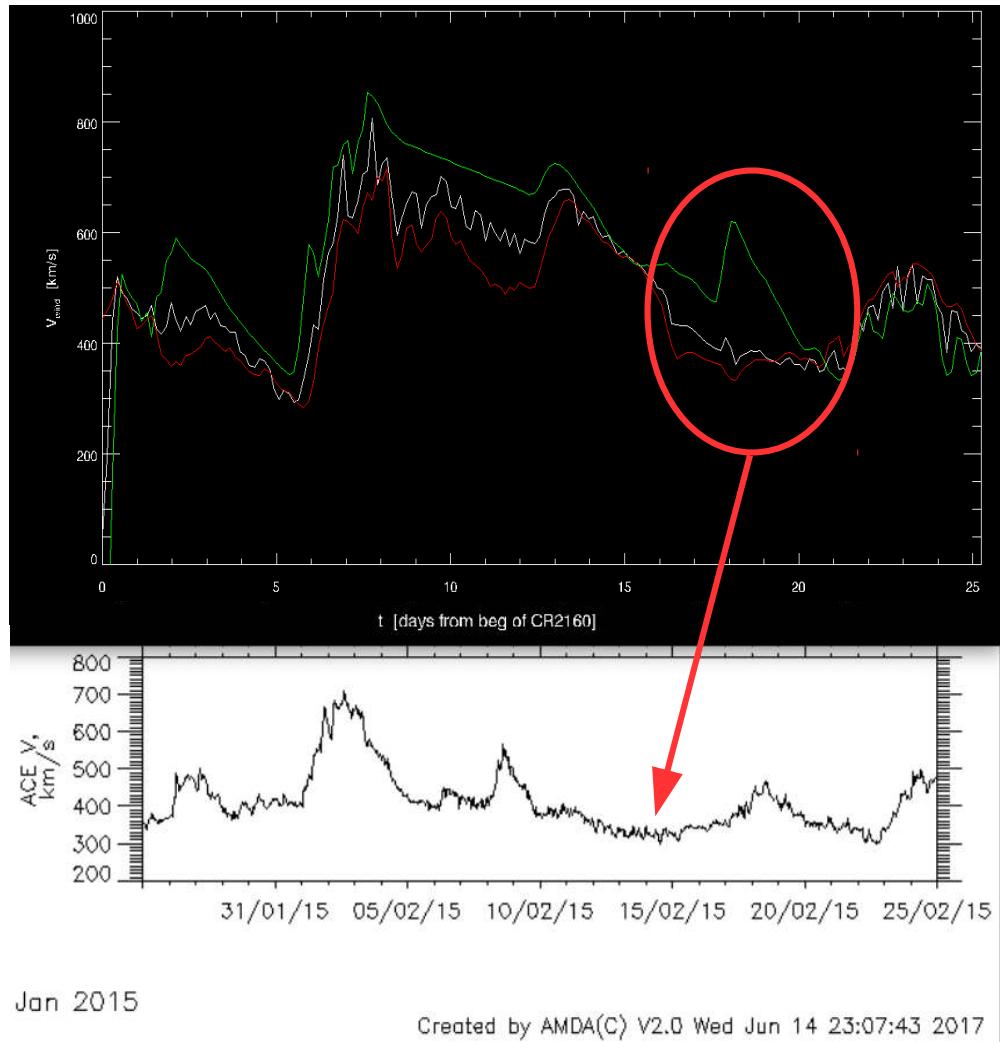
Multiple time-series, at **different orbital positions**,
rotation + background field evolution

WIP: SWiFT with 1D propagation

Simulations vs. OMNI



SWiFT with 1D propagation: some issues



- **Hits and misses i.r.t. in-situ data**
due, e.g., to errors in latitude of wind streams
- **Dependence on mag. field reconstruction**
background field evolution, B-map cross-calibration
- **No impulsive perturbations**
but rotation, intermittent heating, small δv waves OK

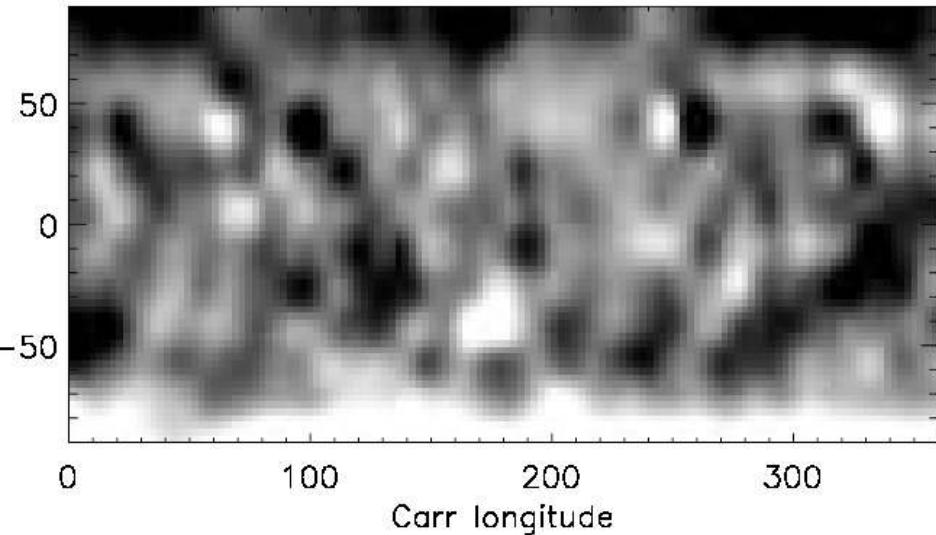
*wind streams measured at
higher latitude / lower latitude*

Synchronic magnetograms

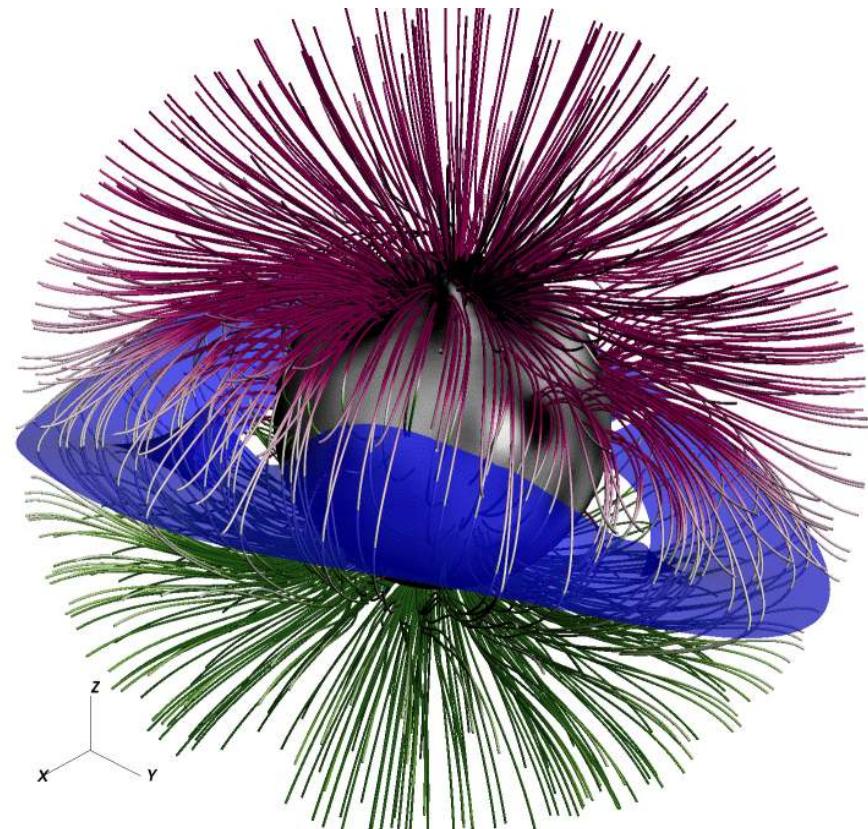
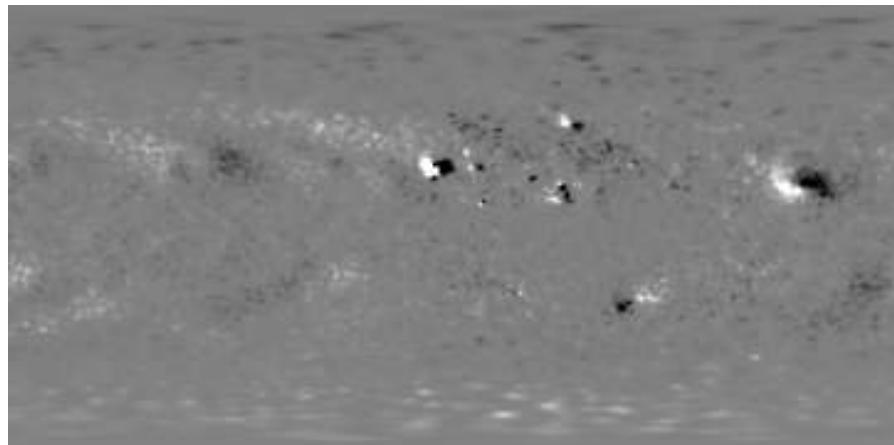
Surface flux-transport methods to evolve magnetograms (full disk)

Simple large-scale shear+diffusion (cf. Wang et al)

wso2079 + 1 d



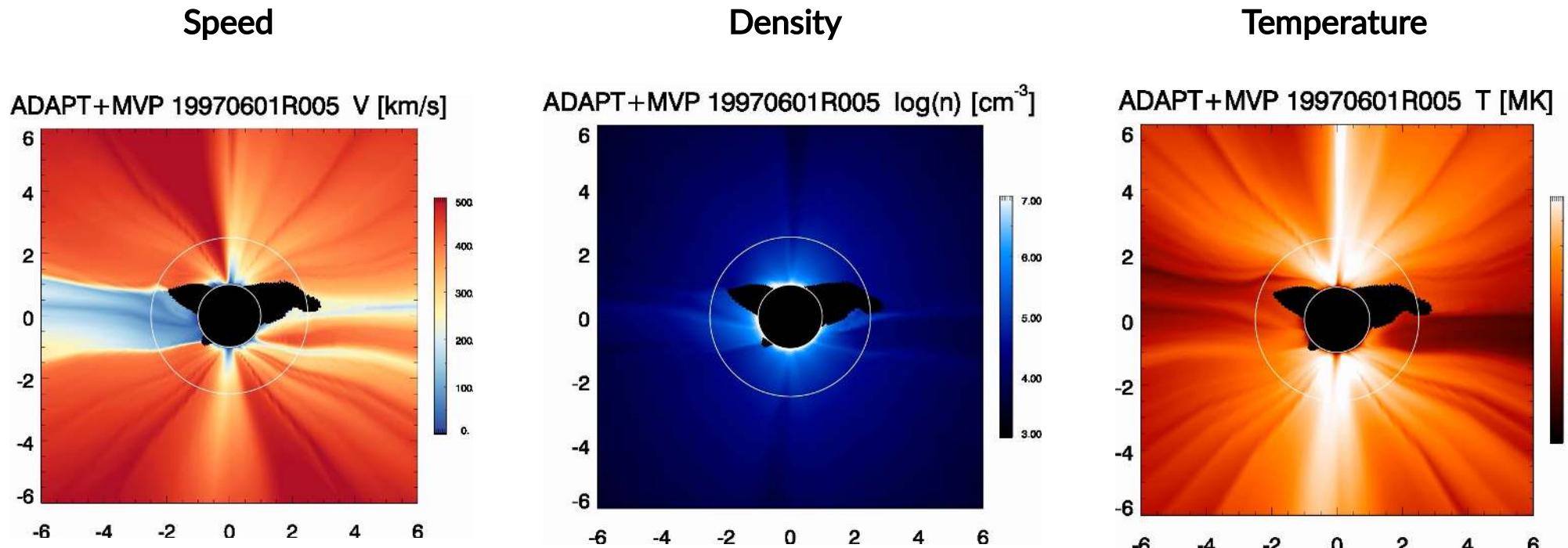
ADAPT ensemble maps (Henney, Arge, et al)



Synchronous wind runs (ADAPT + PFSS-SCS + MULTI-VP)

MULTI-VP (plane-of-sky only)
ADAPT / KPVT maps, PFSS + SCS coronal field reconstruction

1 – 30 June 1997



dark patches: streamers, data gaps
white lines: surface, source-surface

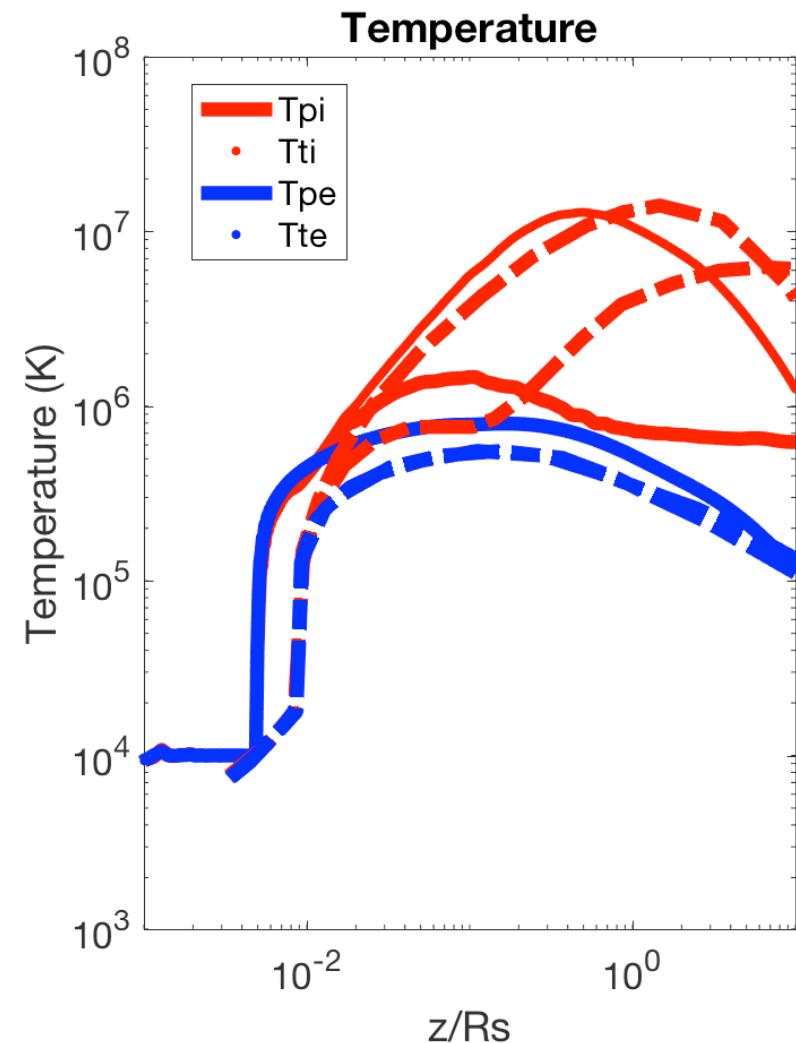
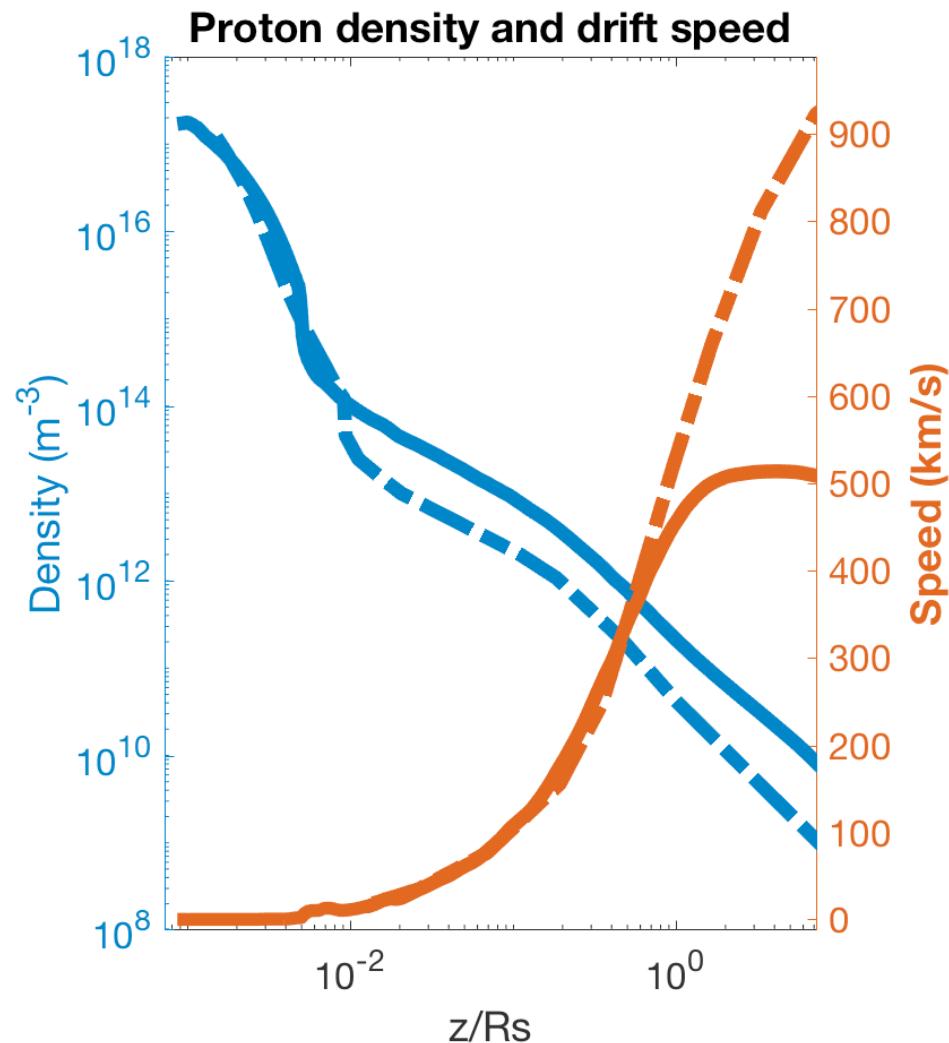
(Pinto, Bemporad, Arge, *in prep*)



- ADAPT-WSA vs. MULTI-VP vs. UVCS data
- synchronous magnetic maps are required
- preparation for Solar Orbiter/METIS

ADAPT magnetogram (Henney, Arge, et al)

Multi-fluid and kinetic fluid solar wind



- better insight into heating mechanisms
- solar wind ion composition
- solar wind sources

(Lavarra, et al, *in prep*)

WIP: Continuous solar wind forecast (7-10 days ahead)

1. Early-on magnetogram data

2. East-limb nowcast

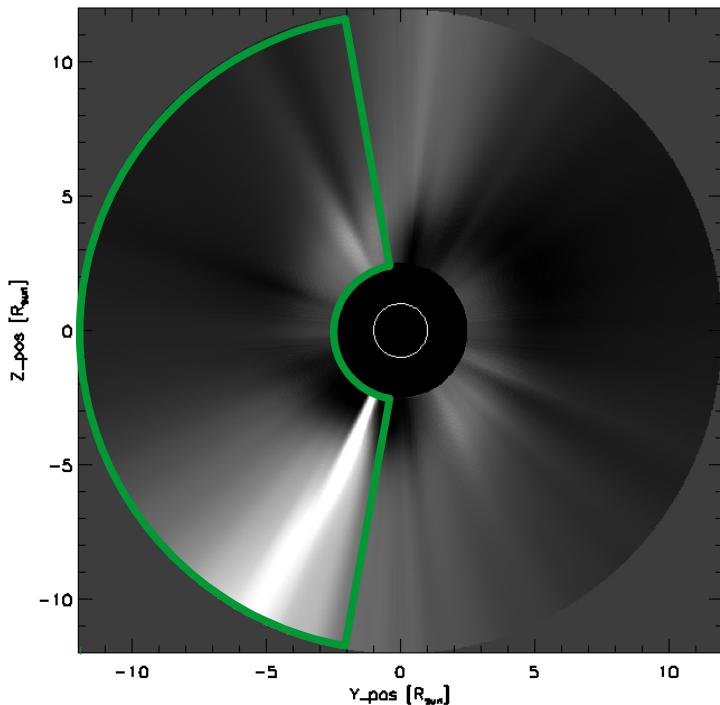
synthetic vs. real coronographic imagery
calibration, re-iteration, forecast quality flags

3. L5 in-situ cross-check

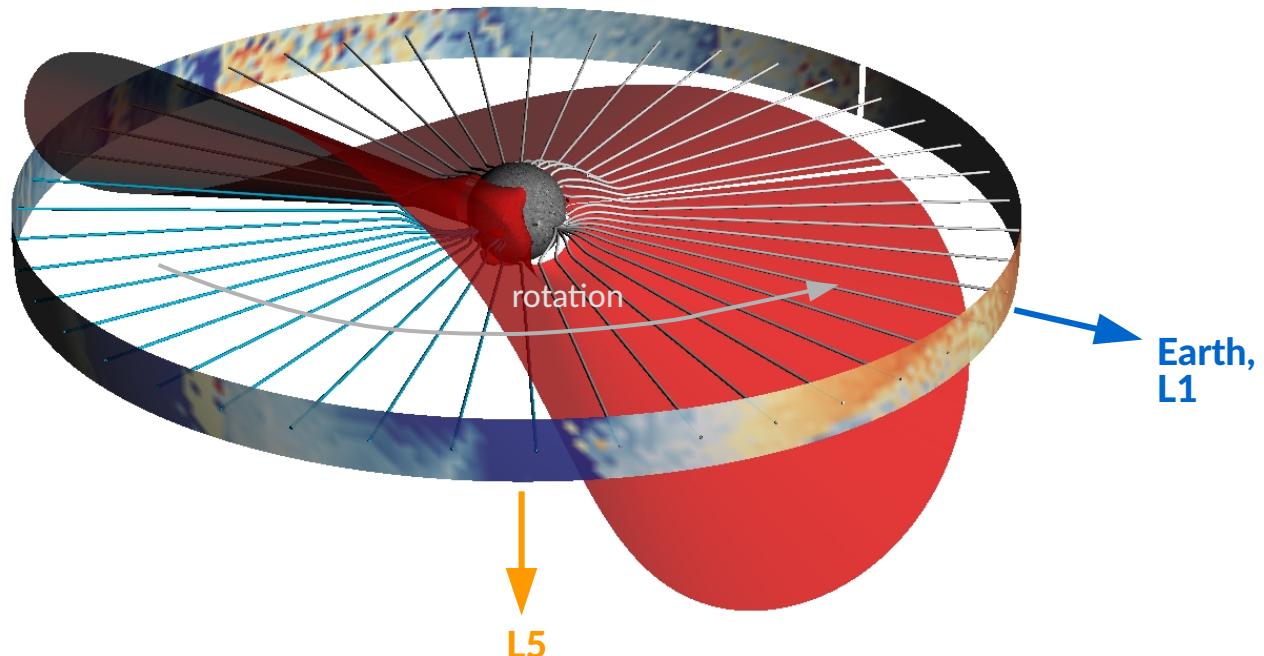
verify and (re-)flag propagated wind solutions

4. Sun-Earth path on west limb

real-time monitoring, Cor/HI imaging



East limb nowcast
(synthetic coron. imaging)



SWiFT pipeline

WIP: Continuous solar wind monitoring and forecasting

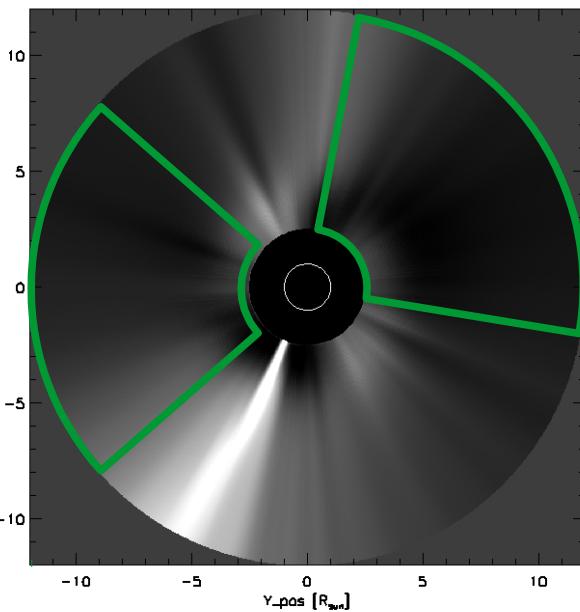
1. Early-on/forecast magnetogram data

2. East/west limb WL nowcast

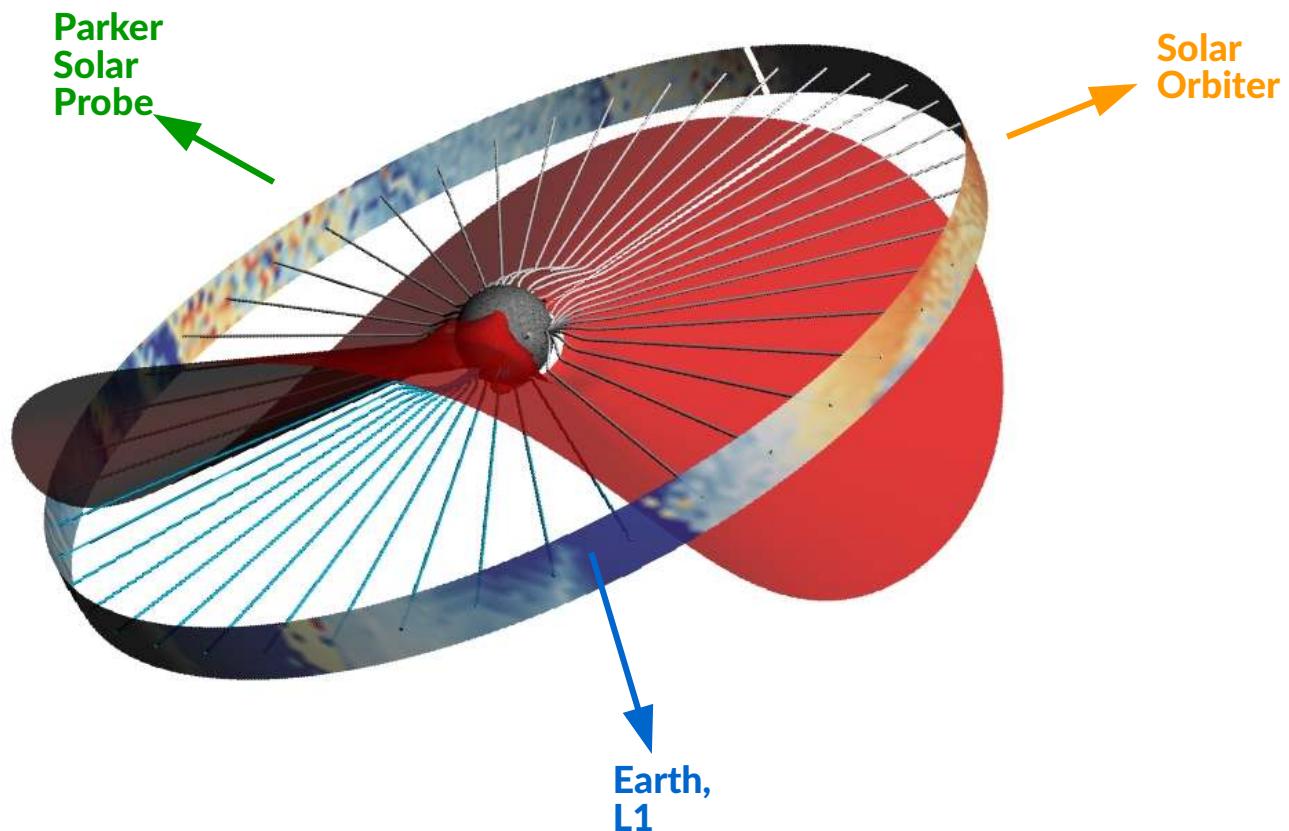
synthetic vs. real coronographic imagery
calibration, re-iteration, forecast quality flags

3. Intermediary in-situ cross-check
verify and (re-)flag propagated wind solutions

4. Sun-spacecraft path on Cor/HI fov
real-time monitoring



East/west limb nowcast
(synthetic coron. imaging)



SWiFT pipeline

Very soon: continuous solar wind monitoring and forecasting

SWIFT-WINDCAST (IRAP / CDPP / STORMS)

► Time Interval

► Model Input

▼ Model Output

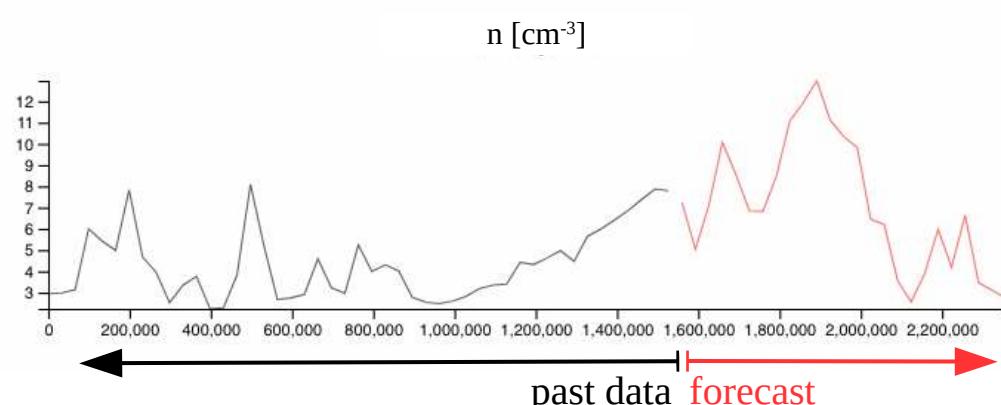
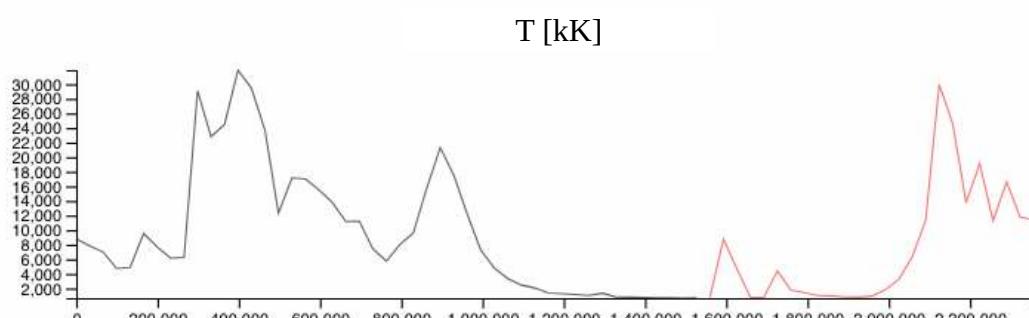
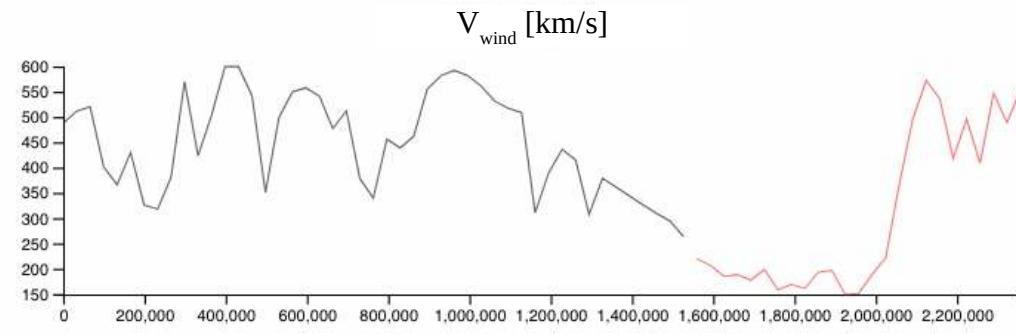
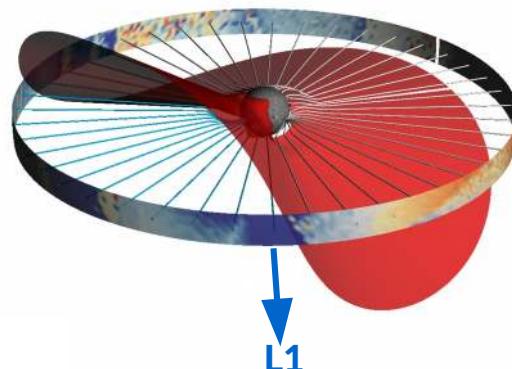
Velocity

Temperature

Density

Magnetic field

► Planets



← past data → forecast



▶ Time Interval

▶ Model Input

▼ Model Output

Dyn. Pressure

Velocity

B Tangential

Temperature

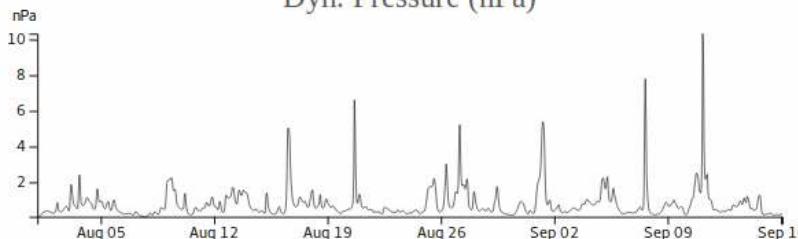
Density

Angle T-S-E

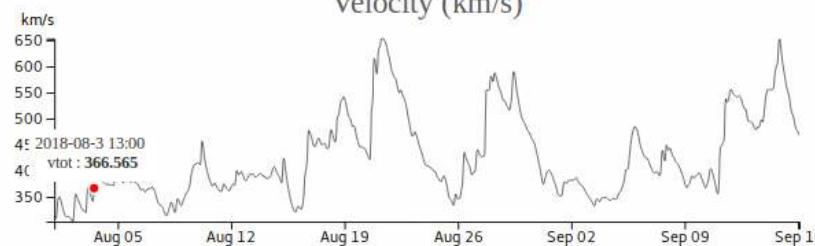
▶ Planets

▶ Pics of Magnetic Fields

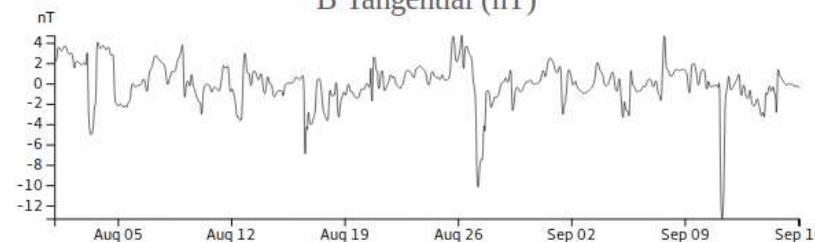
Dyn. Pressure (nPa)



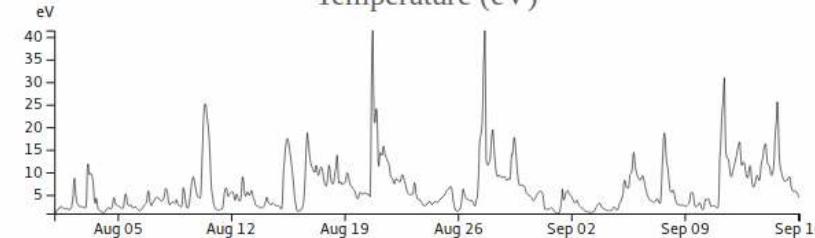
Velocity (km/s)



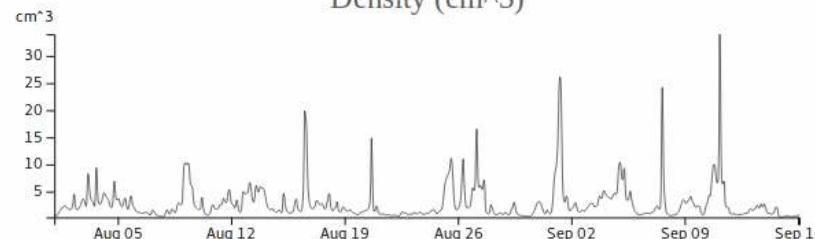
B Tangential (nT)



Temperature (eV)



Density (cm^3)





► Time Interval

► Model Input

▼ Model Output

Dyn. Pressure

Velocity

B Tangential

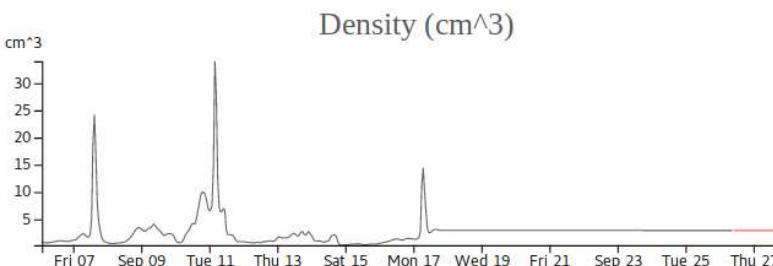
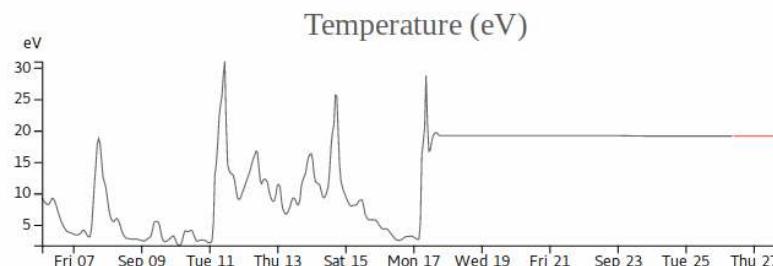
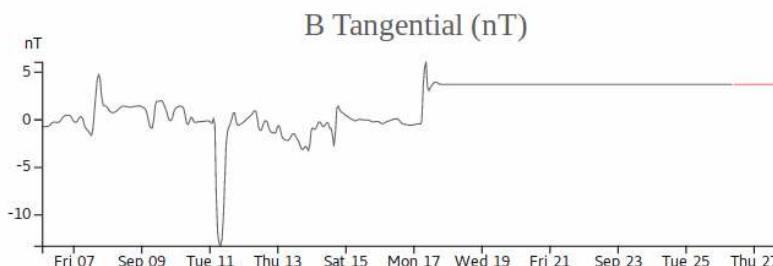
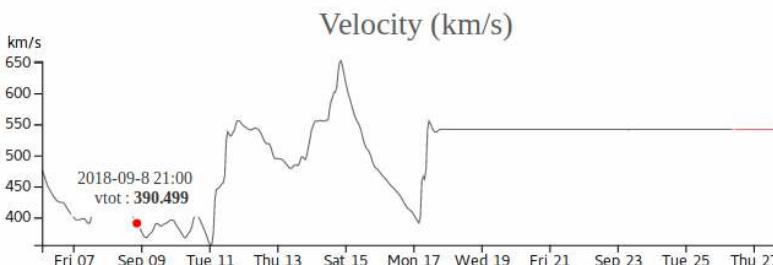
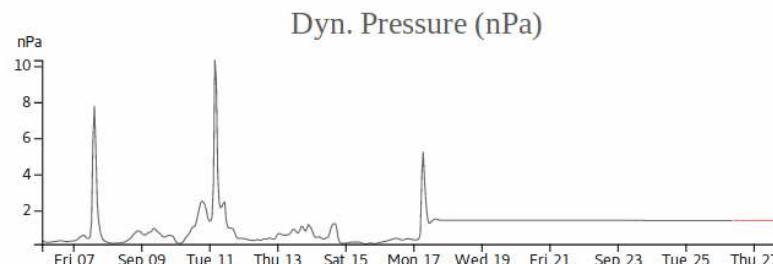
Temperature

Density

Angle T-S-E

► Planets

► Pics of Magnetic Fields



Legend

— In-situ

— Forecast Model

The IRAP and the CDPP

CDPP – Data centre for plasma physics

- European space mission data
- Tools for data analysis
- Services (public and community)

Connectivity Tool

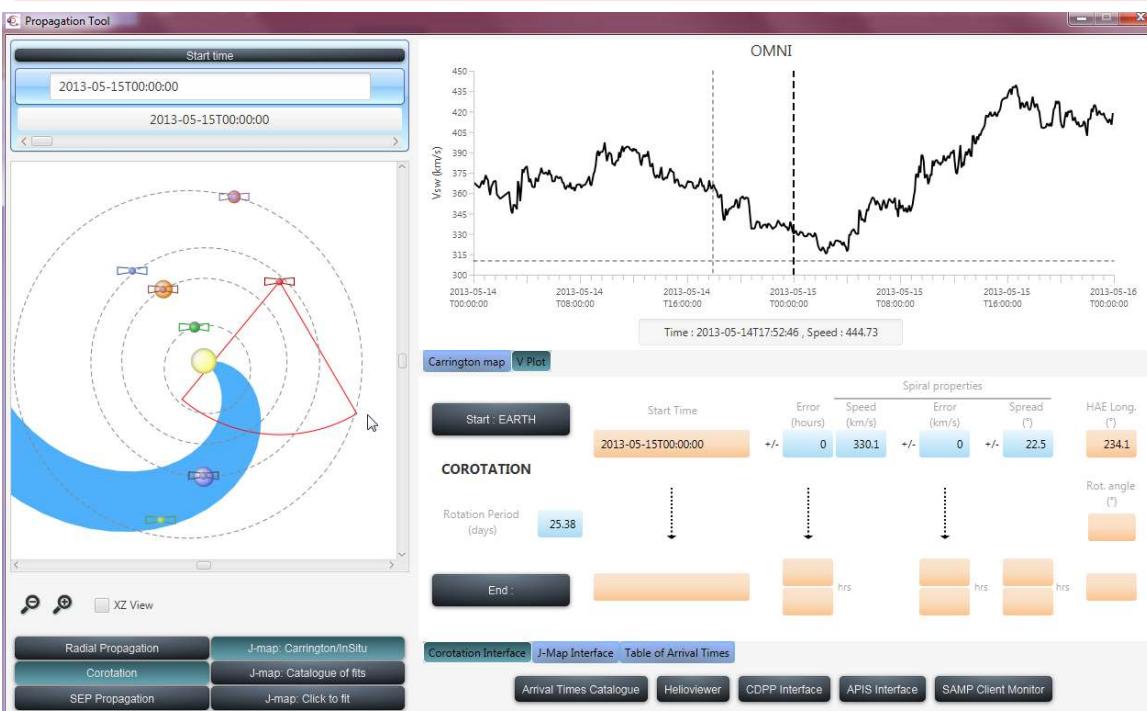
<http://storms-connectsolo.irap.omp.eu>

Sun to s/c connectivity

Propagation Tool

<http://propagationtool.cdpp.eu/>

Propagates solar perturbations : CME, SIR and SEP



SWiFT-WINDCAST (SOON!)

<http://swiftwindcast.irap.omp.eu>

Solar wind forecasts at Earth or s/c

HELIOPROPA

<http://heliopropa.irap.omp.eu/>

Solar wind propag. in-situ to s/c and planets

AMDA

<http://amda.cdp.eu/>

Object-oriented data analysis (space plasmas)

3DView

<http://3dview.cdpp.eu/>

Orbitography, num. simulations, space data

STORMS

<http://stormsweb.irap.omp.eu/>

Solar-Terrestrial ObseRvation and Modelling Service

Conclusions

Connect Tool: Determine surface – s/c connectivity, multiple methods
Testing with SolO orbit, SDO data with 10 yrs offset
Issues: quantify uncertainties, criteria for forecast quality

SWiFT/ MULTI-VP: global wind model ($1 - 32 R_{\text{sun}}$), fast computation,
alternative to semi-empirical (WSA) and to full 3D MHD models

Full set of background solar wind properties,
at all latitudes and azimuths, coronal rotation

Diagnostics (e.g white-light, EUV, in-situ; pre/intermediate/final checkpoints)

Corona to Heliosphere (ENLIL, EUFHORIA, SW1D)

Work-in-progress:

Adding other models: contributions from the community are welcome

Data-driving: synoptic to synchronic magnetic maps, adding local mag maps

Forecast mode: continuous few-days forecast at s/c position

STORMS/SWIFT modélisation X | Home ○ Heliopropia * Solar S X | propagationtool.cdpp.eu X | amda.cdpp.eu X | +

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21

**Amda**

Versatile web tool for
Space Physics

MULTI DATASET VISUALISATION
AND DOWNLOAD

VISUAL AND AUTOMATED EVENT
SEARCH AND DATA MINING

CATALOGUE GENERATION
AND EXPLOITATION

REMOTE ACCESS TO DATA,
MODEL AND IMAGE CENTRES
VIA VO TOOLS AND
STANDARDS

First visit, demo tour

Rules of the road

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**COPP**

Try AMDA as a guest
login: guest
password : your e-mail address

Announcements

09/07/2018

New instance :

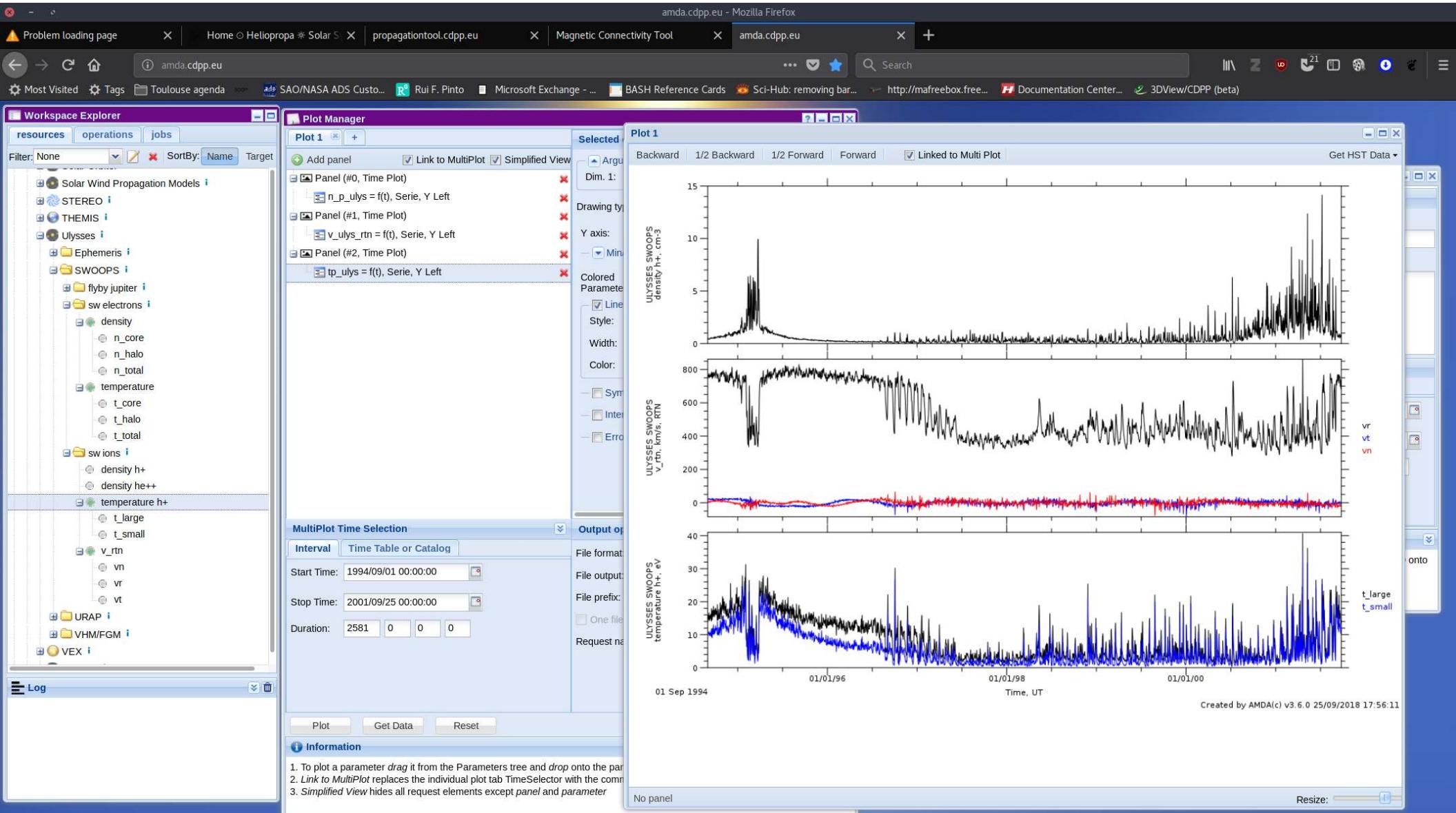
New plot module with more complete options

New analysis functionalities: statistics, catalogues

New data: JUNO, EISCAT, OSCVR

Access to [old version](#)





Welcome to CDPP/Propagation Tool

Tutorials : video (mov files)

- Introduction to the CDPP Propagation Tool (13M)
- Description of the propagation tool main interface (8M)
- Case 1: Using the tool in the Jmap Carrington/in situ mode (radial) (37M)
- Case 2: Using the tool in the Jmap tool click mode (radial) (39M)

Tutorials : video (mpeg files)

- Introduction to the CDPP Propagation Tool (46M)
- Description of the propagation tool main interface (47M)
- Case 1: Using the tool in the Jmap Carrington/in situ mode (radial) (176M)
- Case 2: Using the tool in the Jmap tool click mode (radial) (184M)

Table of available data

- Flare Data, Carrington Maps, J-Maps, Solar Wind Speed

Supported set up

- Java requirements
- Get Java 8
- Linux troubleshoot

What's new ?

[Launch the Propagation Tool](#)

java 8: Mac users: please use Firefox or Safari

A new interactive tool accessible to the solar, heliospheric and planetary science communities to track solar storms, streams and energetic particles in the heliosphere. This tool was defined and developed by IRAP and IAS staff through a subcontract with GFI informatique and CNES financial support. It follows on from and is complementary to the propagation tool developed by the FP7 HELIO project.

Main functionalities are described in the reference paper recently published in Planetary and Space Science.

The propagation tool allows users:

- to propagate solar eruptions (CMEs) radially sunward or anti-sunward (**Radial Propagation**),
- to propagate corotating structures (CIRs) in the heliosphere (**Corotation**),
- to propagate solar energetic particles along magnetic fields lines sunward or anti-sunward (**SEP Propagation**).

The START and END points can be the Sun, planets or probes situated in the interplanetary medium. The times of propagation between the START and END points are based on simple analytic calculations.

The added values of the tool are an easy access to unique datasets and a fast interoperability :

- it integrates the **orbital elements** (using SPICE) of probes and planets. This allows you to determine via simple clicks the position/orientations of imagers that you would like to consider,
- it offers **web-service** access to summary plots of in-situ data stored at the CDPP as well as movies of solar images stored at MEDOC,
- it provides access to a wide range of **Carrington maps** of the solar surface to visualize the location of active regions, coronal holes and solar flares on the Sun

The great novelty of the tool is the immediate visualisation and basic manipulation of maps of solar wind mass flows tracked continuously from the Sun to 1AU. These maps are called J-maps and are generated by extracting bands of pixels in coronal and heliospheric images along the ecliptic planes and stacking them vertically (along the ordinate) with time (along the abscissae). The maps are produced from teraoctets of imagery data that are impossible to manipulate if you are not an expert in the field. The tool was designed to be user friendly and accessible to any scientist interested in locating CMEs/CIRs and particle fluxes in the ecliptic plane.

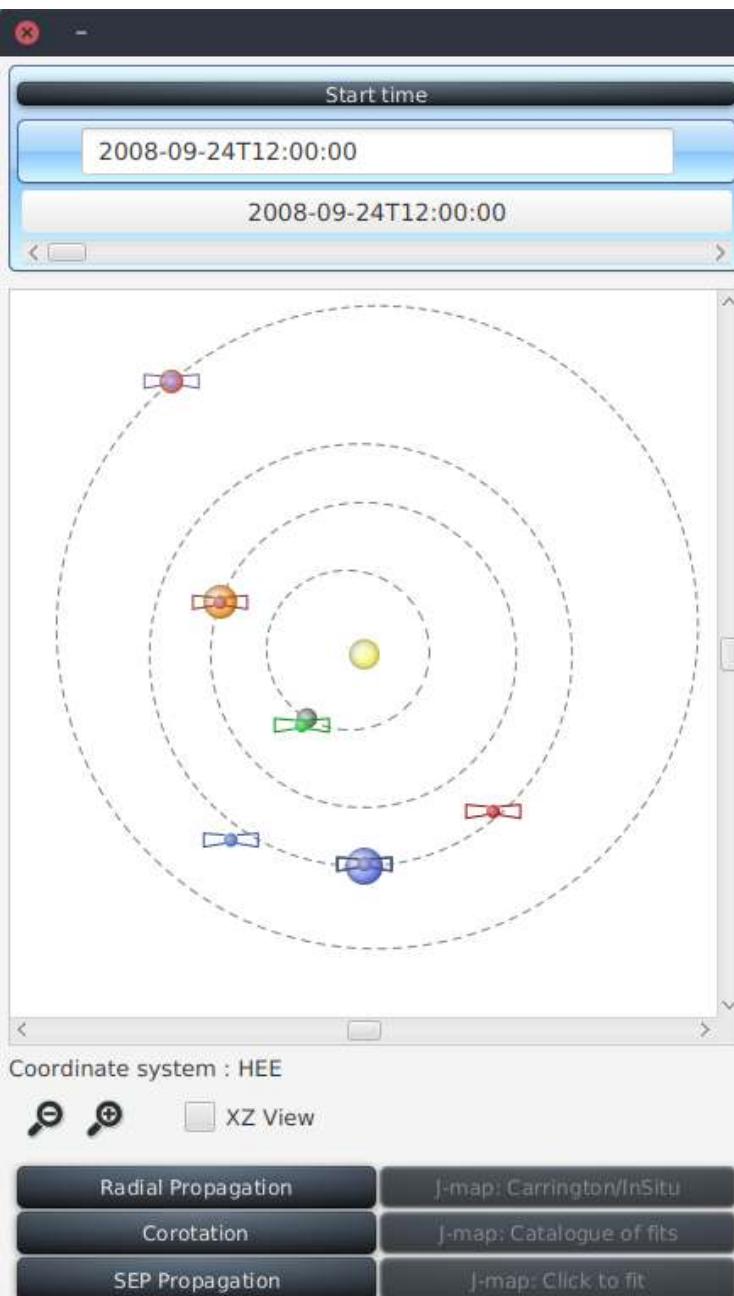
With the tool you can use these maps to:

- cross check your ballistic calculation of CME/CIR propagations,
- carry out your own calculations of CME/CIR trajectories in the ecliptic plane via a few clicks on the map (simple use),
- use pre-calculated CME trajectories to check if a transient emerged from the Sun and impacted a planet or probe

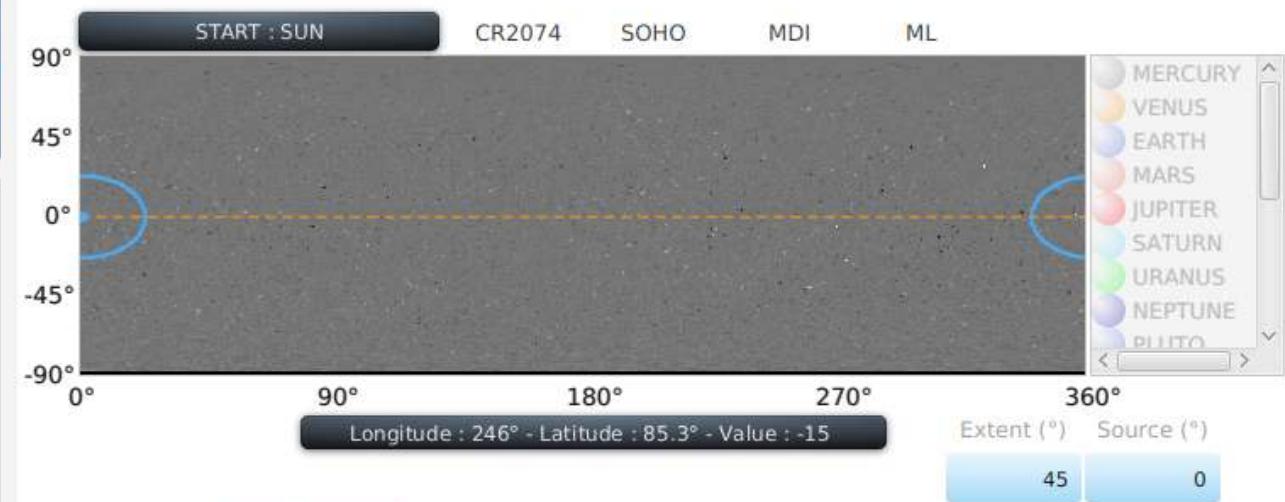
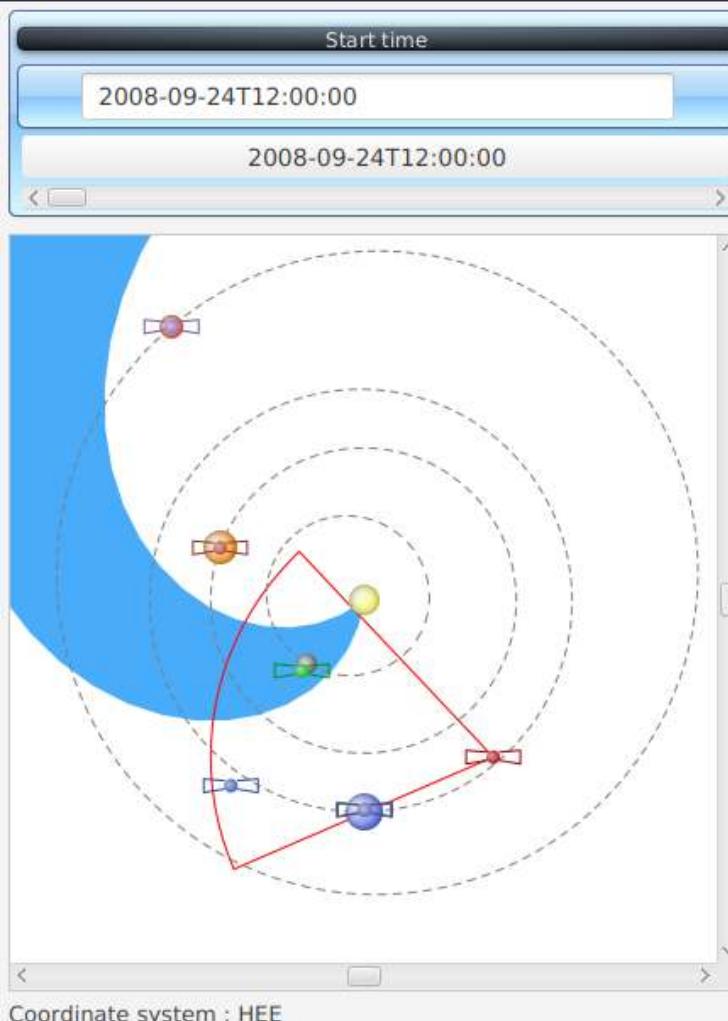


Contact





Propagation Tool



Carrington map V Plot Flux Plot

Start : SUN Start Time 2008-09-24T12:00:00 Error (hours) +/- 0 Speed (km/s) 500 Error (km/s) +/- 0 Spread (°) +/- 22.5 Spiral properties HAE Long. (°) 332.7

COROTATION

Rotation Period (days)

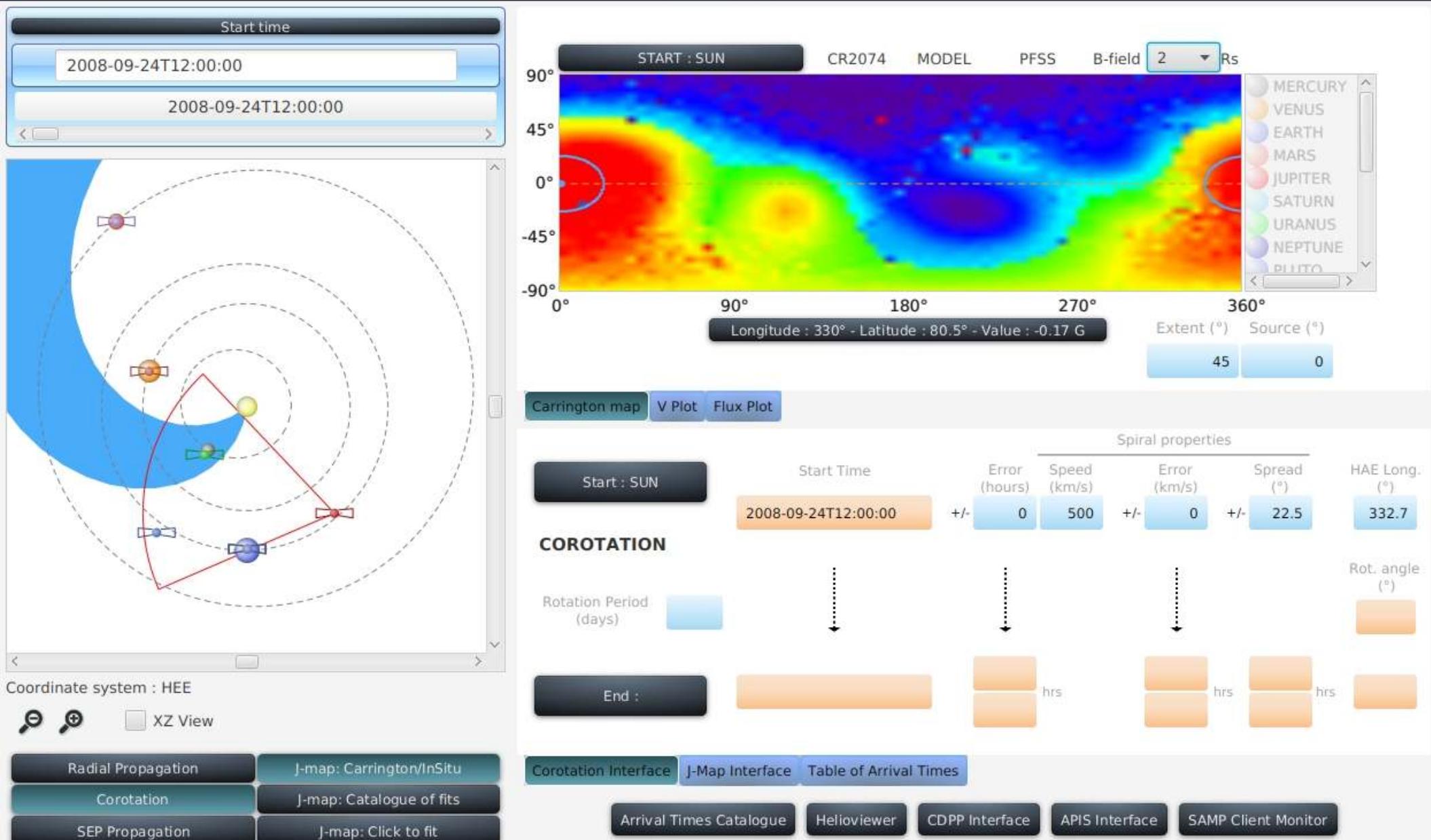
End : hrs hrs hrs hrs

Rot. angle (°)

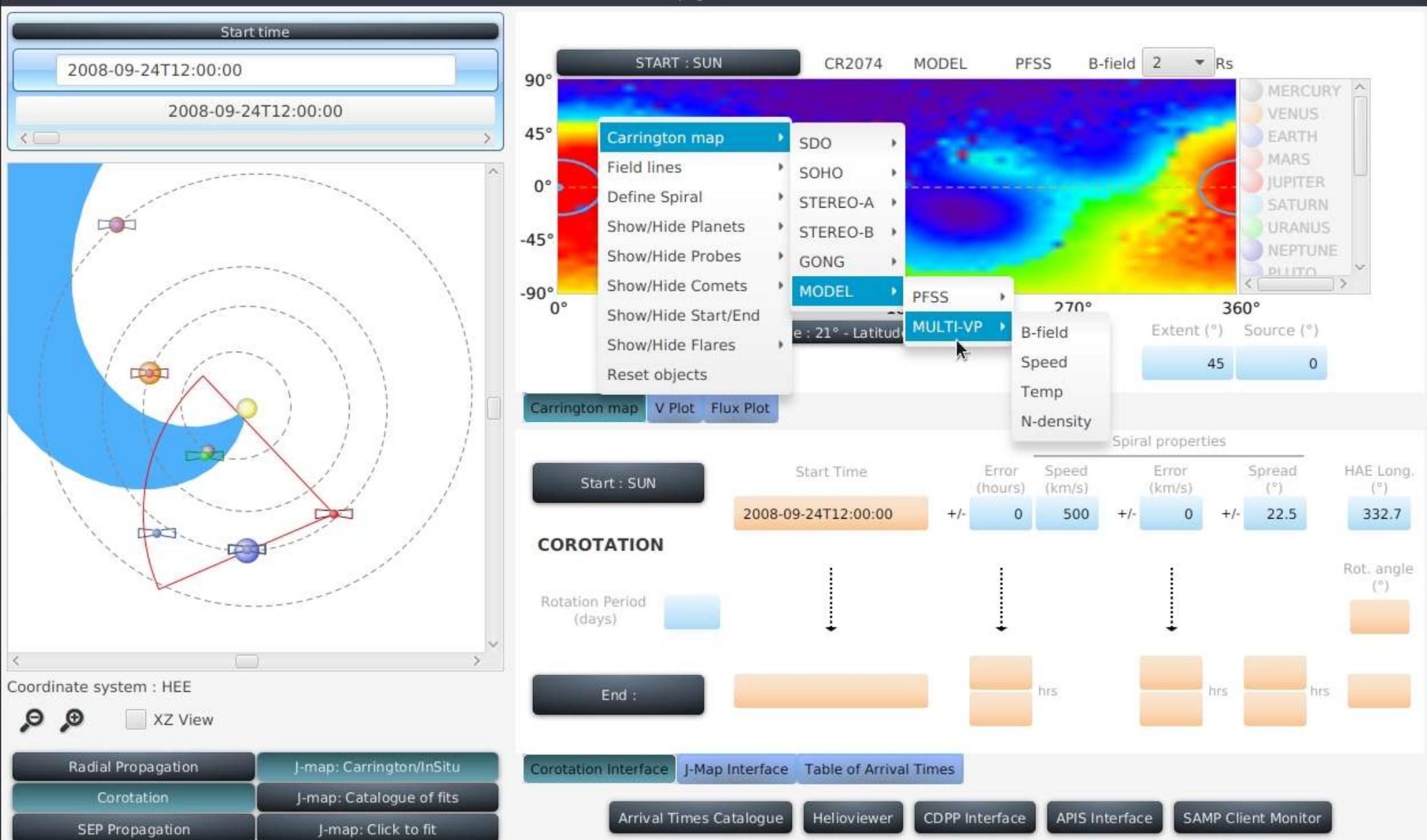
Corotation Interface J-Map Interface Table of Arrival Times

Arrival Times Catalogue Helioviewer CDPP Interface APIS Interface SAMP Client Monitor

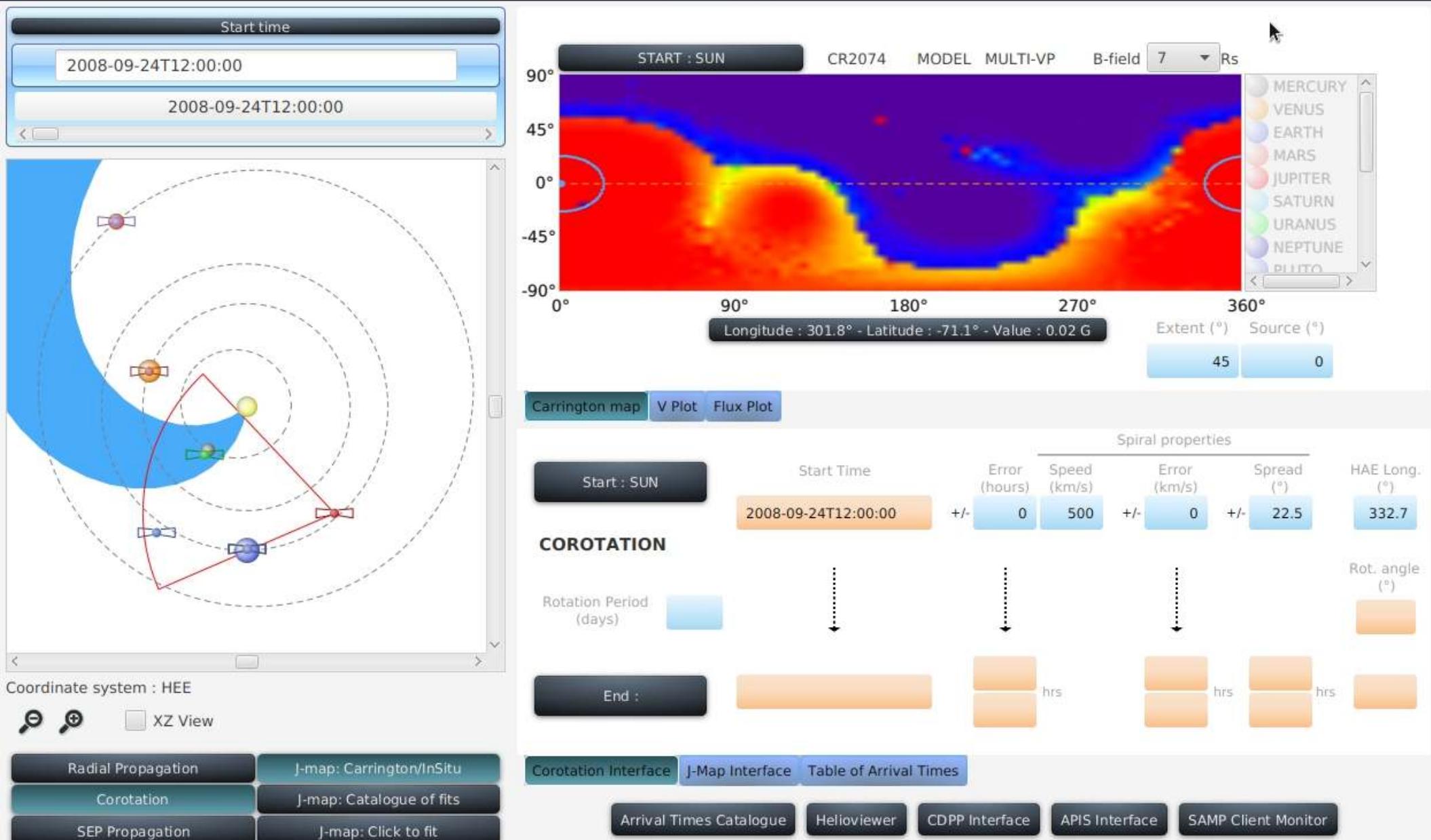
Propagation Tool



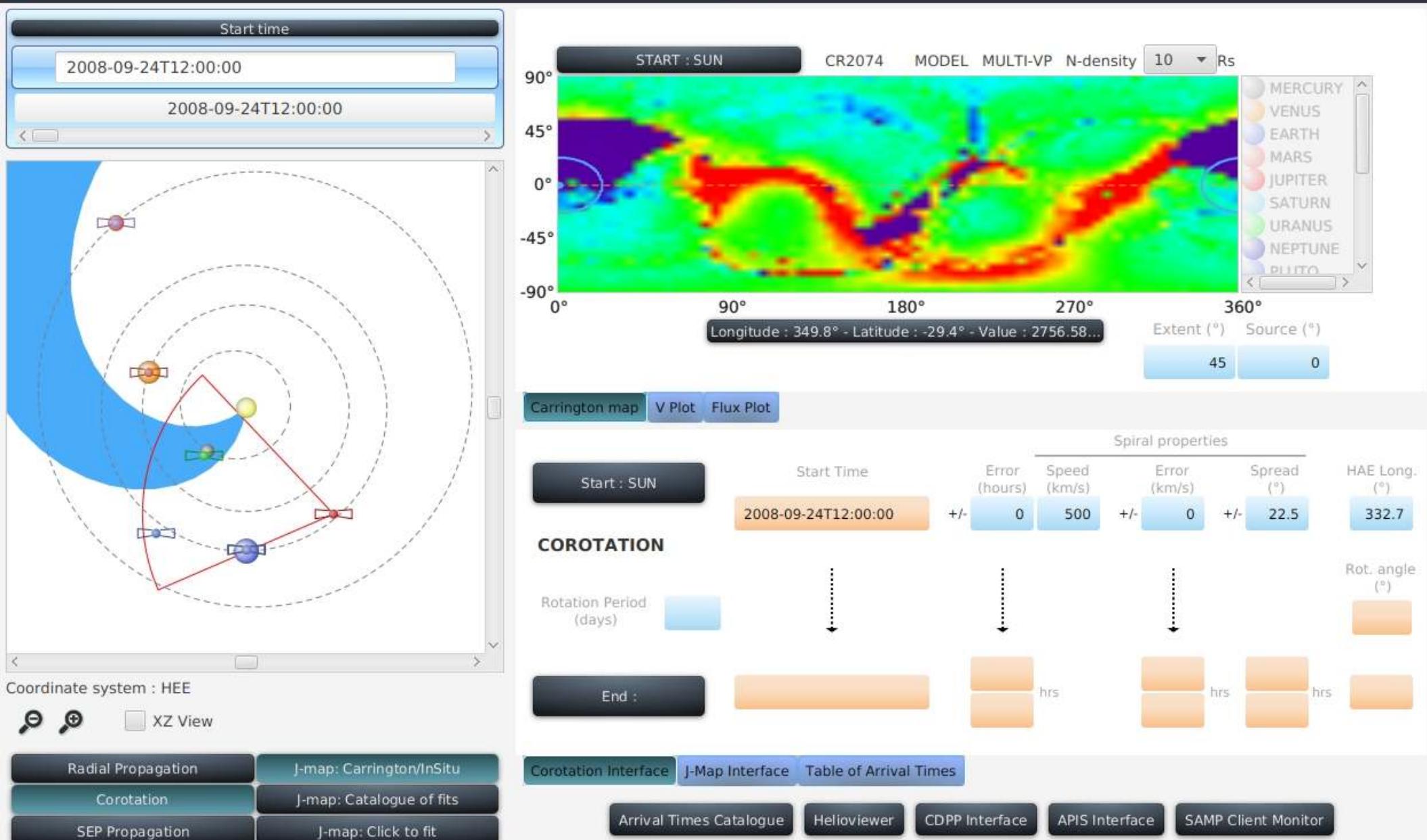
Propagation Tool



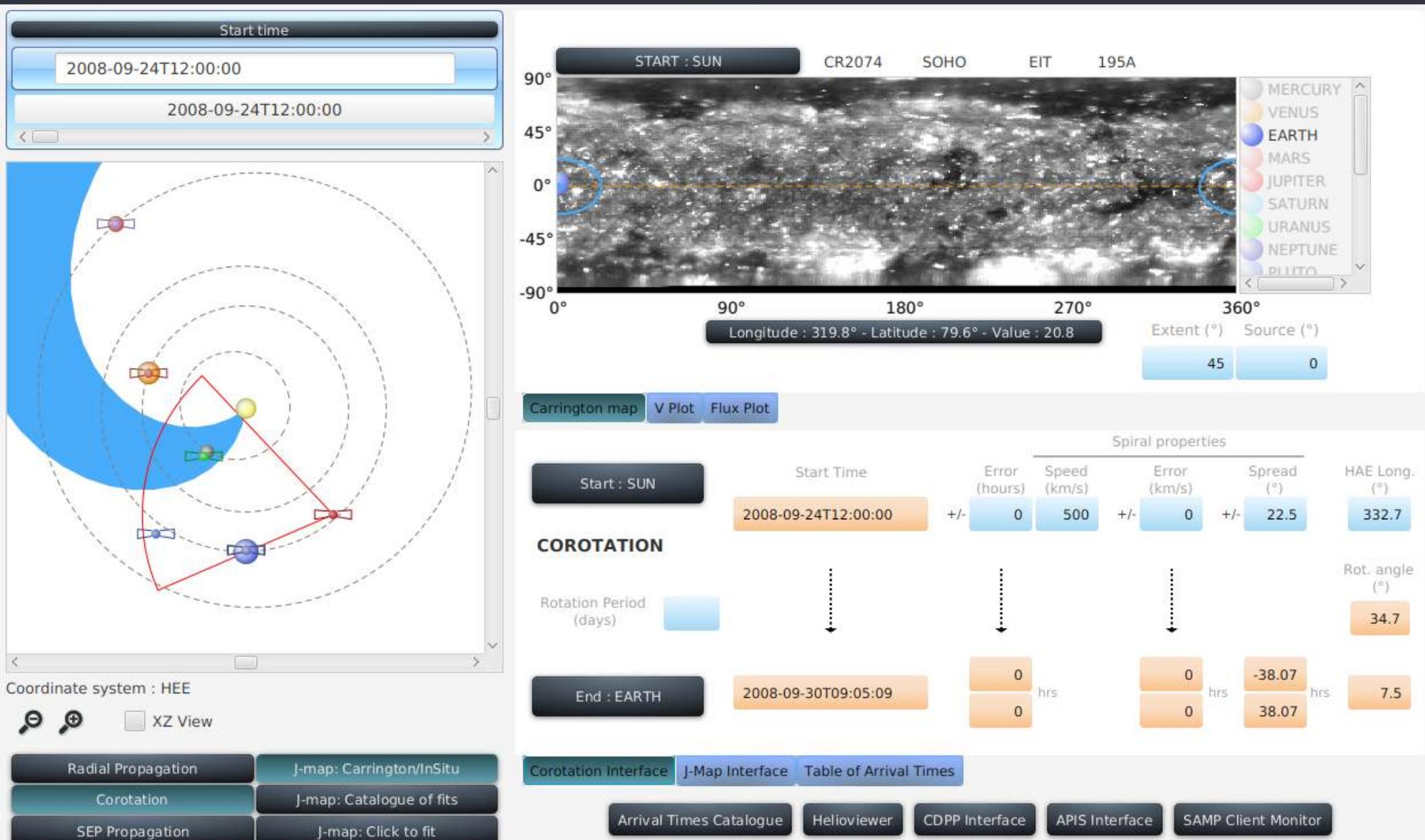
Propagation Tool

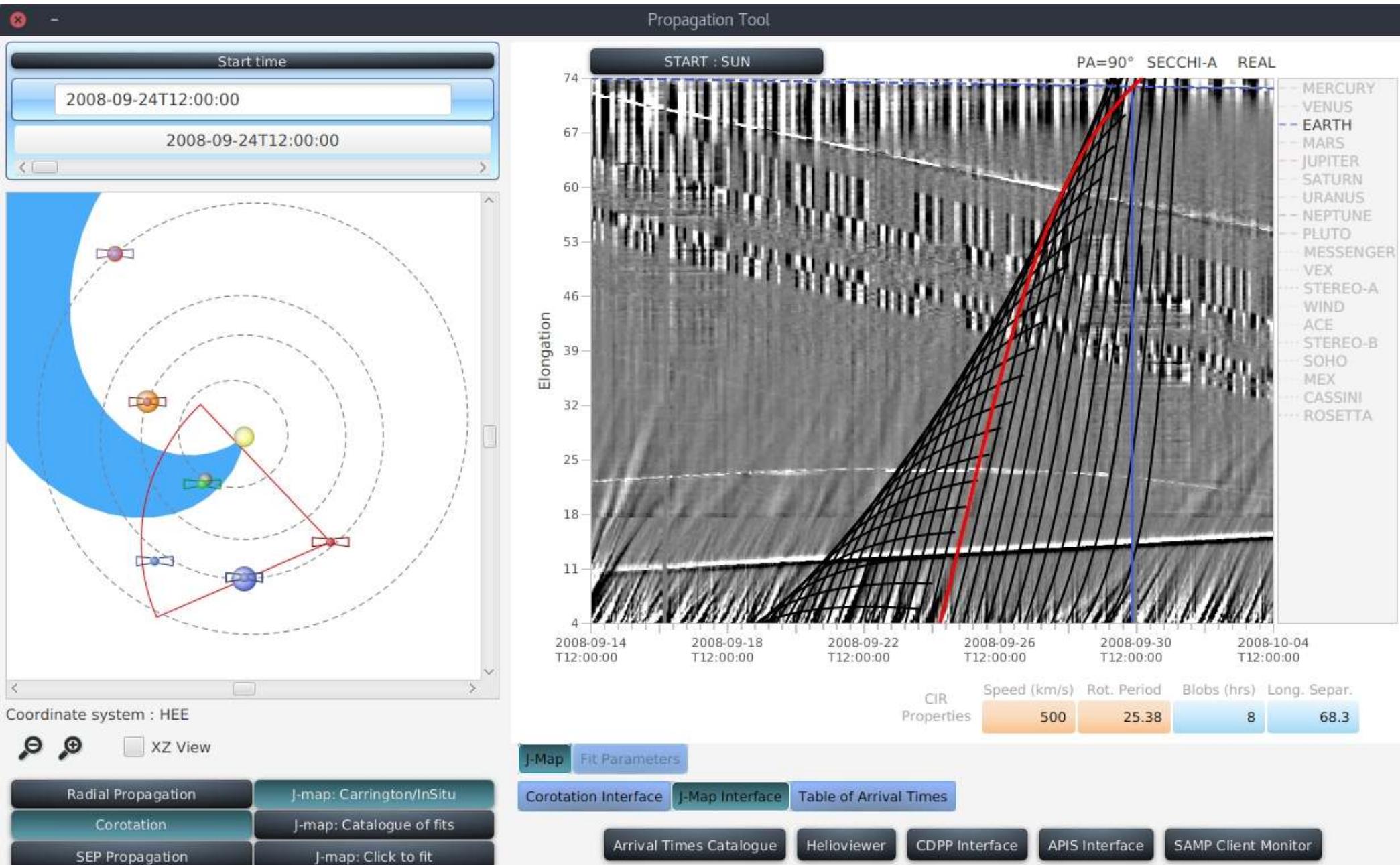


Propagation Tool

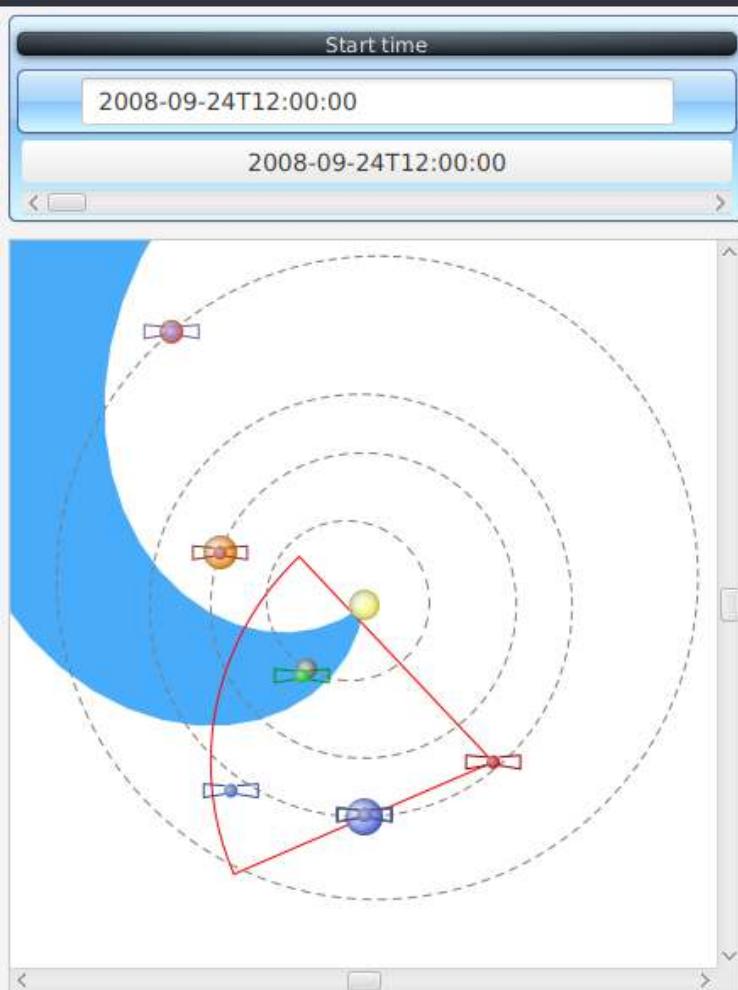


Propagation Tool





Propagation Tool



Coordinate system : HEE



XZ View

Radial Propagation

J-map: Carrington/InSitu

Corotation

J-map: Catalogue of fits

SEP Propagation

J-map: Click to fit

Target	t'	t'min(Δt)	t'max(Δt)	t'min(ΔV)	t'max(ΔV)	t'min(Δφ)	t'max(Δφ)	φEnd(t'...)
SUN	2008-09-24T12:00:00	0	0	0	0	-38.07	38.07	0
Probes								
MESSEN...	2008-09-25T07:54:00	0	0	0	0	-38.07	38.07	350.38
VEX	2008-10-19T03:38:48	0	0	0	0	-38.07	38.07	318.16
STEREO-A	2008-10-03T05:33:18	0	0	0	0	-38.07	38.07	77.44
WIND	2008-09-30T08:03:39	0	0	0	0	-38.07	38.07	34.53
ACE	2008-09-30T08:08:57	0	0	0	0	-38.07	38.07	34.68
STEREO-B	2008-09-27T22:56:18	0	0	0	0	-38.07	38.07	356.24
SOHO	2008-09-30T08:40:44	0	0	0	0	-38.07	38.07	34.96
MEX	2008-10-17T23:10:42	0	0	0	0	-38.07	38.07	255.74
CASSINI	2008-10-14T20:53:48	0	0	0	0	-38.07	38.07	190.48
ROSETTA	2008-09-26T13:45:31	0	0	0	0	-38.07	38.07	282.63
Planets								
MERCURY	2008-09-25T03:50:59	0	0	0	0	-38.07	38.07	349.99
VENUS	2008-10-19T03:38:18	0	0	0	0	-38.07	38.07	318.16
EARTH	2008-09-30T09:05:09	0	0	0	0	-38.07	38.07	34.72
MARS	2008-10-17T23:10:32	0	0	0	0	-38.07	38.07	255.74
JUPITER	2008-10-09T15:57:17	0	0	0	0	-38.07	38.07	322.4
SATURN	2008-10-14T20:57:26	0	0	0	0	-38.07	38.07	190.49
URANUS	2008-10-14T14:26:50	0	0	0	0	-38.07	38.07	18.14
NEPTUNE	2008-09-26T06:56:13	0	0	0	0	-38.07	38.07	350.24
PLUTO	2008-09-27T17:39:58	0	0	0	0	-38.07	38.07	297.55
Comets								

Given defined width, targets in red are impacted by Spiral

Corotation Interface J-Map Interface Table of Arrival Times

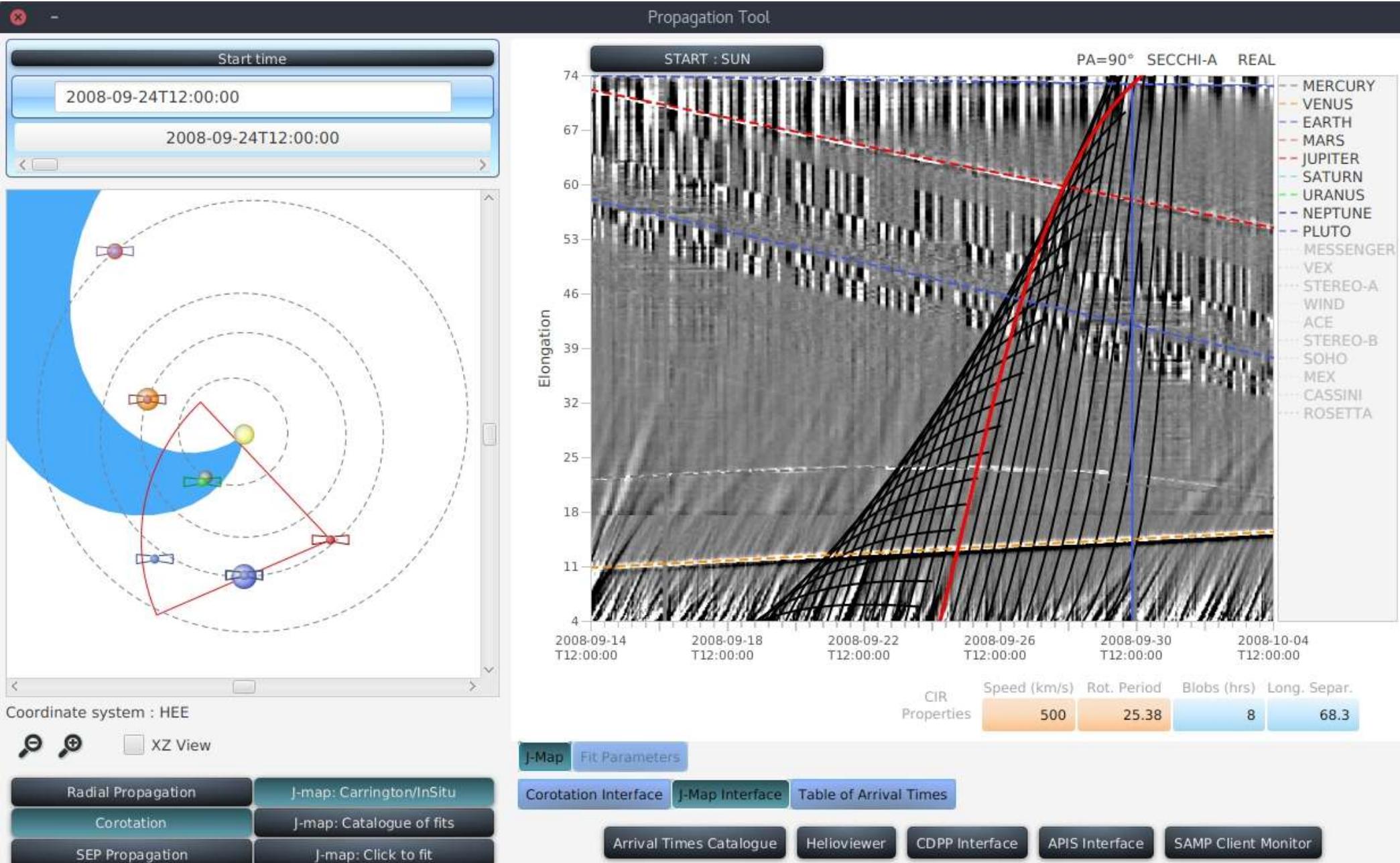
Arrival Times Catalogue

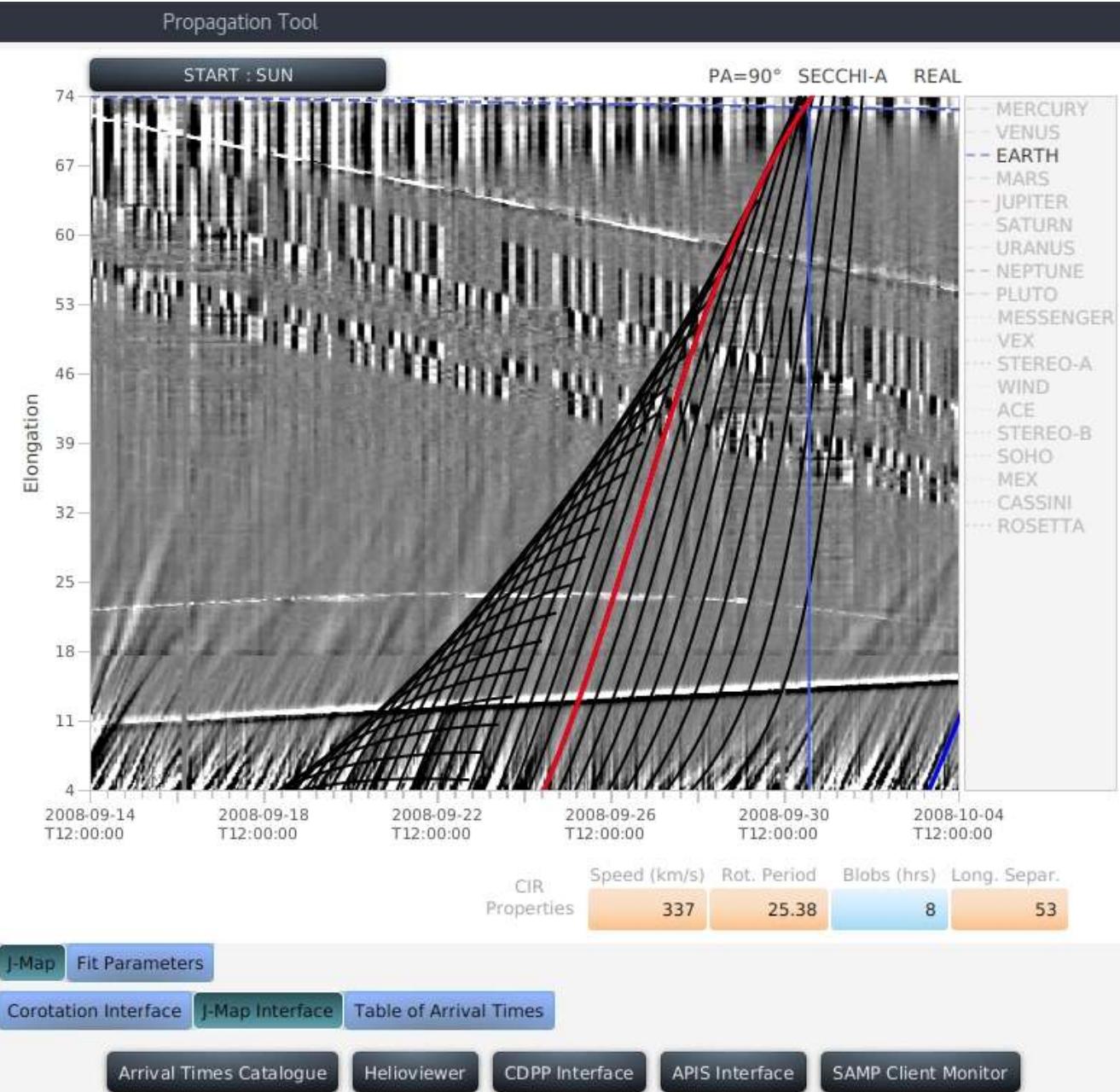
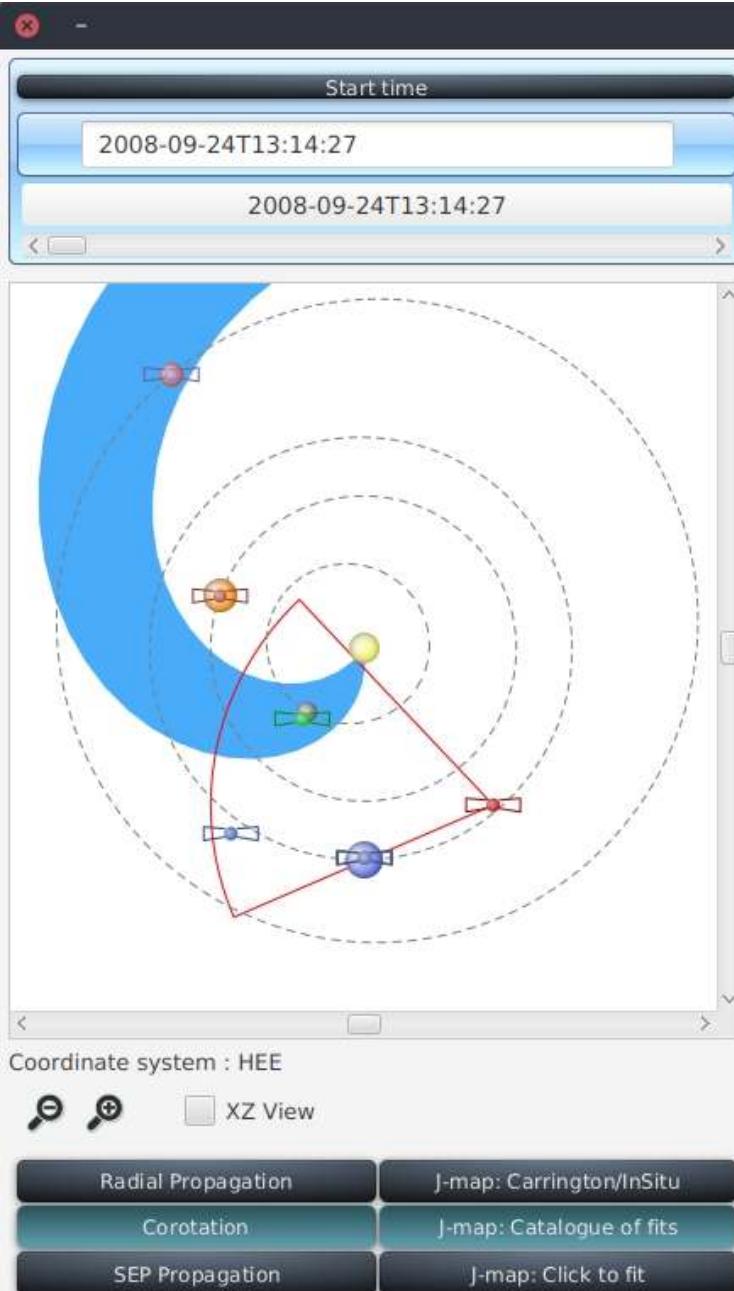
Helioviewer

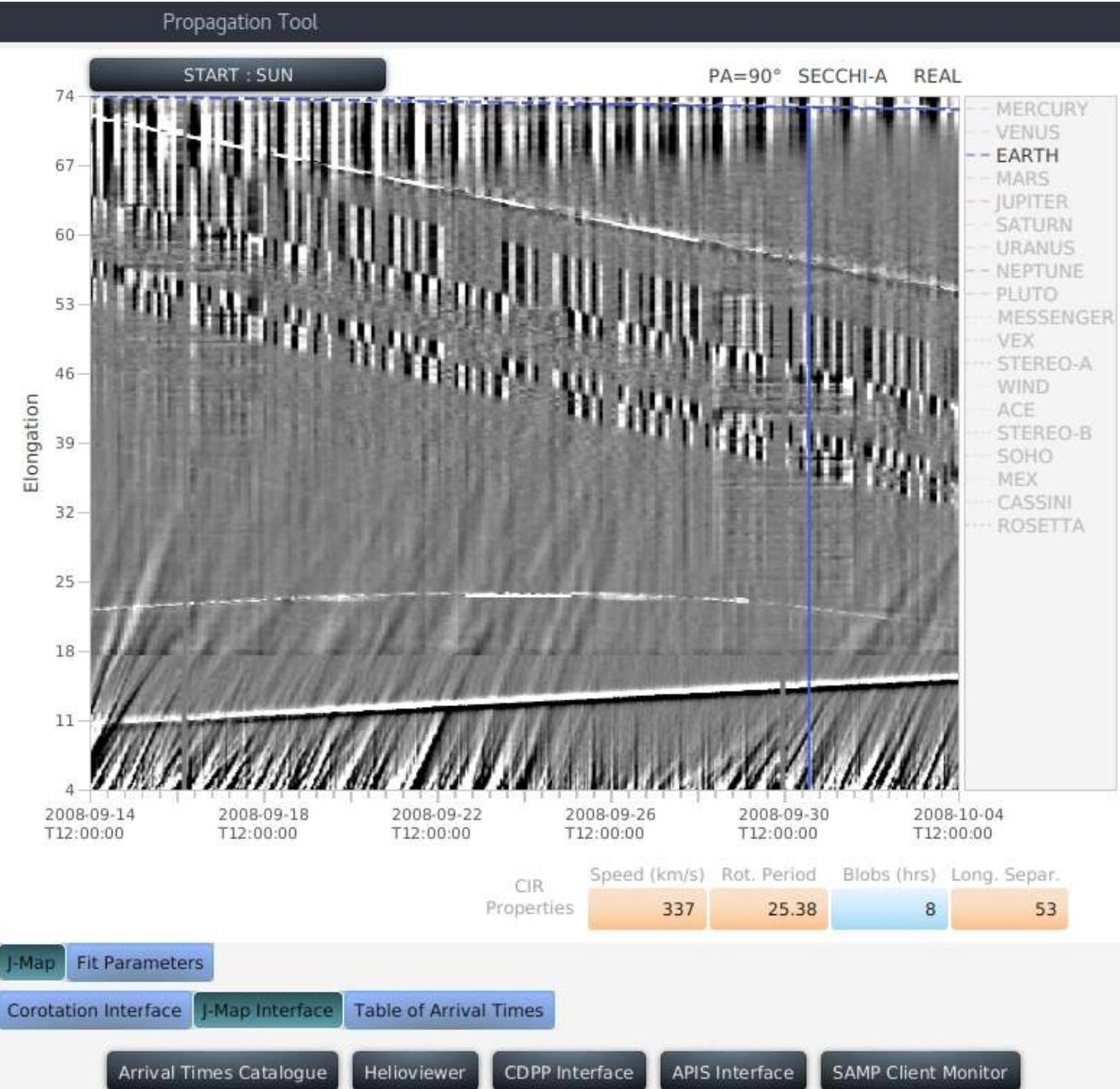
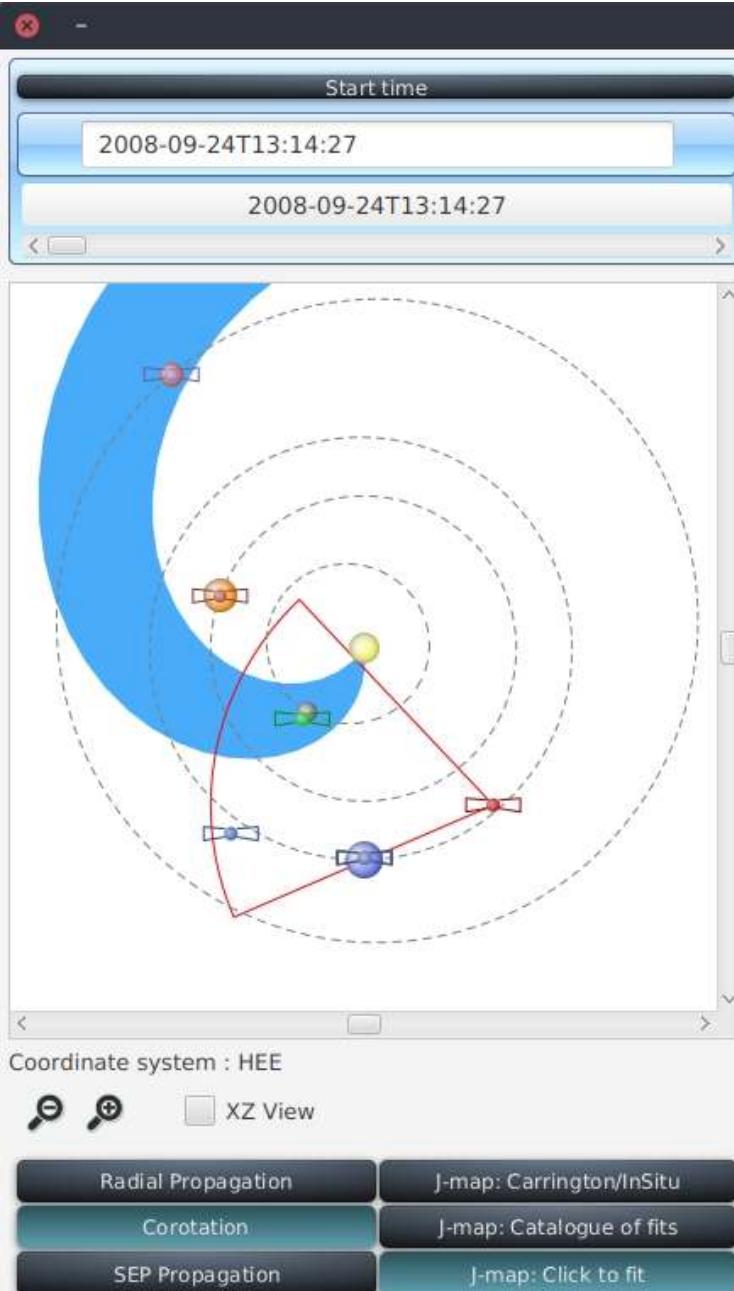
CDPP Interface

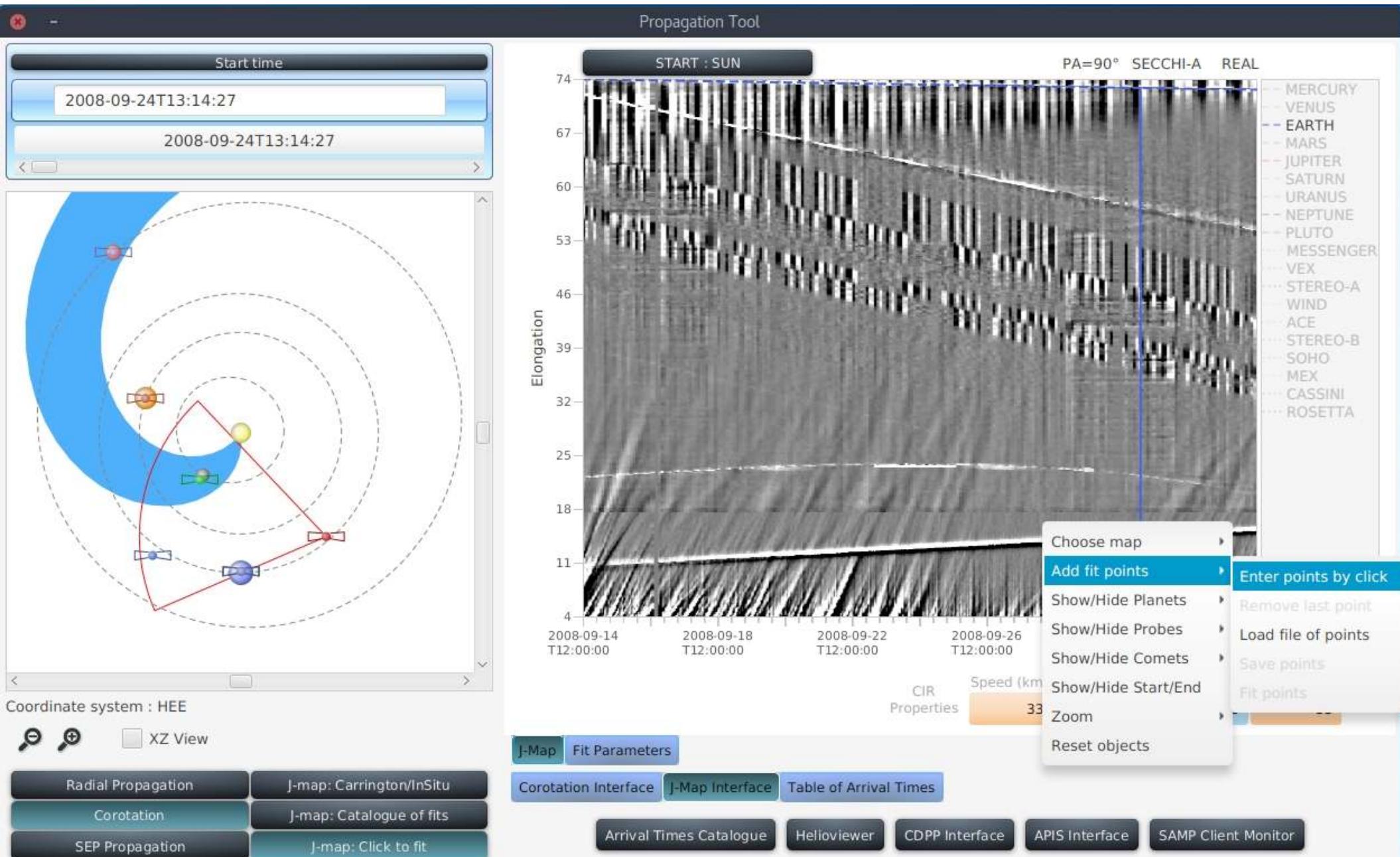
APIS Interface

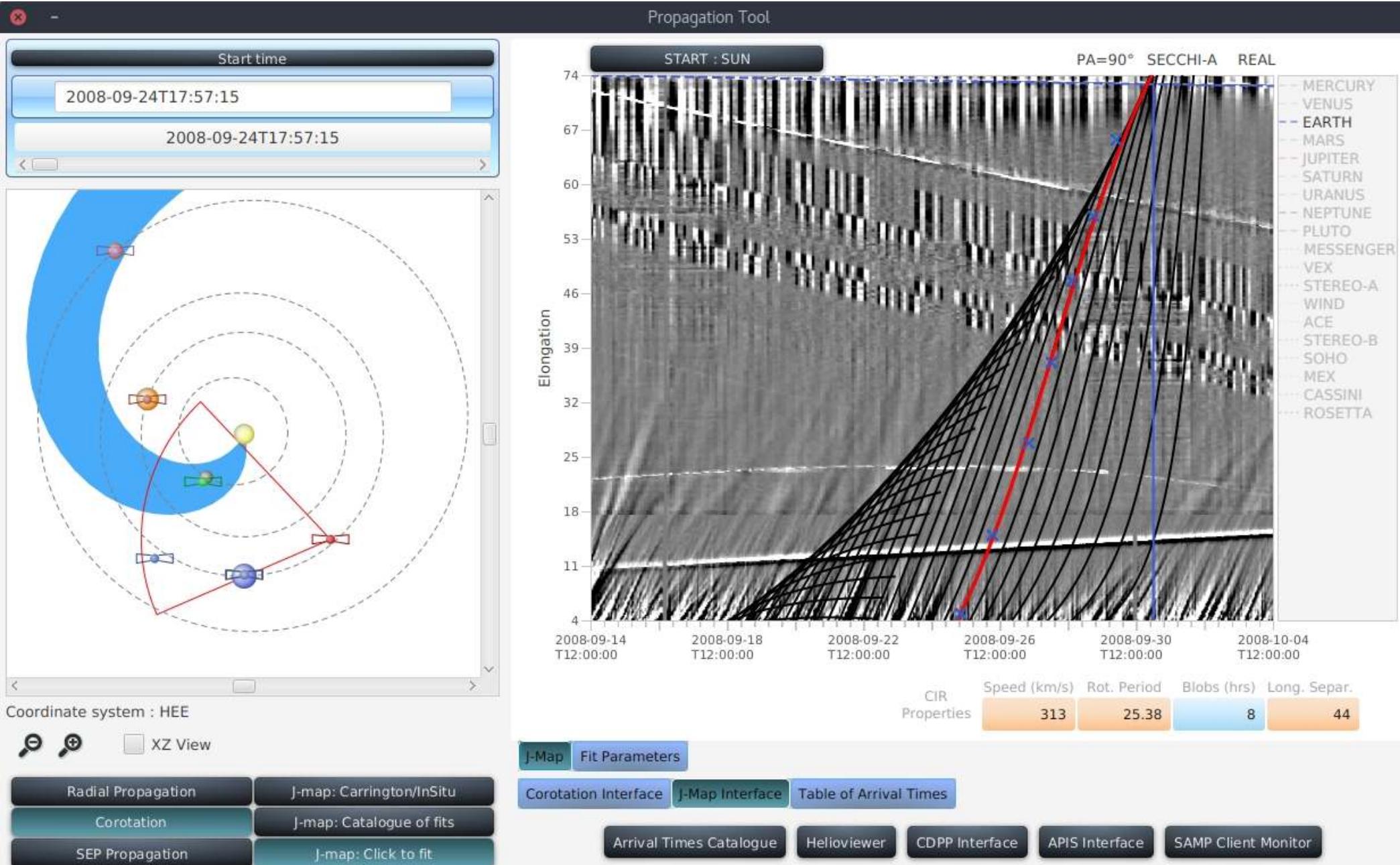
SAMP Client Monitor



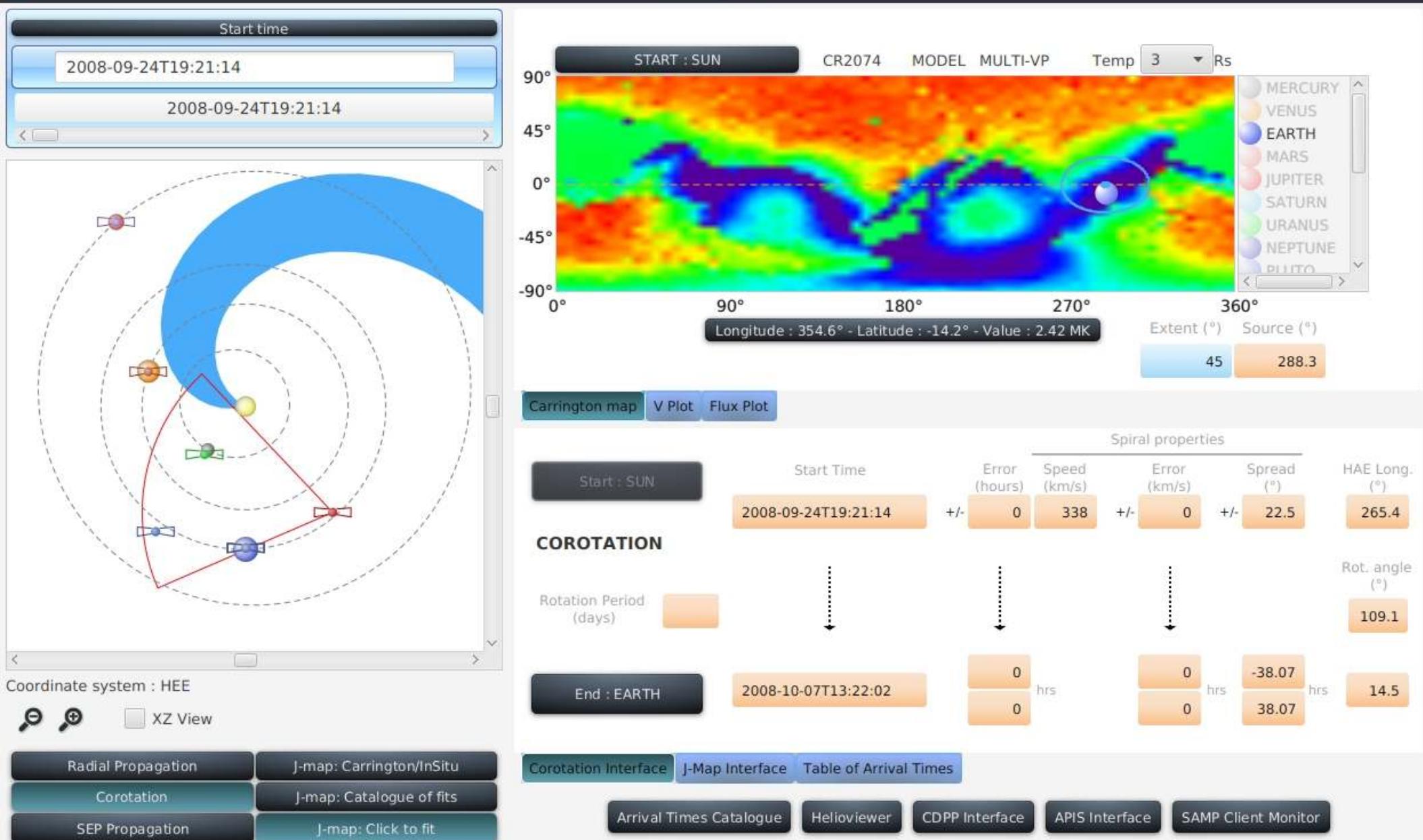








Propagation Tool

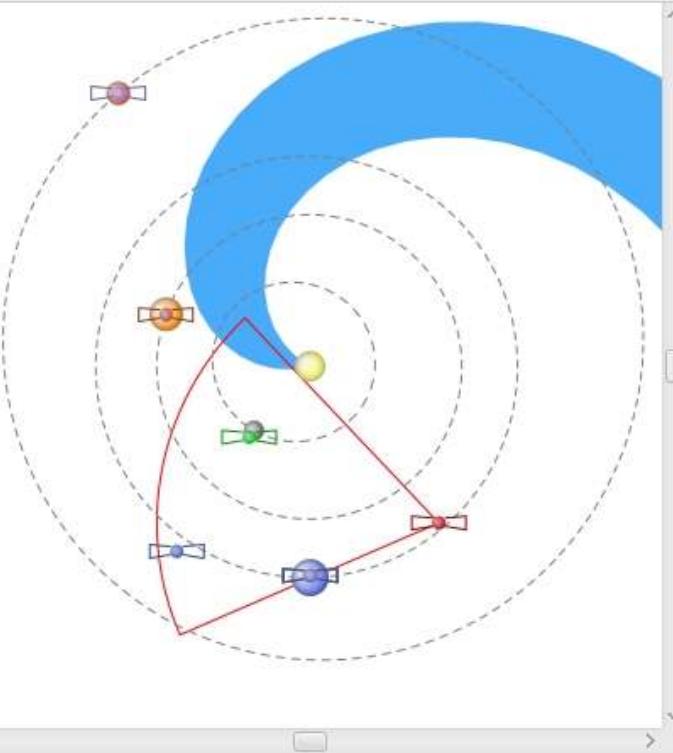


Propagation Tool

Start time

2008-09-24T19:21:14

2008-09-24T19:21:14



Coordinate system : HEE



XZ View

Radial Propagation

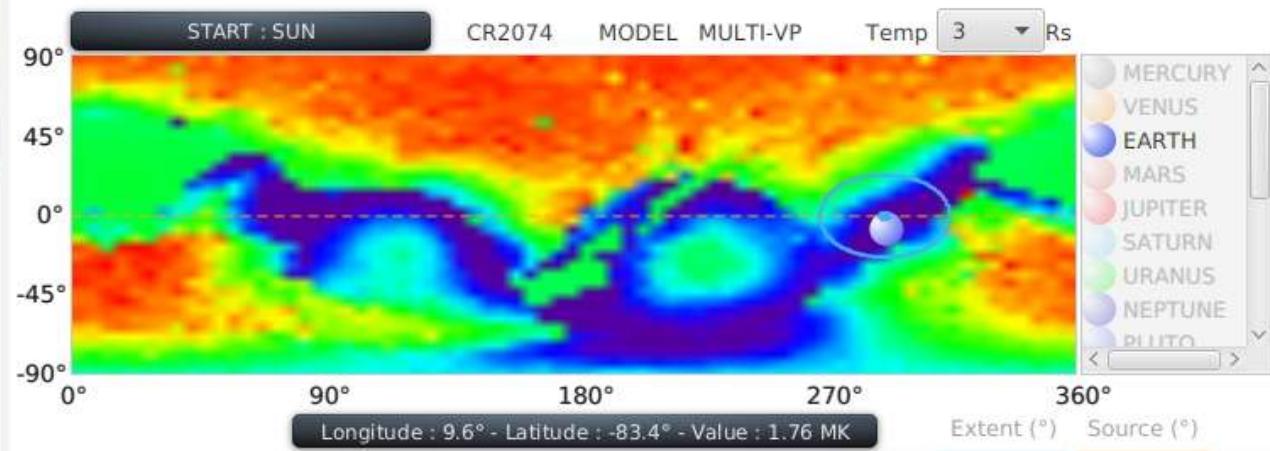
J-map: Carrington/InSitu

Corotation

J-map: Catalogue of fits

SEP Propagation

J-map: Click to fit



Start : SUN

Start Time

2008-09-24T19:21:14

COROTATION

Rotation Period
(days)

End : EARTH

2008-10-07T13:22:00

CDPP Interface

Date	2008-09-24T19:21:14	Center on start time
Time Interval	4	days
Observatory	OMNI	
Parameters	B, Bx, By, Bz, N, T, V	

AMDA plot

Corotaton Interface

J-Map Interface

Table of Arrival Times

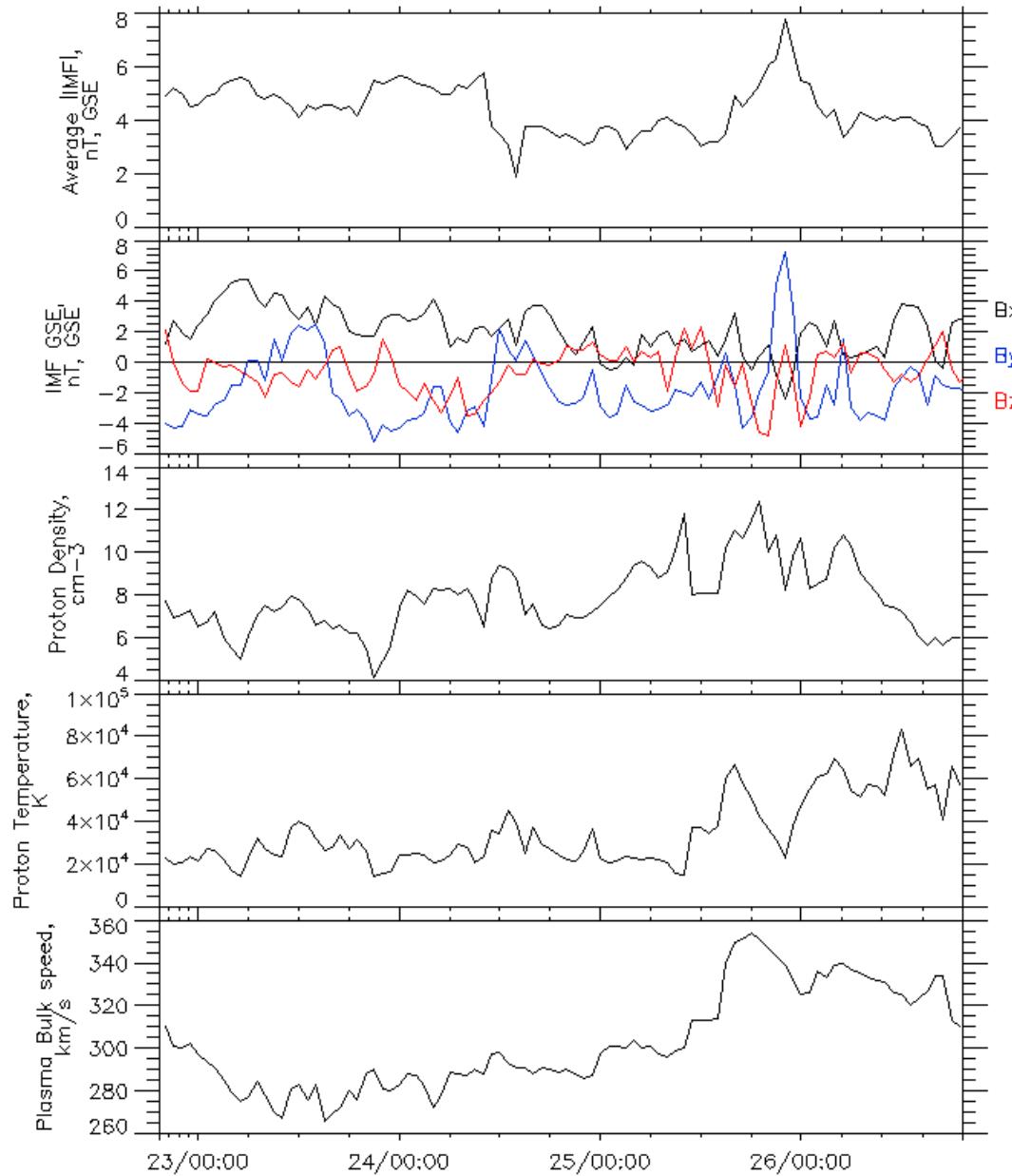
Arrival Times Catalogue

Heliographer

CDPP Interface

APIS Interface

SAMP Client Monitor



Sep 2008

Created by AMDA(C) V2.0 Wed Sep 26 00:16:59 2018