



KARL-FRANZENS-UNIVERSITÄT GRAZ  
UNIVERSITY OF GRAZ



# Evolution of plasma parameters during the early acceleration phase of the June 13 2010 CME event

J. Seibezeder<sup>1</sup>, A. Veronig<sup>1,2</sup>, K. Dissauer<sup>1</sup>, M. Temmer<sup>1</sup>, K. Vanninathan<sup>1</sup>, B. Vršnak<sup>3</sup>

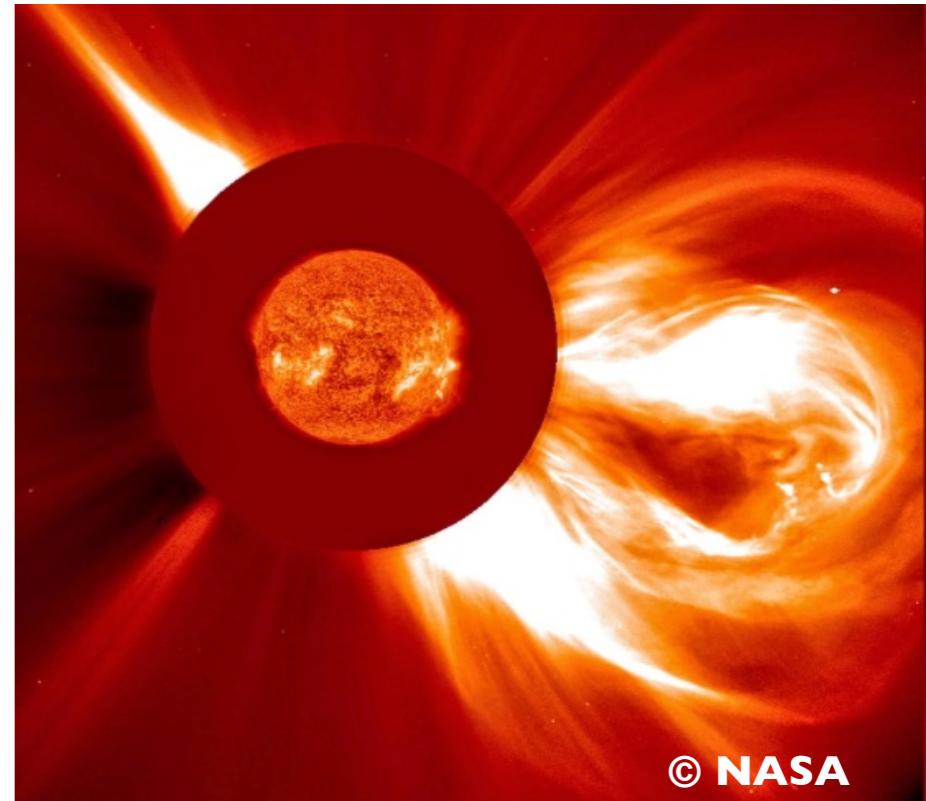
<sup>1</sup> Institute of Physics, University of Graz, Austria

<sup>2</sup> Kanzelhöhe Observatory, University of Graz, Austria

<sup>3</sup> Hvar Observatory, Faculty of Geodesy, University of Zagreb, Croatia

# Coronal Mass Ejections

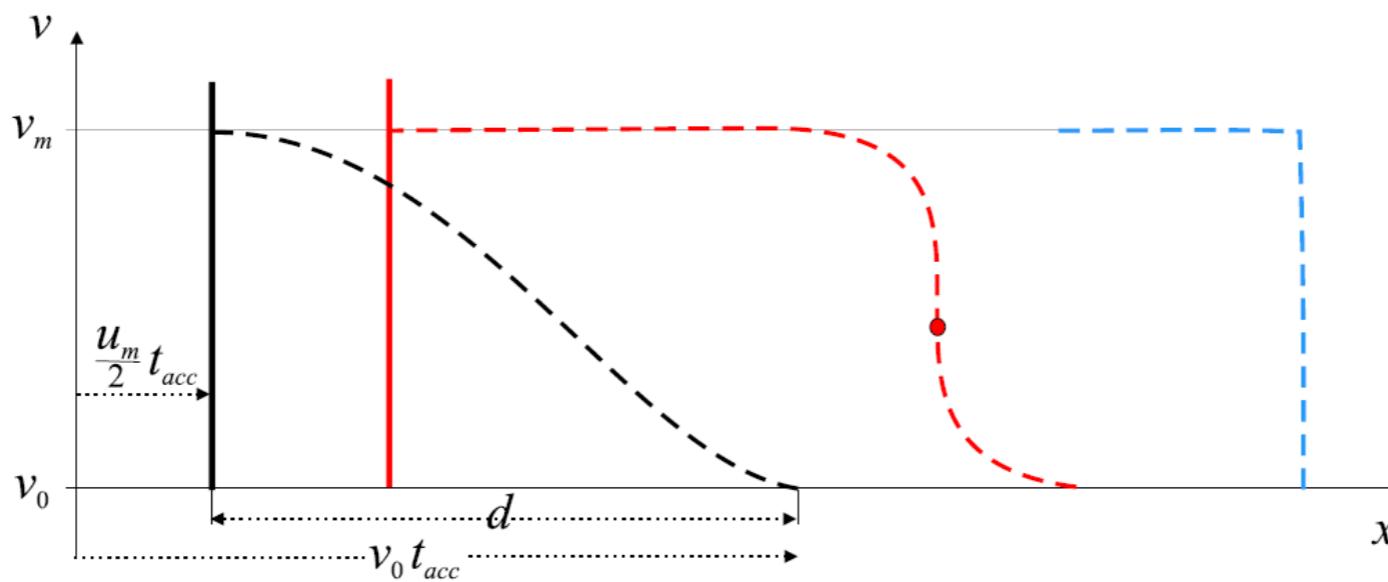
- Structures containing plasma and magnetic field expelled from Solar Corona
- Significant influence on IP space, space weather
- Earth's magnetosphere  
→ geomagnetic storms<sup>[1,2]</sup>



**Characterizing their evolution  
helps us understand their origins**

# CME Shocks

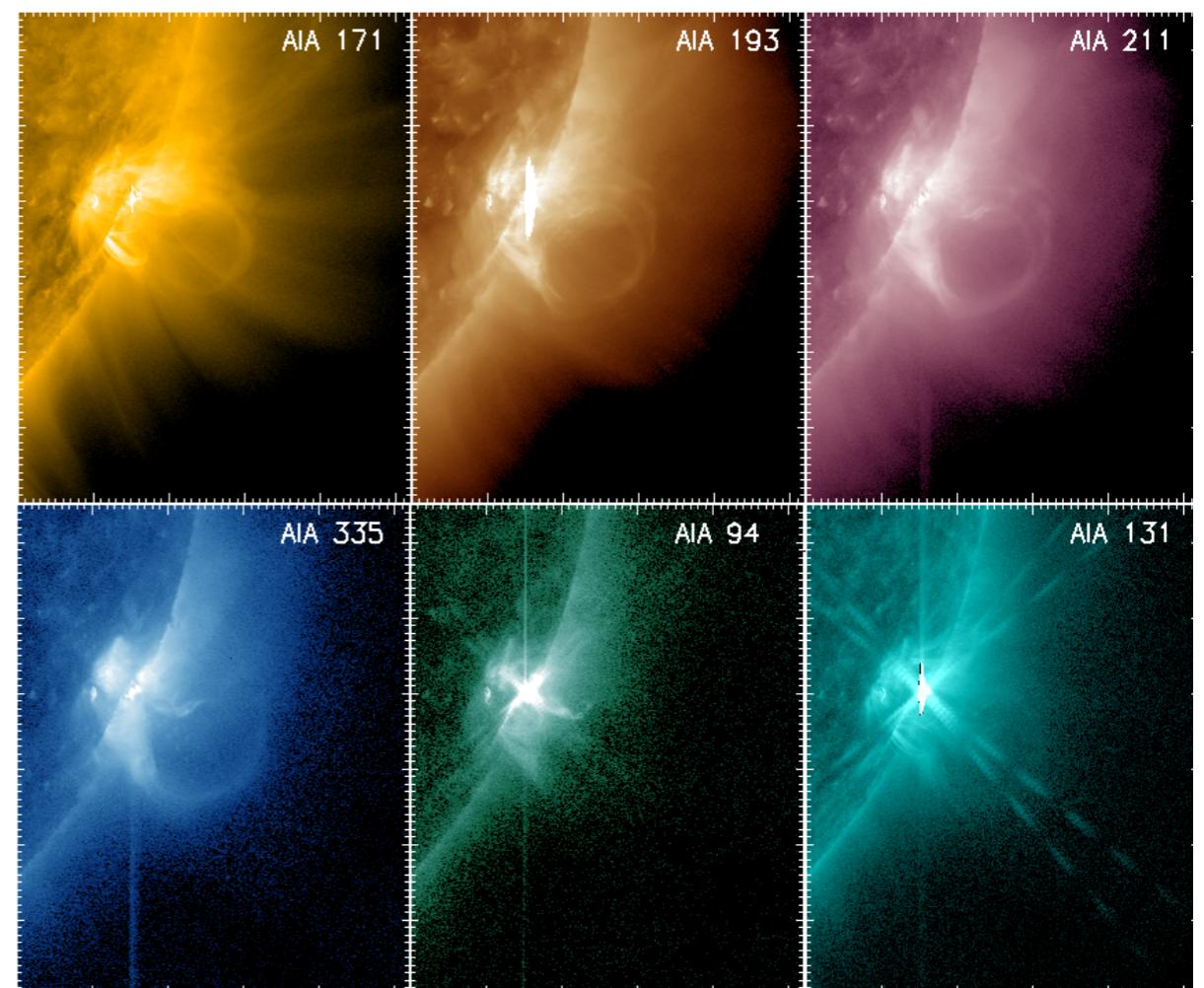
- Explosive expansion due to CMEs → acts as a 3D piston (Lulic+, 2013)
  - If impulsive enough:
    - Perturbation is created
    - Nonlinear evolution of the wave front → perturbation steepens  
→ transformation into shock wave (Vršnak+, 2008)



**B. Vršnak, E.W. Cliver,  
2008**

# The event: 13 June 2010

- Well-studied event
- M1.0 flare, AR 11079
- Strong, short-lived acceleration  
(Patsourakos+, 2010)
- Type II radio burst (Ma+, 2011;  
Kozarev+, 2011)  
→ indicates coronal shocks
- Very slow off-limb event



# Method

- Data: SDO/AIA spacecraft
- Channels: 6 EUV (94 Å, 131 Å, 171 Å, 194 Å, 211 Å, 335 Å)
- Differential Emission Measure (DEM) technique  
by Hannah & Kontar, 2012

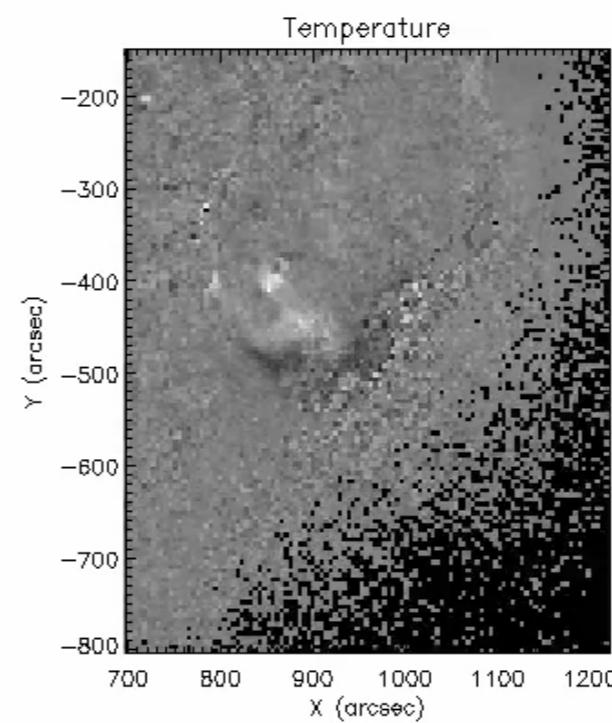
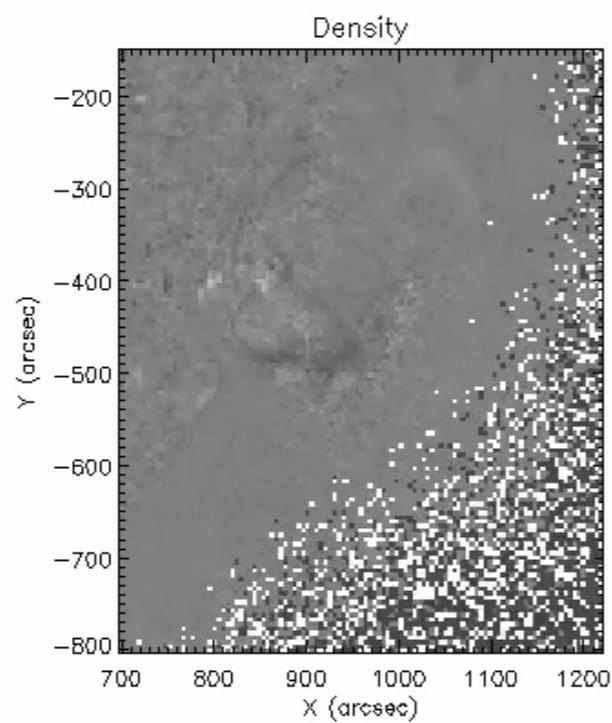
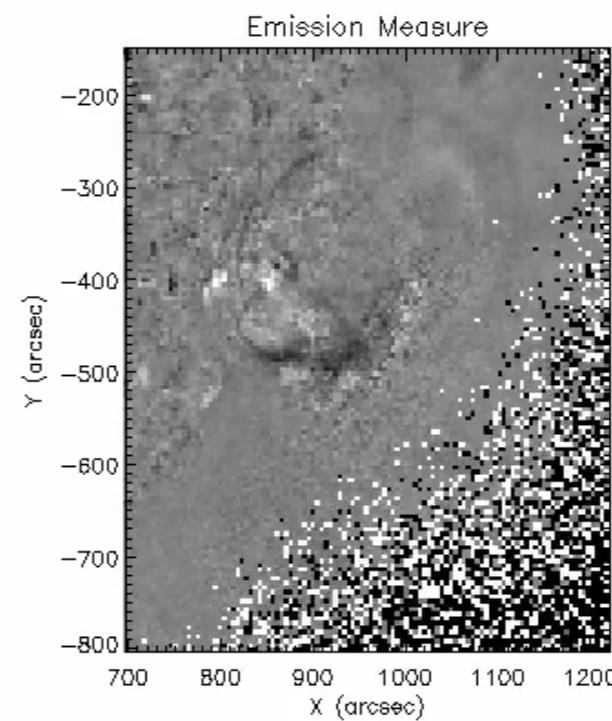
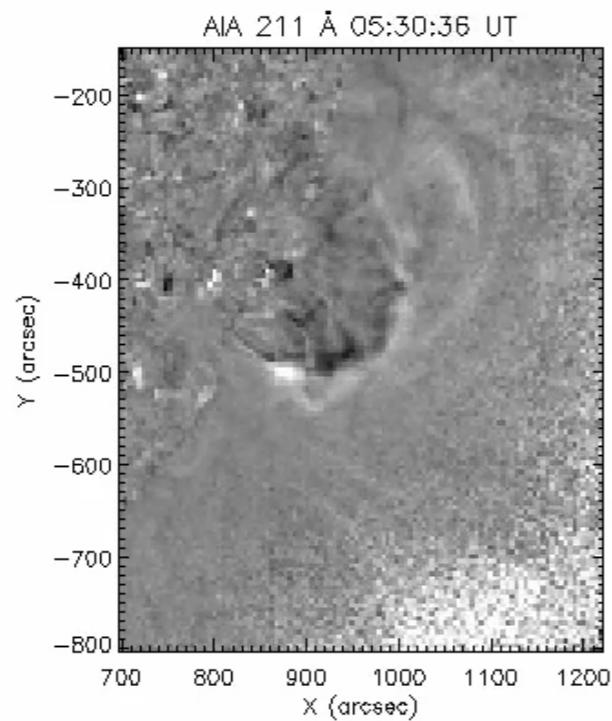
→ Plasma temperature:  $\bar{T} = \frac{\int DEM(T) T dT}{\int DEM(T) dT}$

→ Emission Measure:  $EM = \int DEM(T) dT$

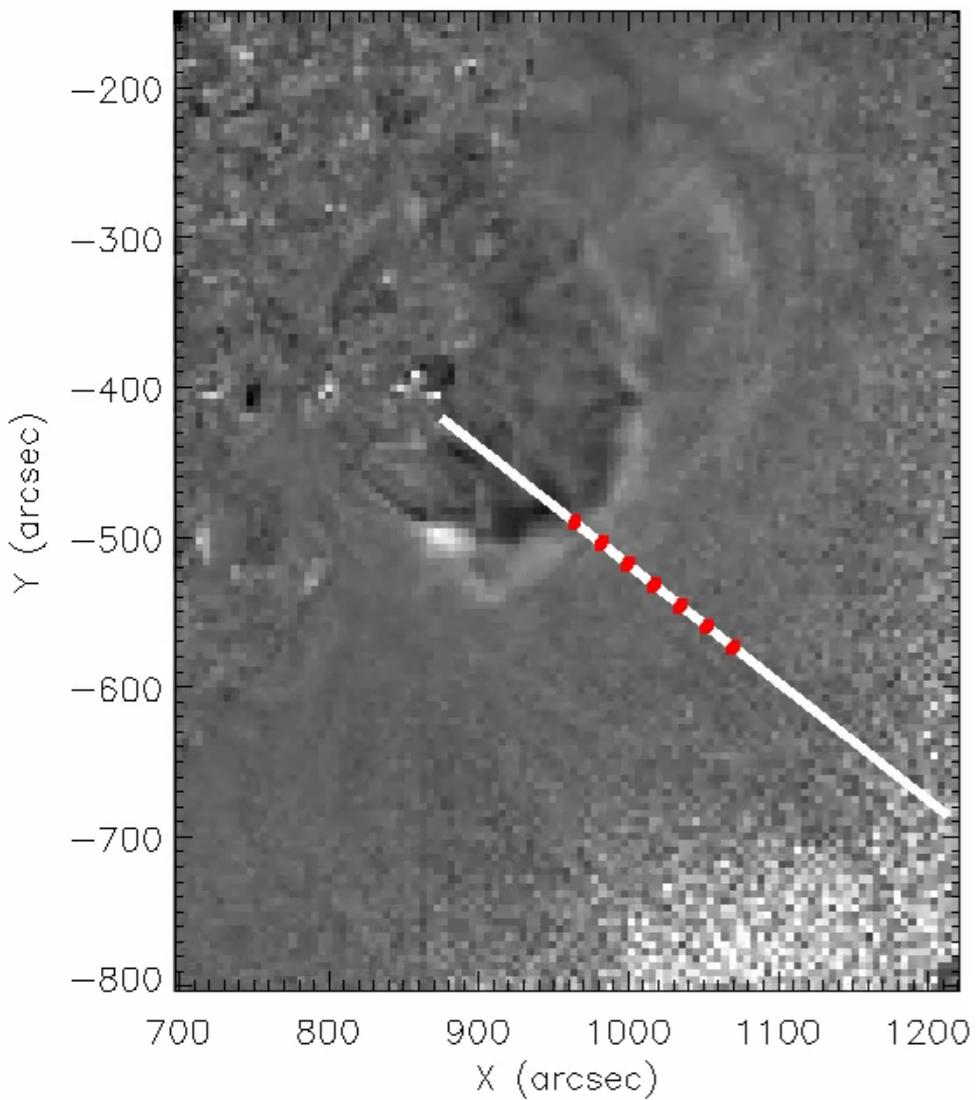
→ Plasma Density:  $\bar{n} = \sqrt{\frac{\int DEM(T) dT}{H}} = \sqrt{\frac{EM}{H}}$

# Calculated DEM maps

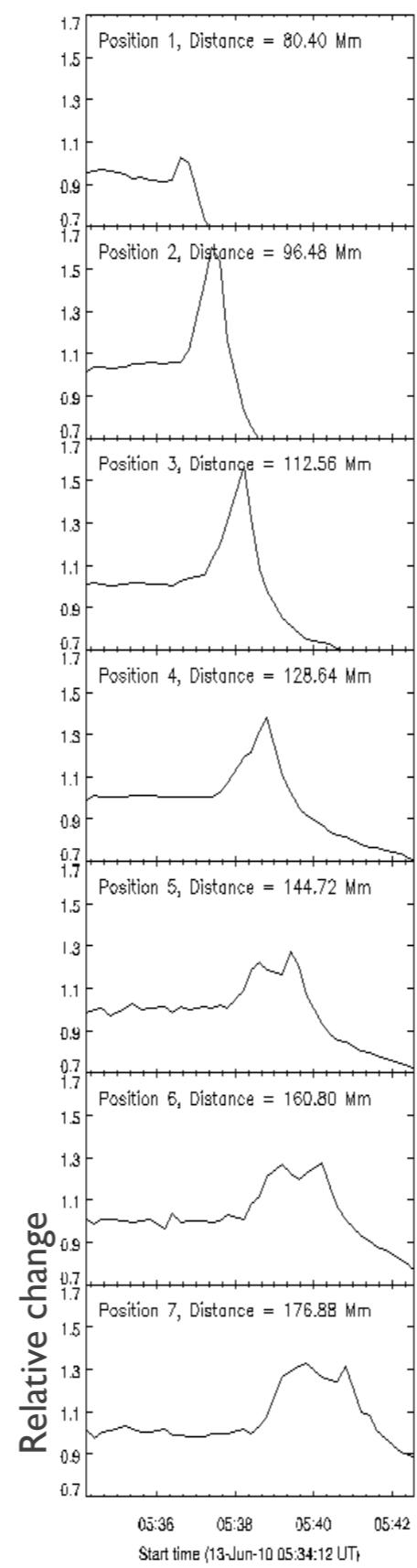
13/06/2010



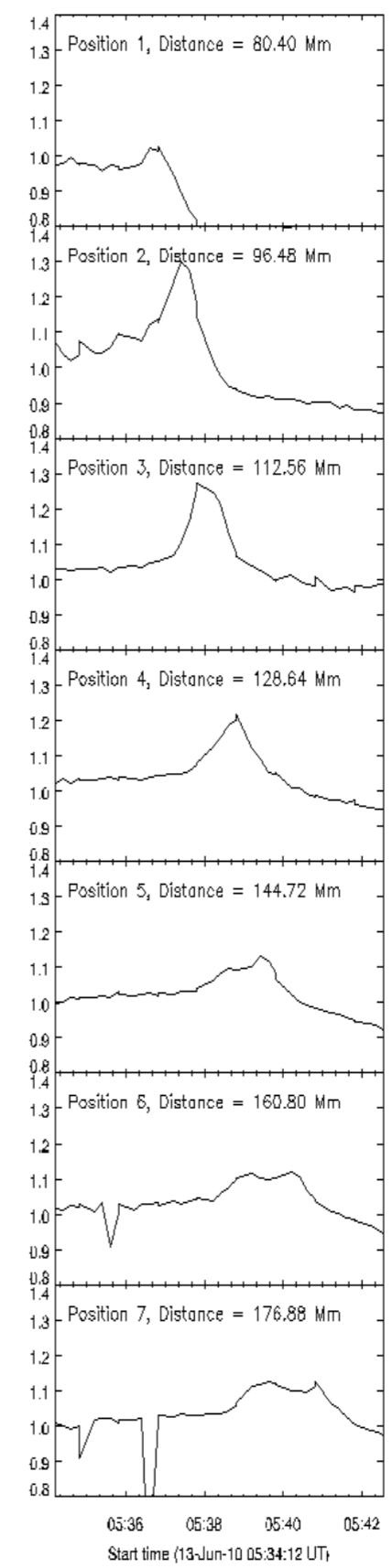
SDO AIA\_2 211 13-Jun-2010 05:32:48.620 UT



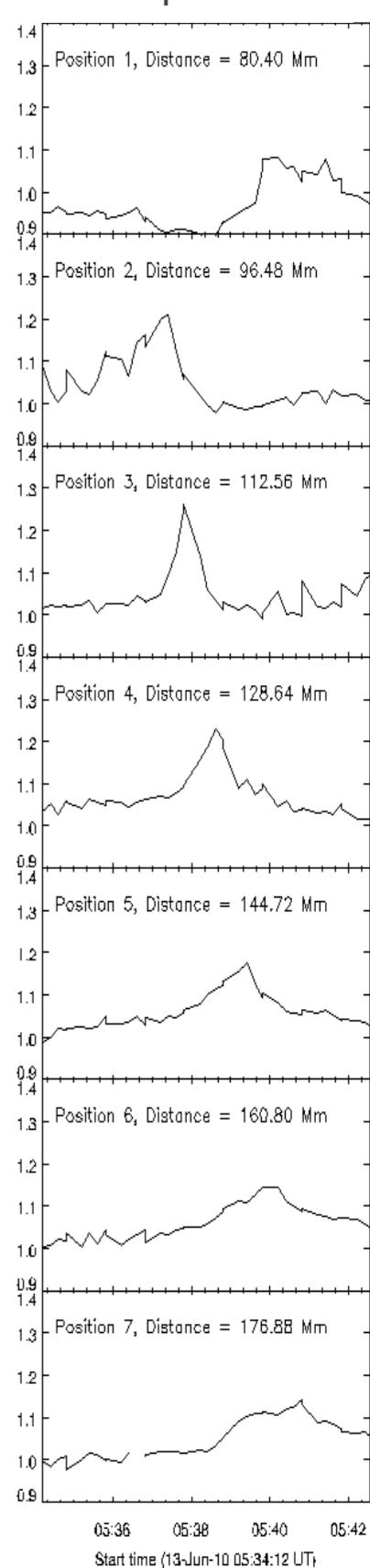
211 Å



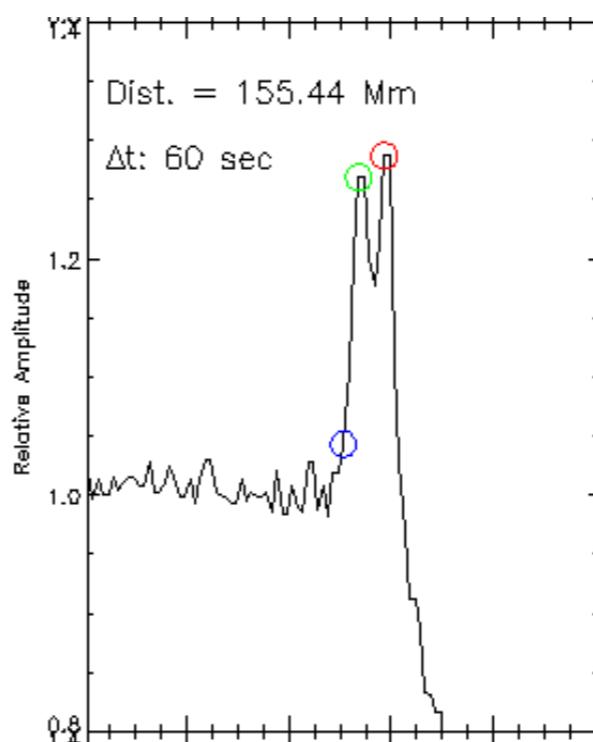
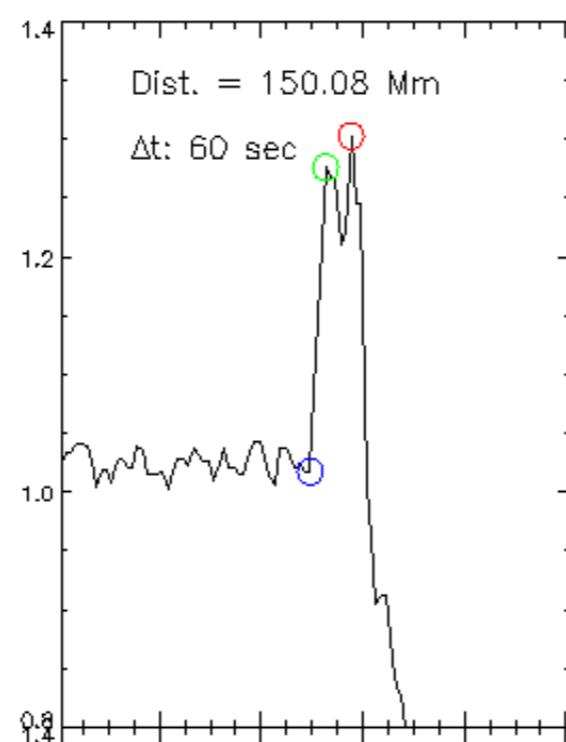
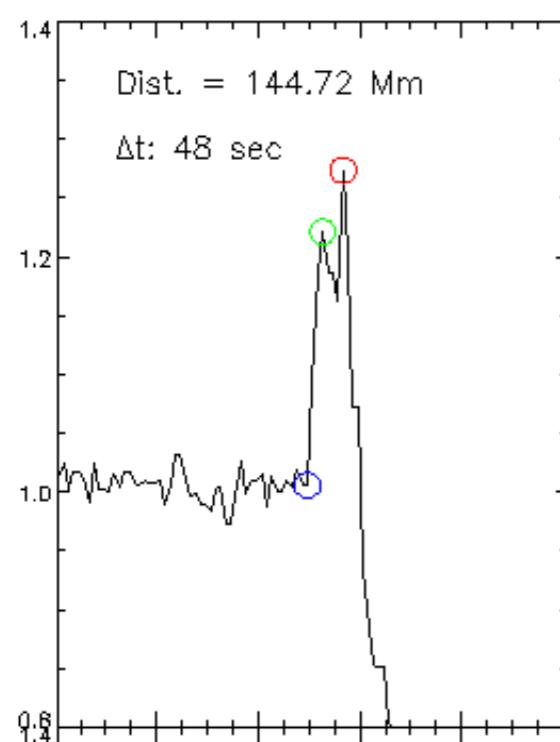
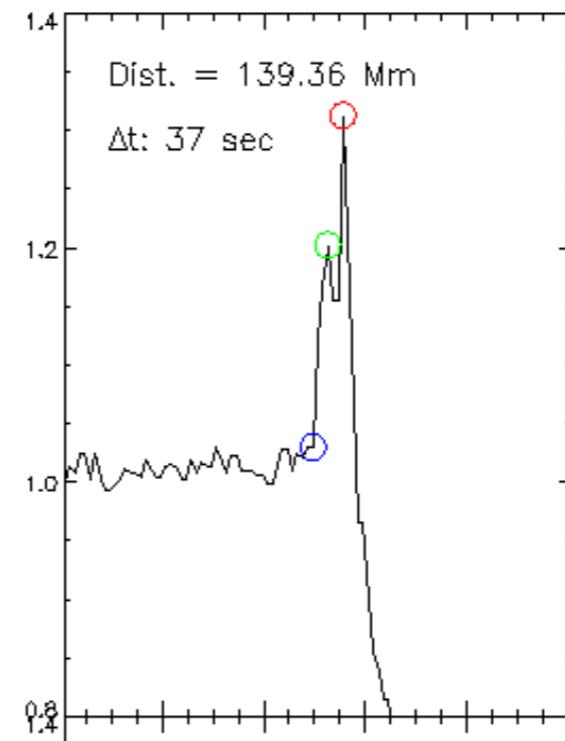
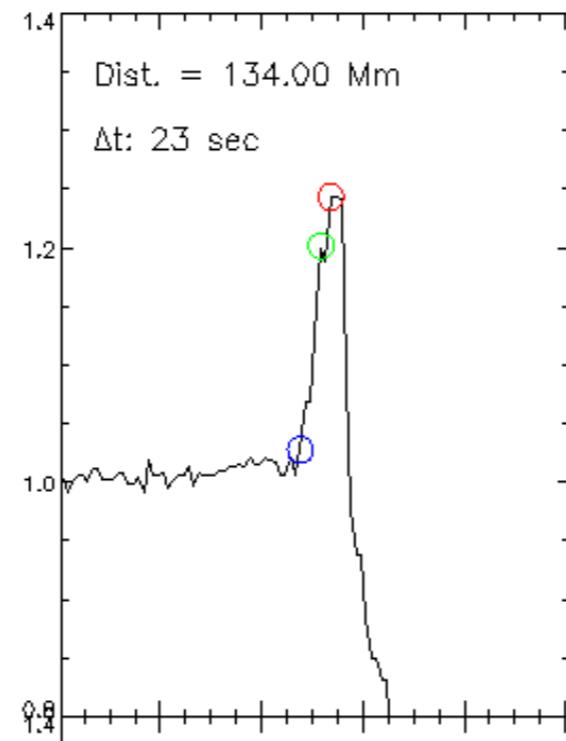
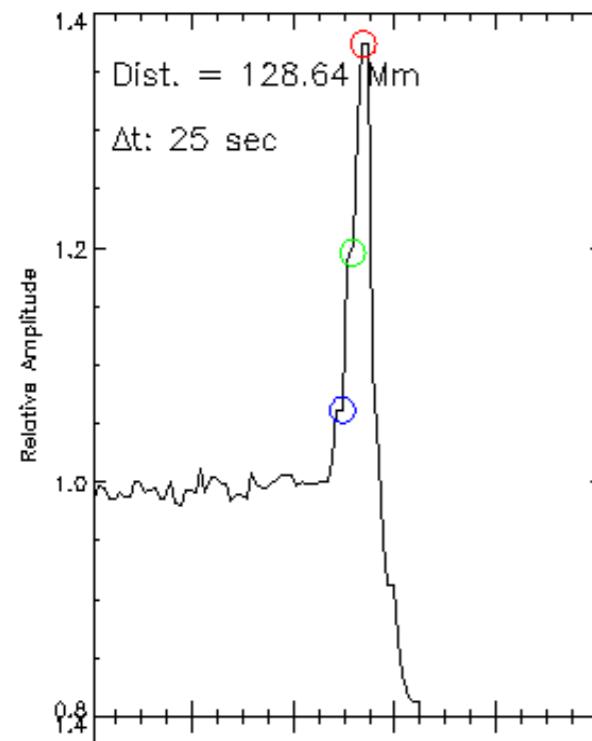
Density



Temperature



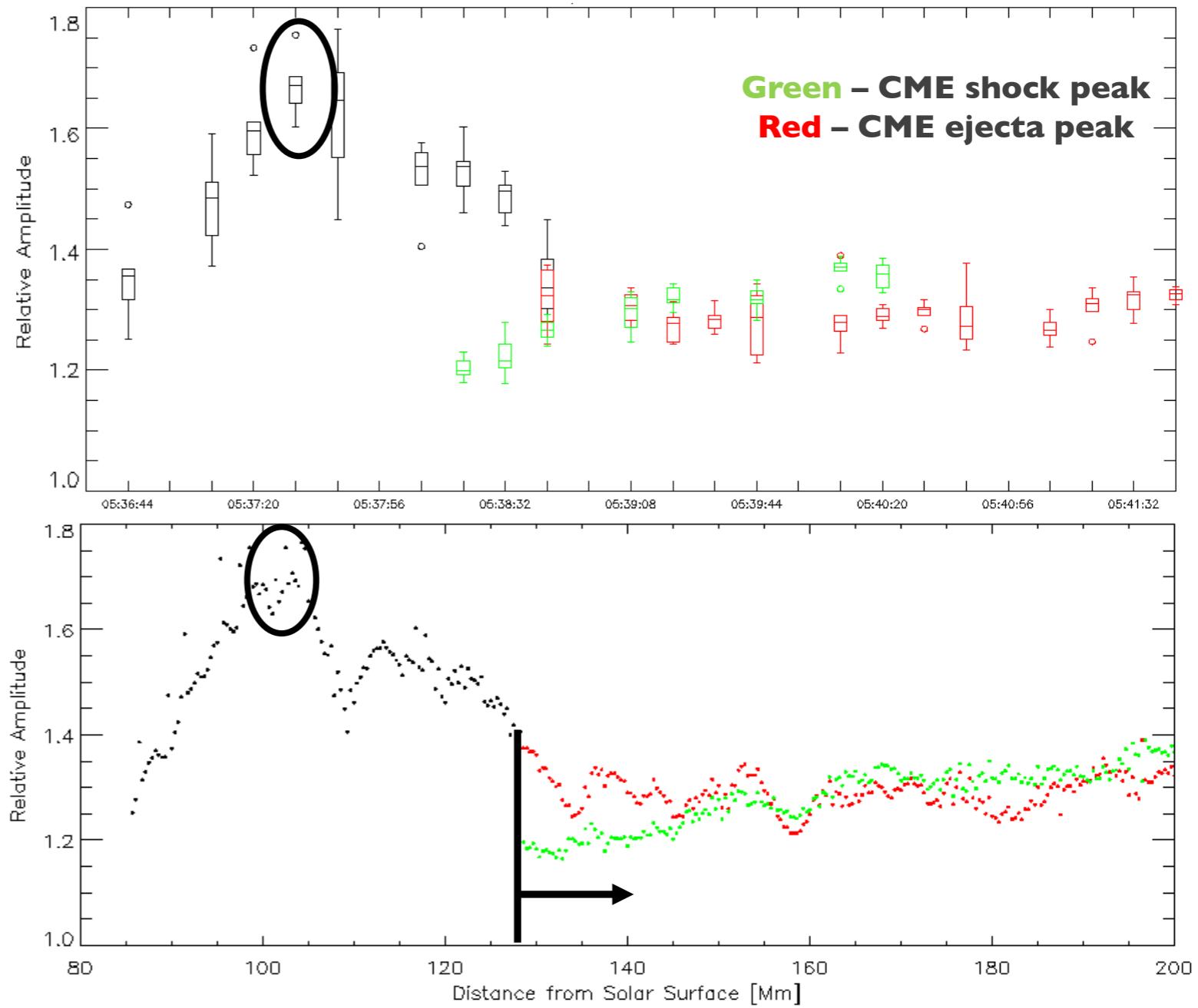
# Signature in 211 Å



Blue – CME shock (outer front)  
Green – CME shock (peak)  
Red – CME ejecta (peak)

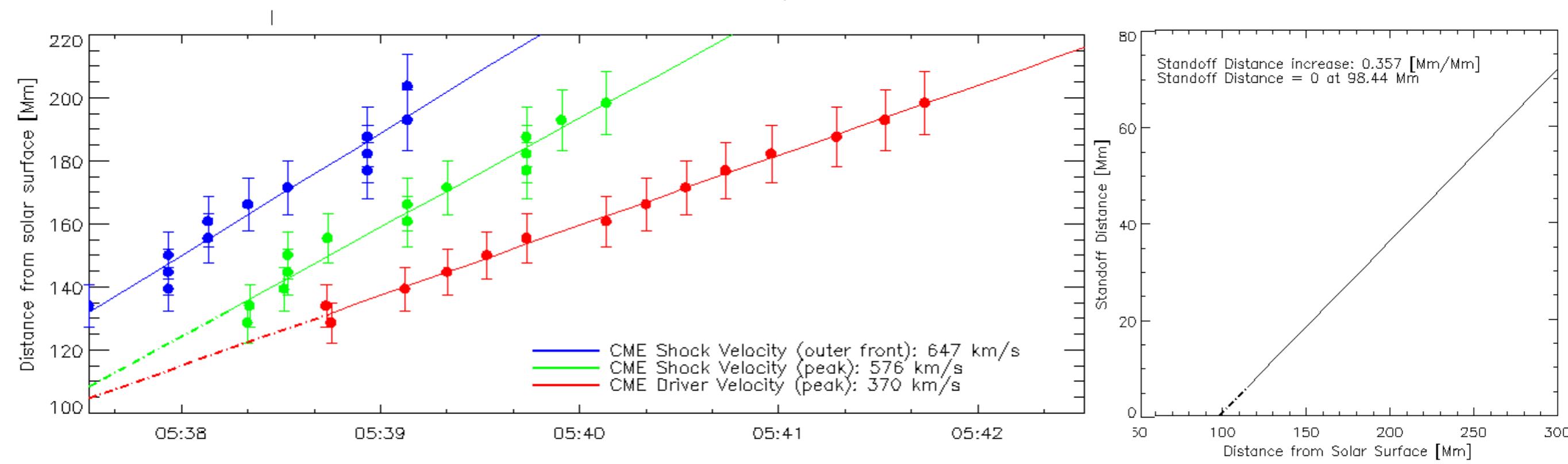
# Signature in 211 Å

- Maximum amplitude increase: 60%  
at 105 Mm  
05:37:32
- 128 Mm onwards: clear double peak structure  
→ de-coupling

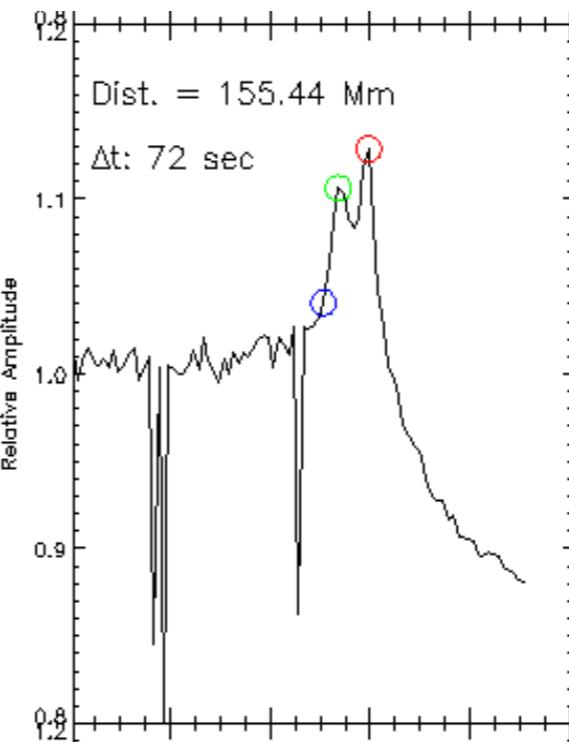
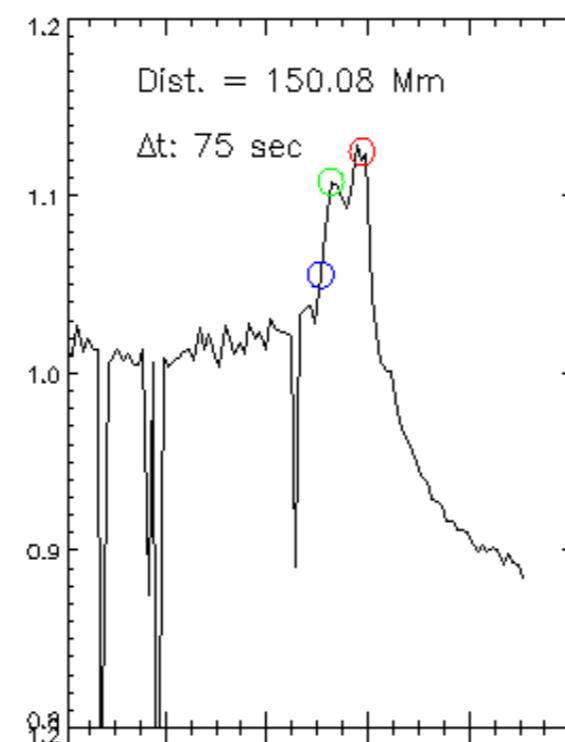
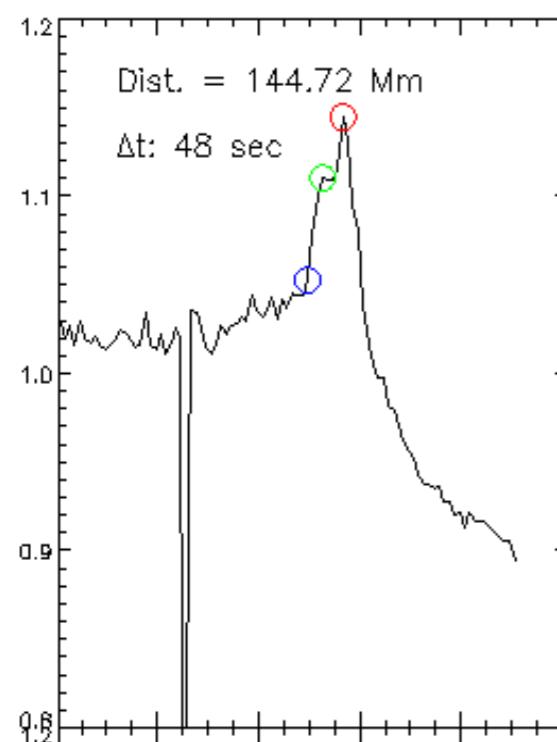
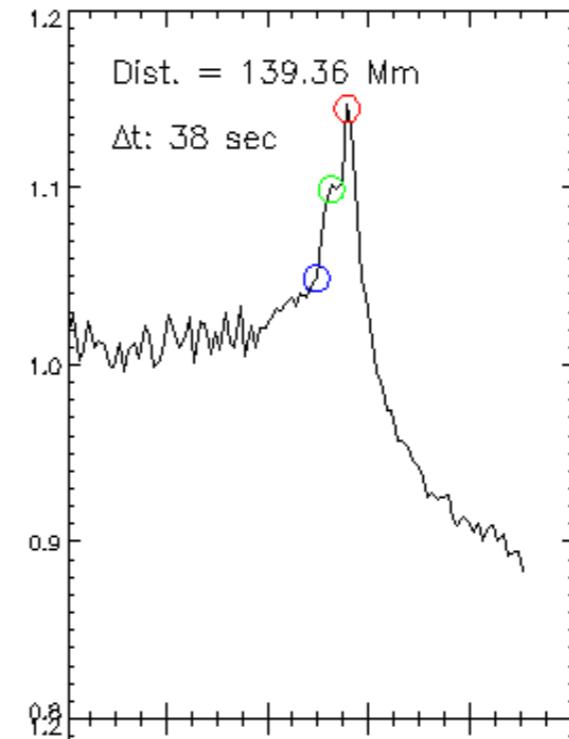
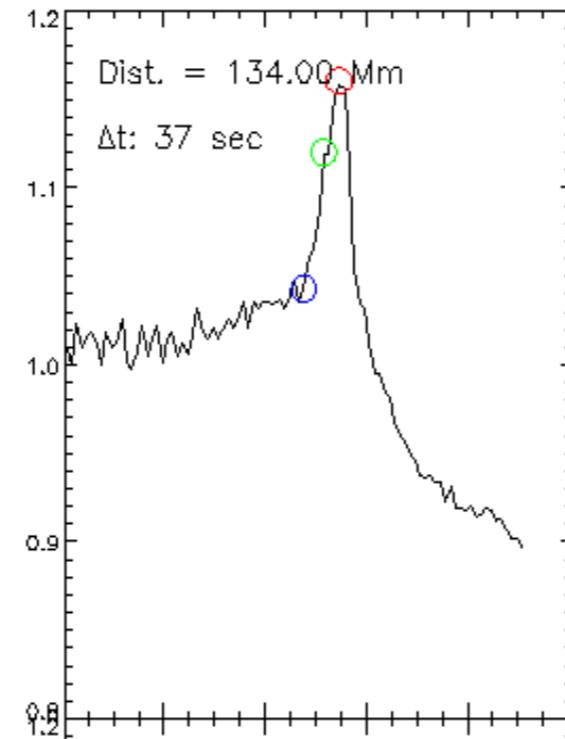
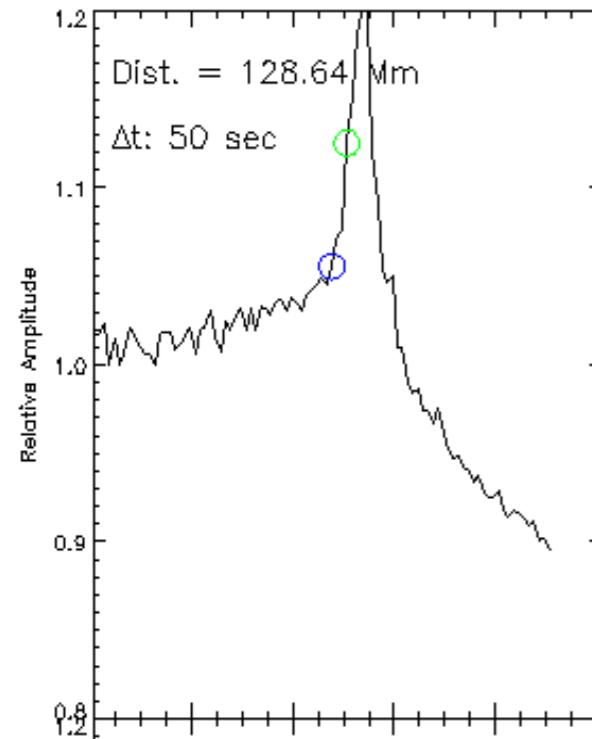


# Signature in 211 Å

- CME ejecta velocity: 370 km/s
- CME shock velocity: 576 km/s
- Shock formation: 98 Mm from solar surface
- Standoff distance increase: 206 km/s ( $\approx 0.36 \text{ Mm/Mm}$ )

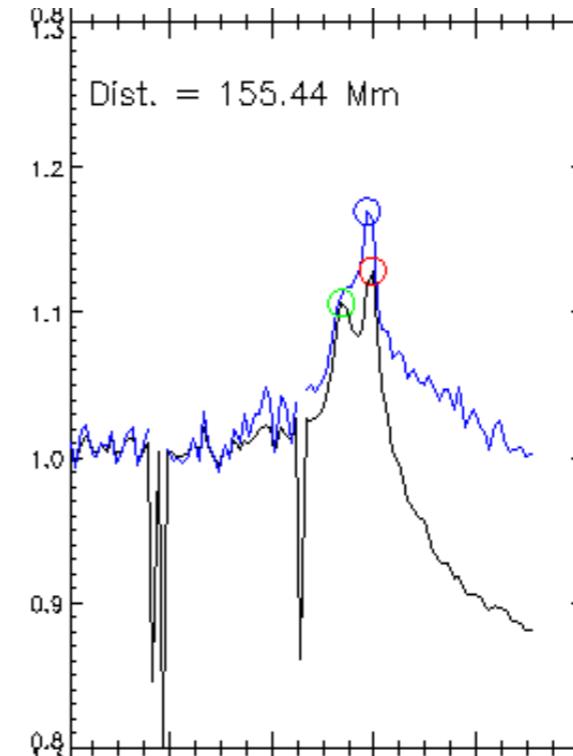
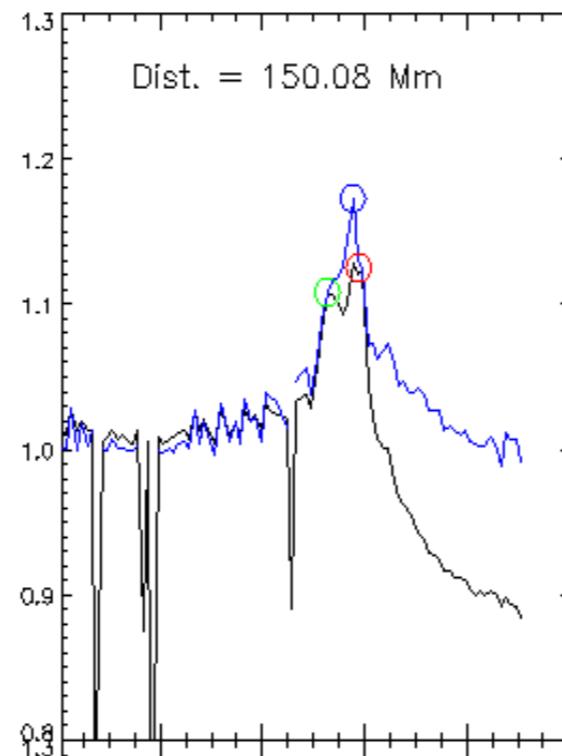
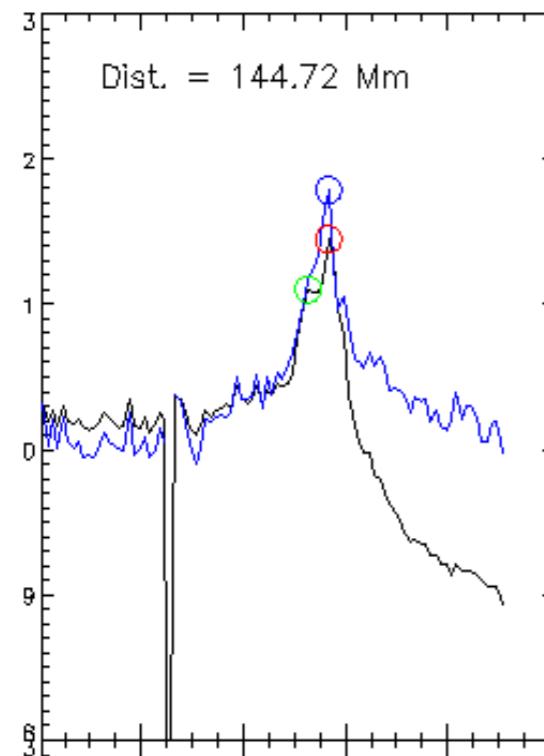
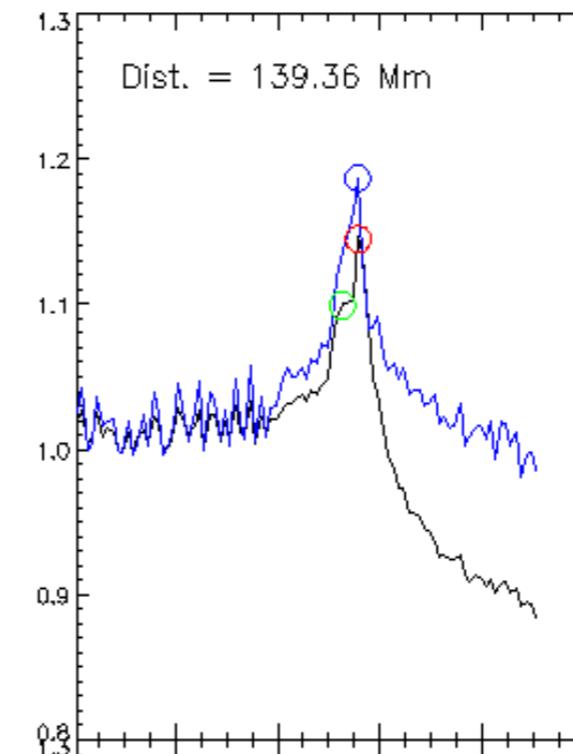
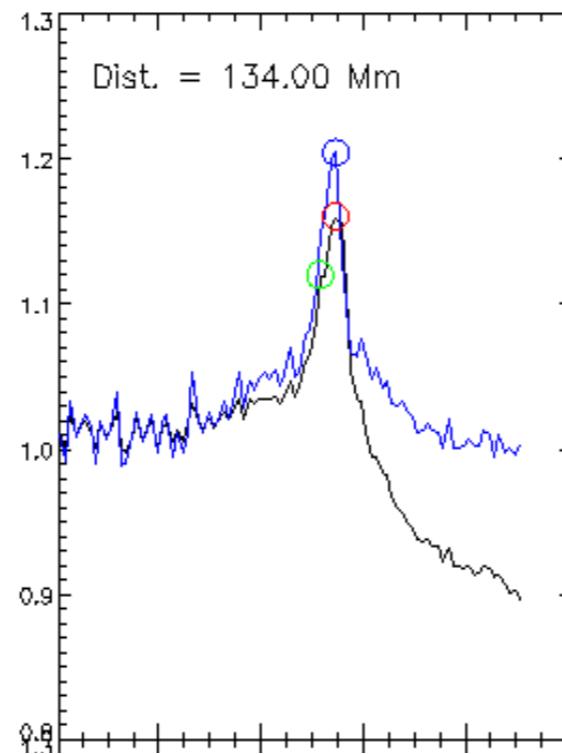
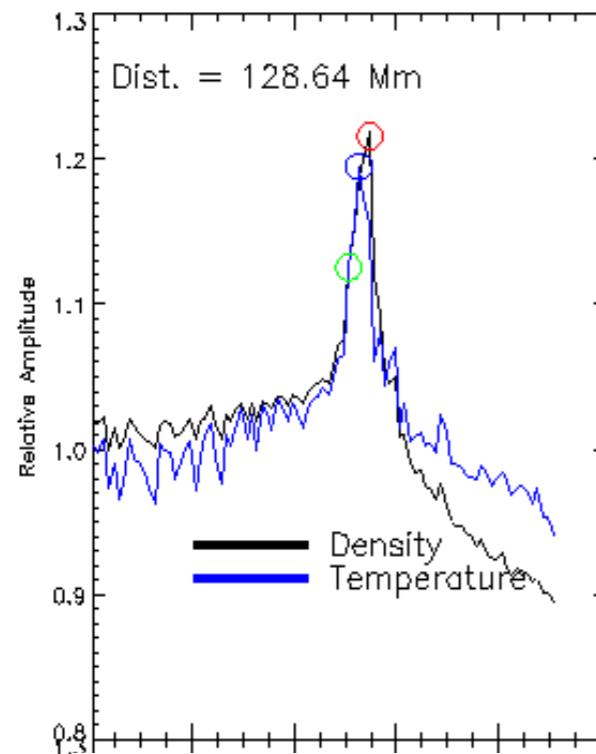


# Density



Green – CME shock (peak)  
Red – CME ejecta (peak)  
Blue – CME shock (outer front)

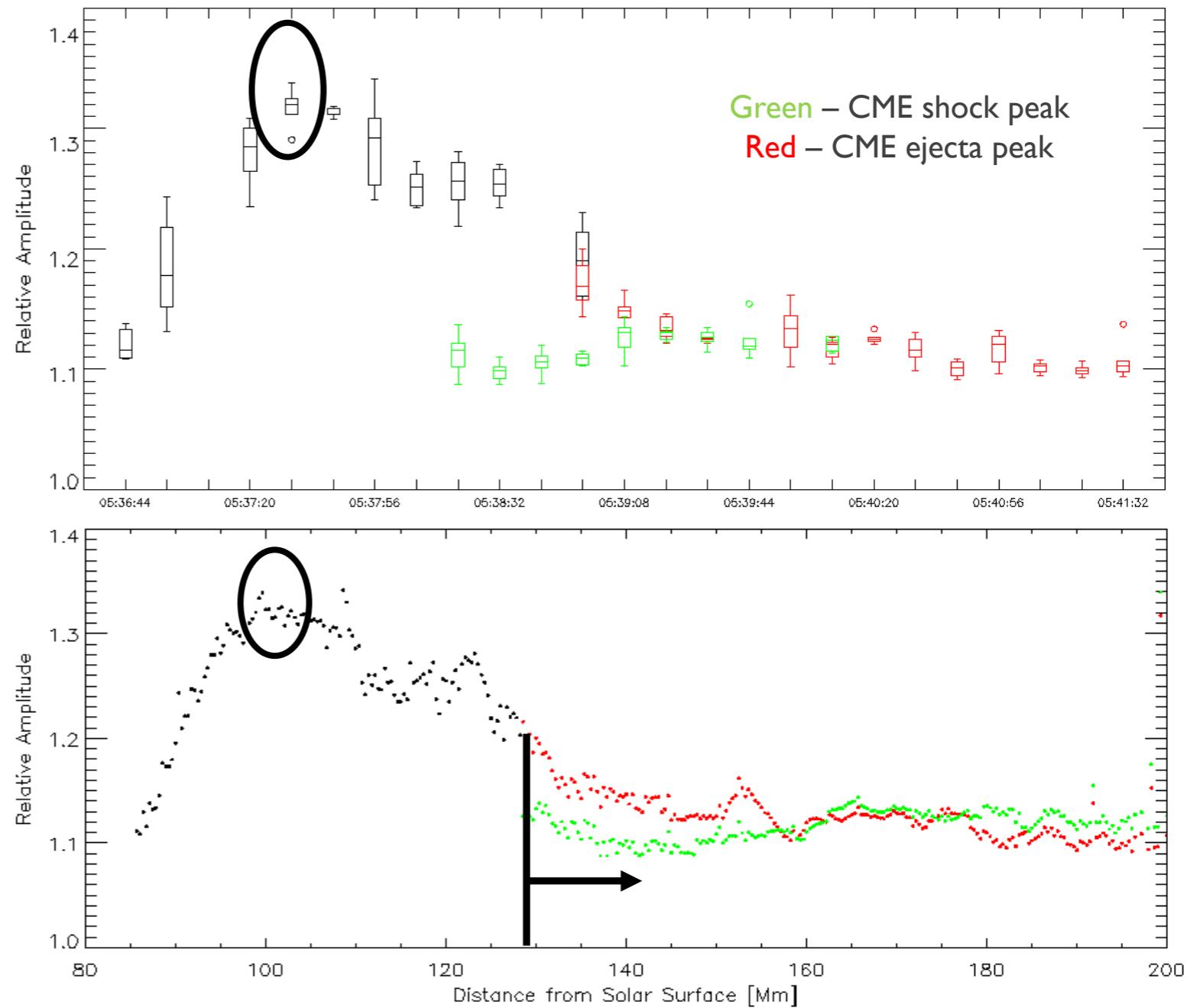
# Density vs. Temperature



- Green** – CME shock,  
Density (peak)
- Red** – CME ejecta,  
Density (peak)
- Blue** – CME ejecta,  
Temperature (peak)

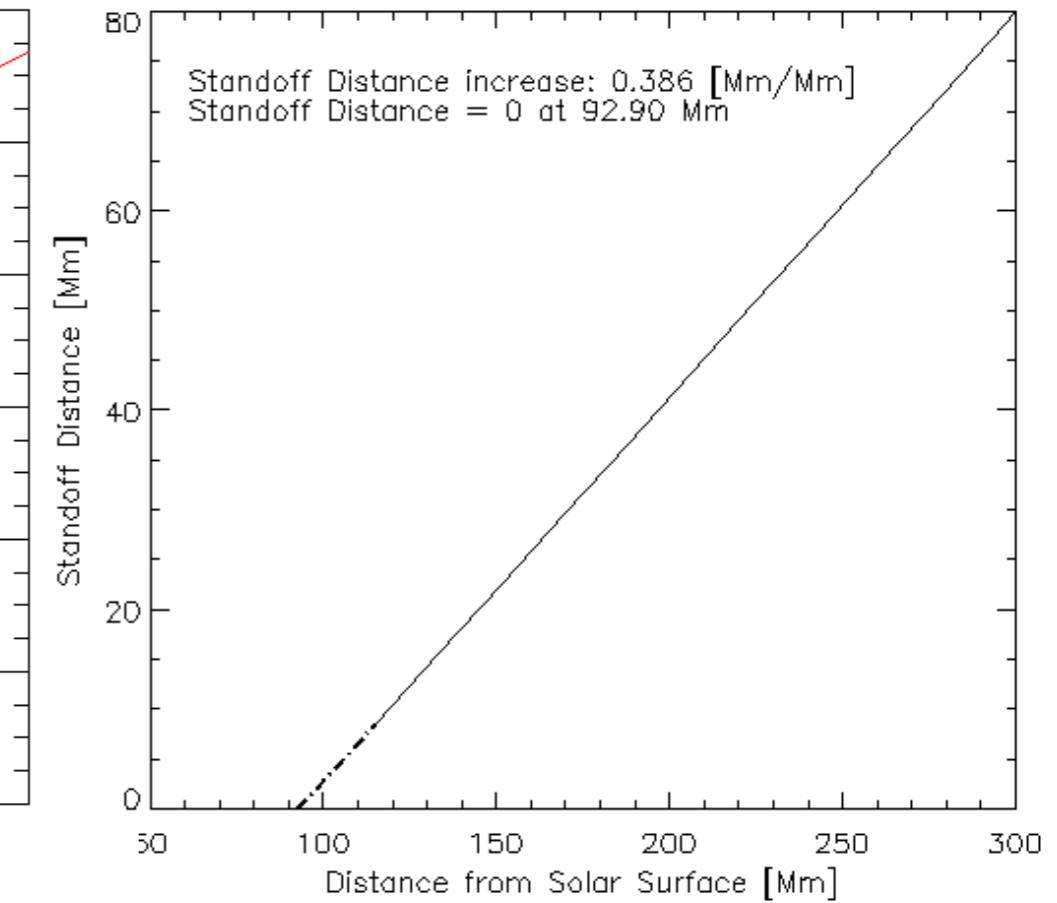
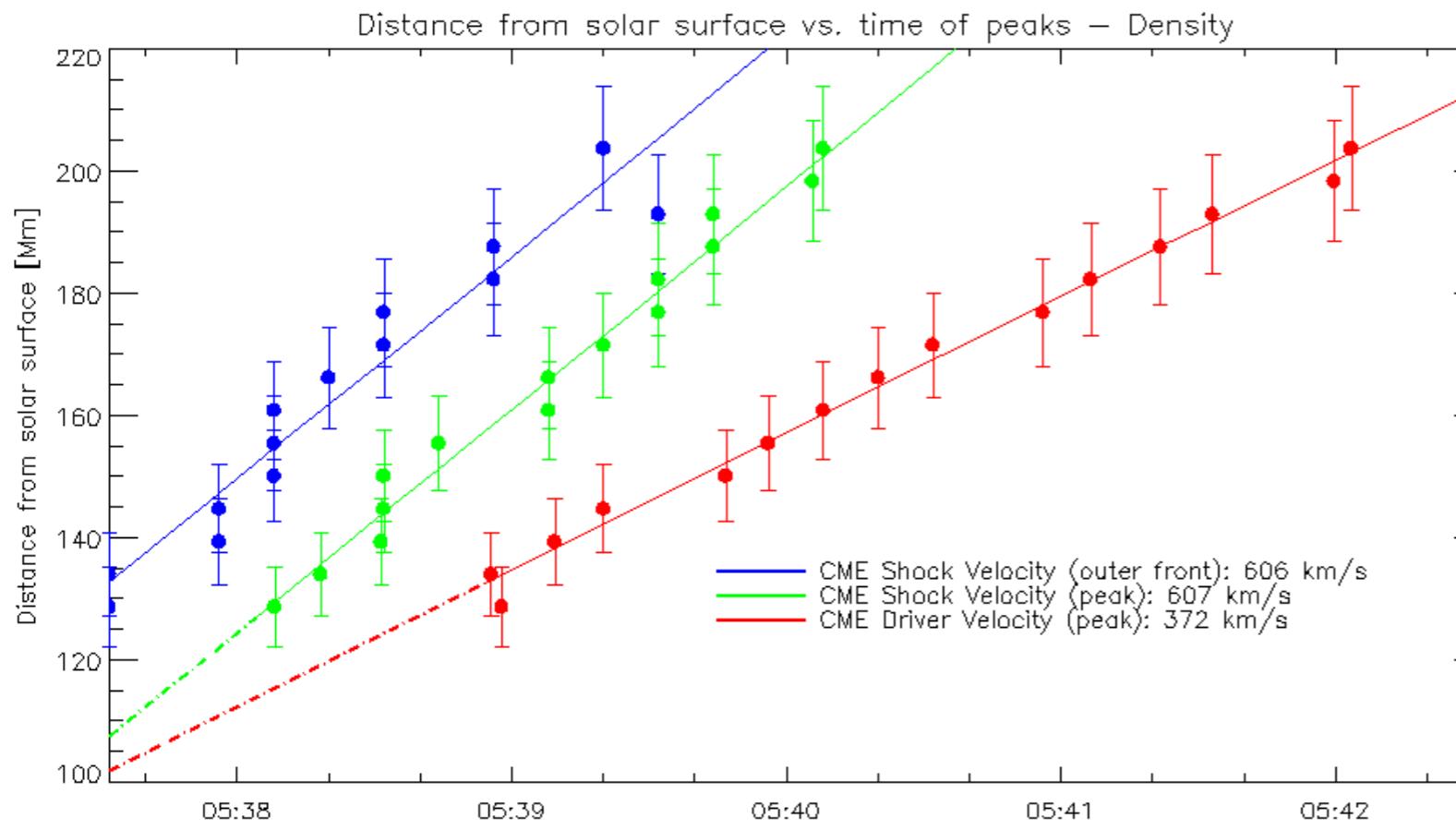
# Density

- Maximum amplitude increase: 30%  
at 103 Mm  
05:37:32
- 128 Mm onwards: clear double peak structure  
→ de-coupling



# Density

- CME ejecta velocity: 372 km/s
- CME shock velocity: 607 km/s
- Shock formation: 93 Mm from solar surface
- Standoff distance increase: 234 km/s  
 $(\approx 0.39 \text{ Mm/Mm})$



# Summary

- CME shock velocity:  $\sim 600$  km/s
- CME ejecta velocity:  $\sim 380$  km/s
- Shock formation starts  $\sim 95$  Mm from solar surface
- Maximum compression ratio of 1.3 at  $\sim 105$  Mm from solar surface
- De-coupling low in the corona,  $\sim 128$  Mm from solar surface
- Indication of piston-driven shock due to linear increase of standoff-distance



KARL-FRANZENS-UNIVERSITÄT GRAZ  
UNIVERSITY OF GRAZ



# THANK YOU FOR YOUR ATTENTION