

Parametric functions of the coronal mass ejections properties close to the Sun based on numerical hydrodynamic simulations

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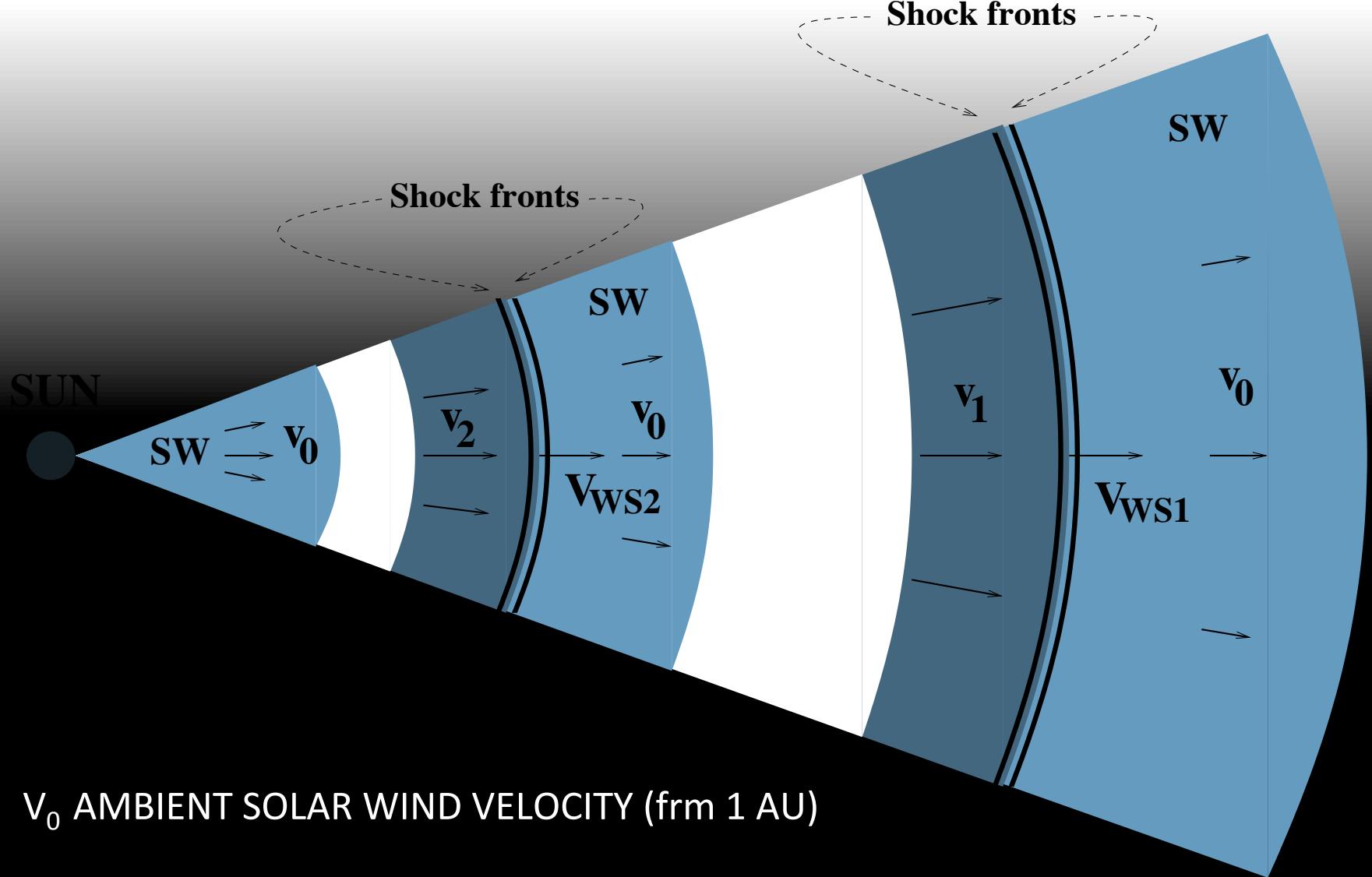
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v_0 AMBIENT SOLAR WIND VELOCITY (frm 1 AU)

v_1 VELOCITY OF THE 1st CME (LEADING EDGE)

v_2 VELOCITY OF THE 1st CME (LEADING EDGE)

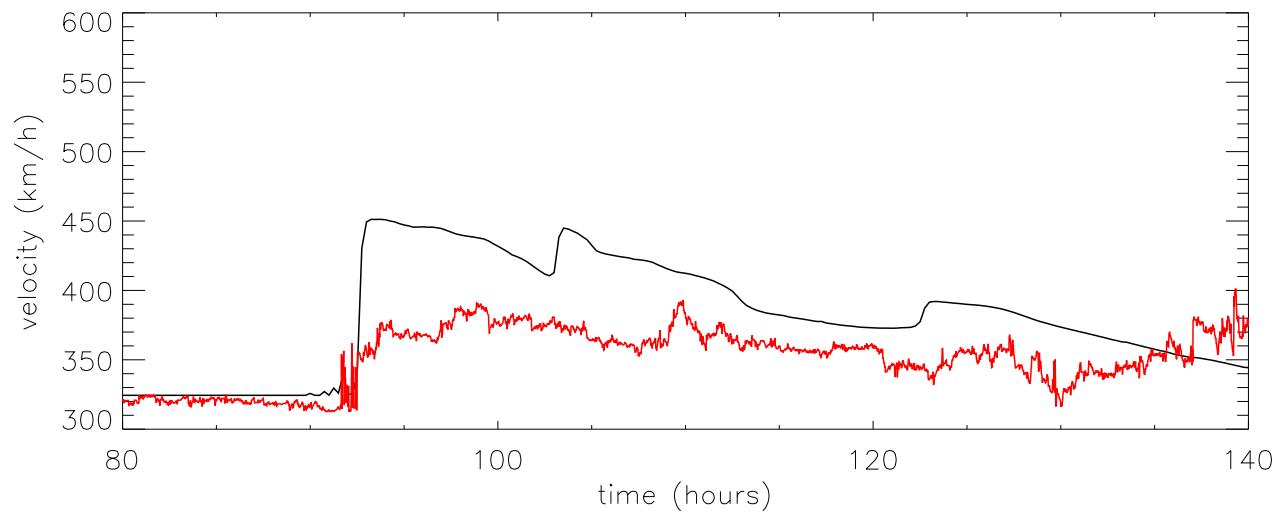
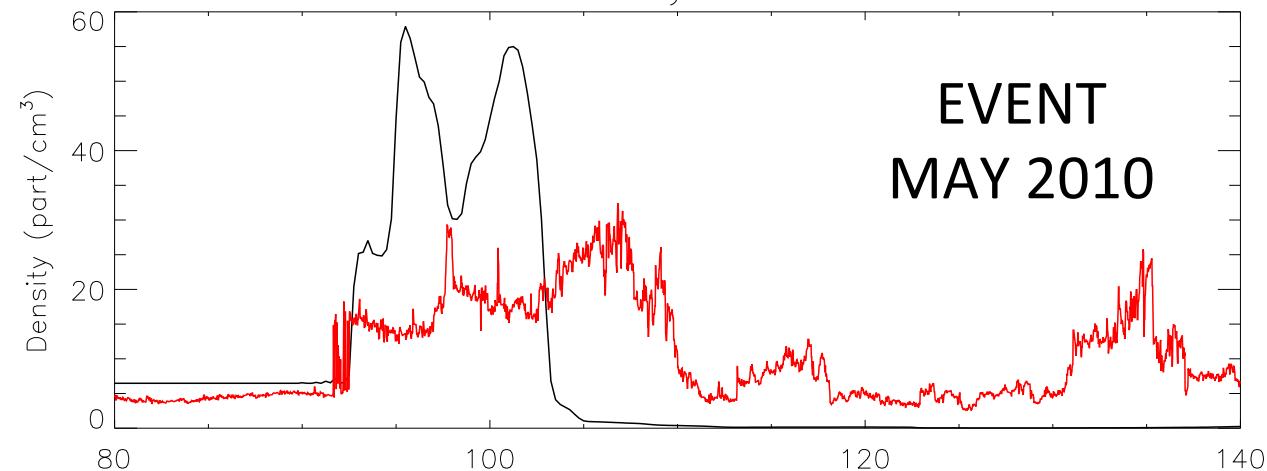
Niembro et al 2015

METHOD

- 1) We set the initial parameters for:
CME: velocity, mass loss rate, injection time
SW: velocity, mass loss rate
- 2) We obtained profiles of density and velocity as function of time at 1 AU
- 3) We compared our results with real data at 1 AU

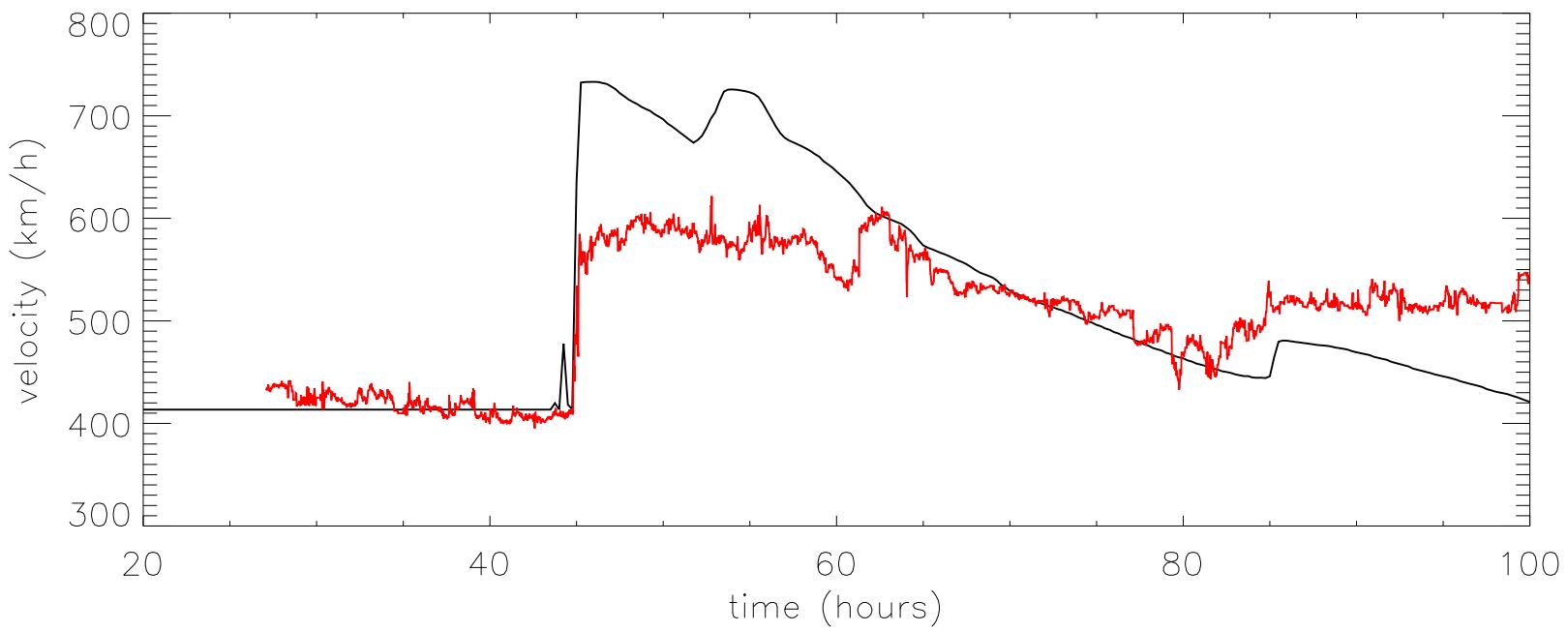
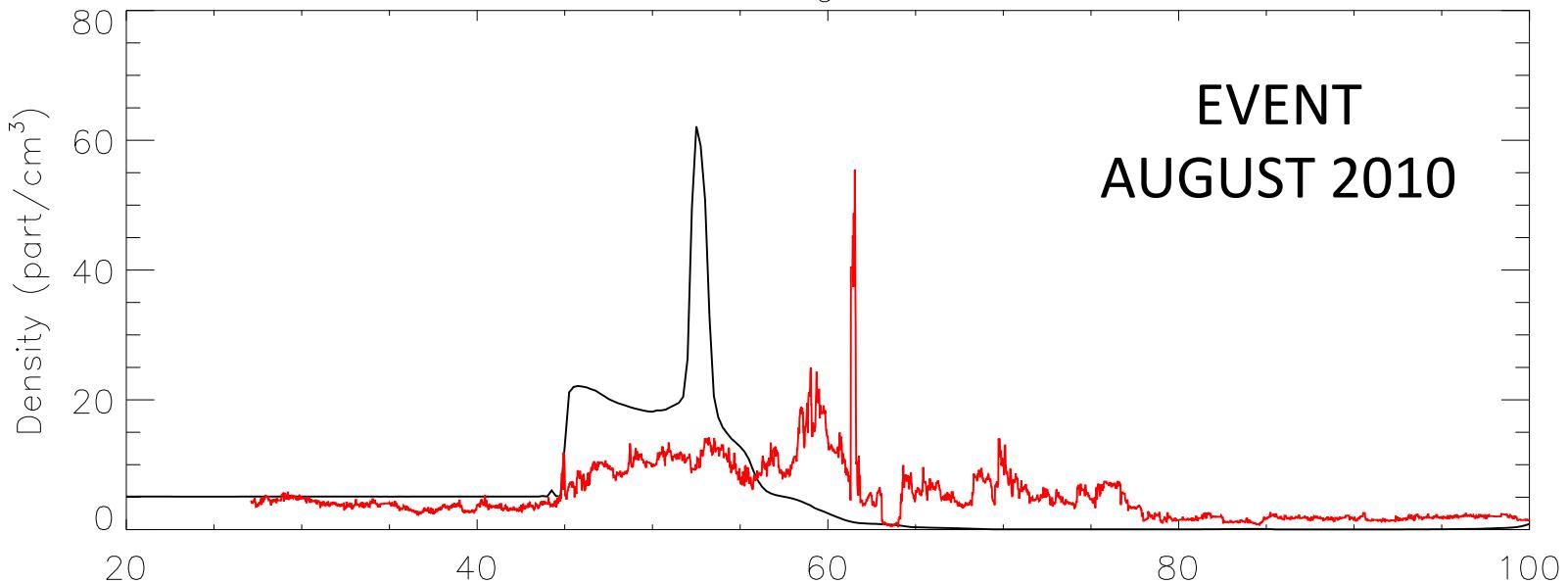
23-May-2010

EVENT
MAY 2010



01-Aug-2010

EVENT
AUGUST 2010



FITTING PROBLEMS

Arrival Time +/- 8 hours

Arrival Velocity +- 100 km s⁻¹

Arrival Density +50 cm⁻³

Compression Region Duration -10 hours

Rarefaction Region Duration -15 hours

WHY?

- MODEL (1)

- OBSERVATIONS

CME: velocity, mass loss rate, injection time

SW: velocity, mass loss rate

The mass loss rate depends on the mass, the solid angle and the injection time

BACK TO BASICS

- 1 CME
- 1 Dimension
- Change the input set parameters
(5)
- FIX 4 and 1 free

CME

$$v = 1330 \text{ km s}^{-1}$$

$$m = 1.1 \times 10^{13} \text{ kg}$$

$$n = 1.5 \text{ cm}^{-3}$$

$$Dt = 2 \text{ hr}$$

$$\dot{m} = 1.272 \times 10^{-13} M_{\odot} \text{ yr}^{-1}$$

$$\phi = 1.8 \text{ rad}$$

VRSNAK ET AL 2010

EVENT

July 25, 2004

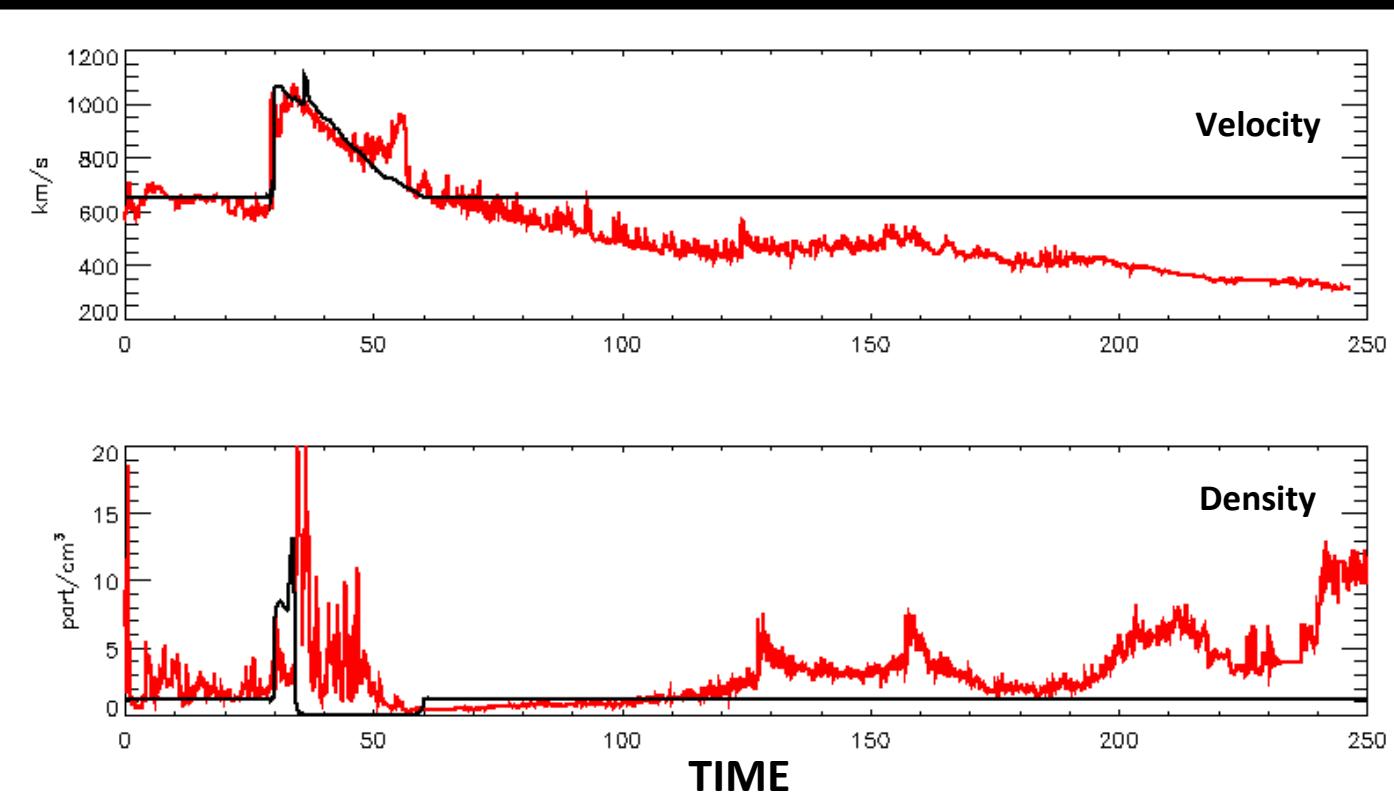
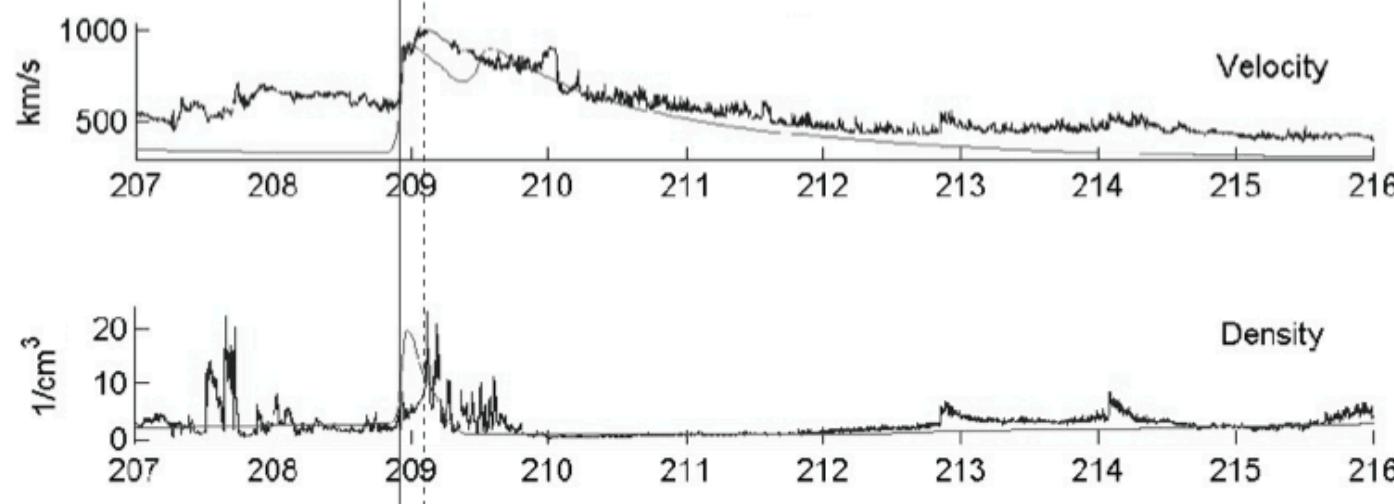
AMBIENT SOLAR
WIND
(SW)

$$v_0 = 650 \text{ km s}^{-1}$$

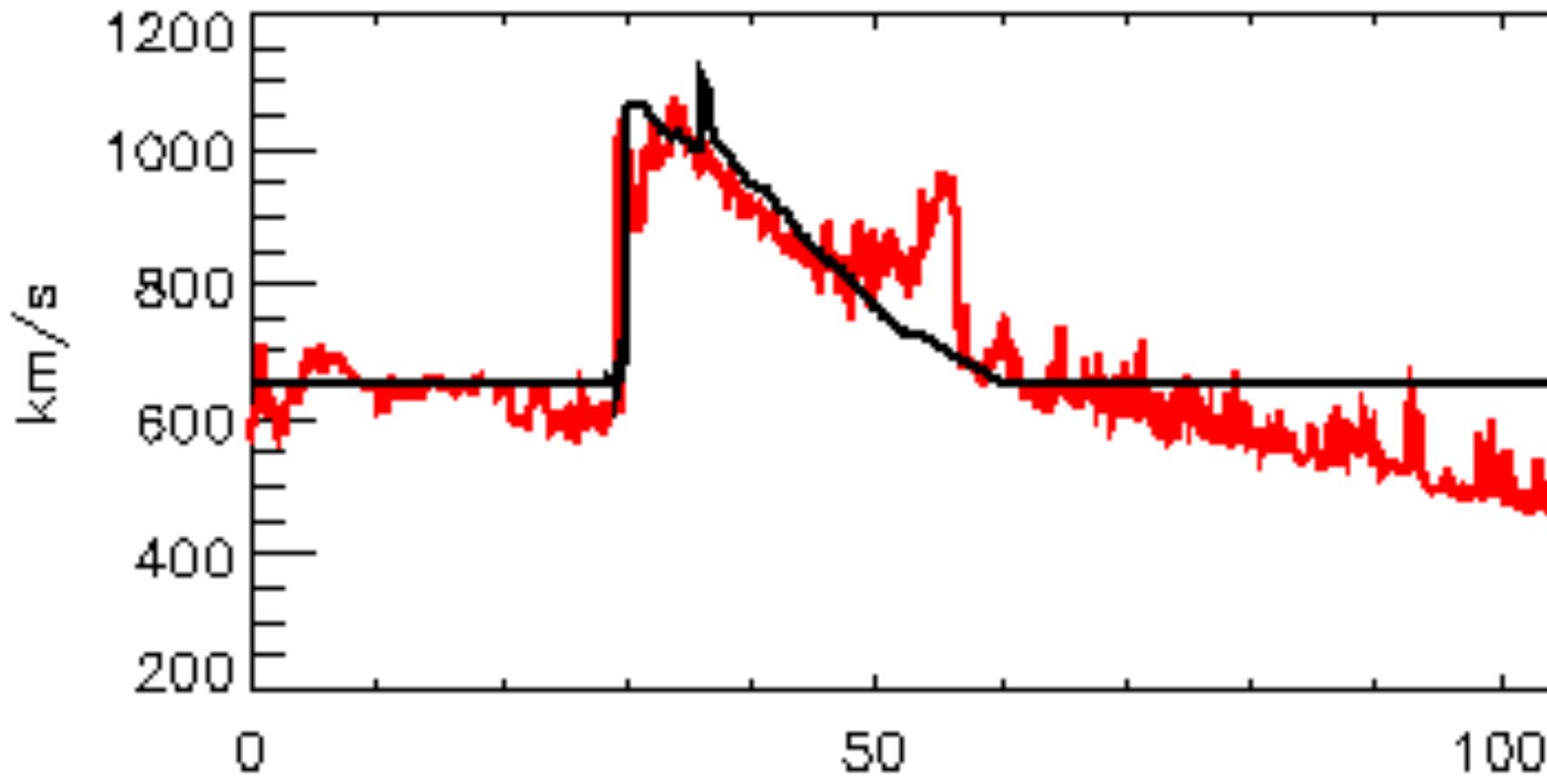
$$\dot{m}_0 = 7.26 \times 10^{-15} M_{\odot} \text{ yr}^{-1}$$

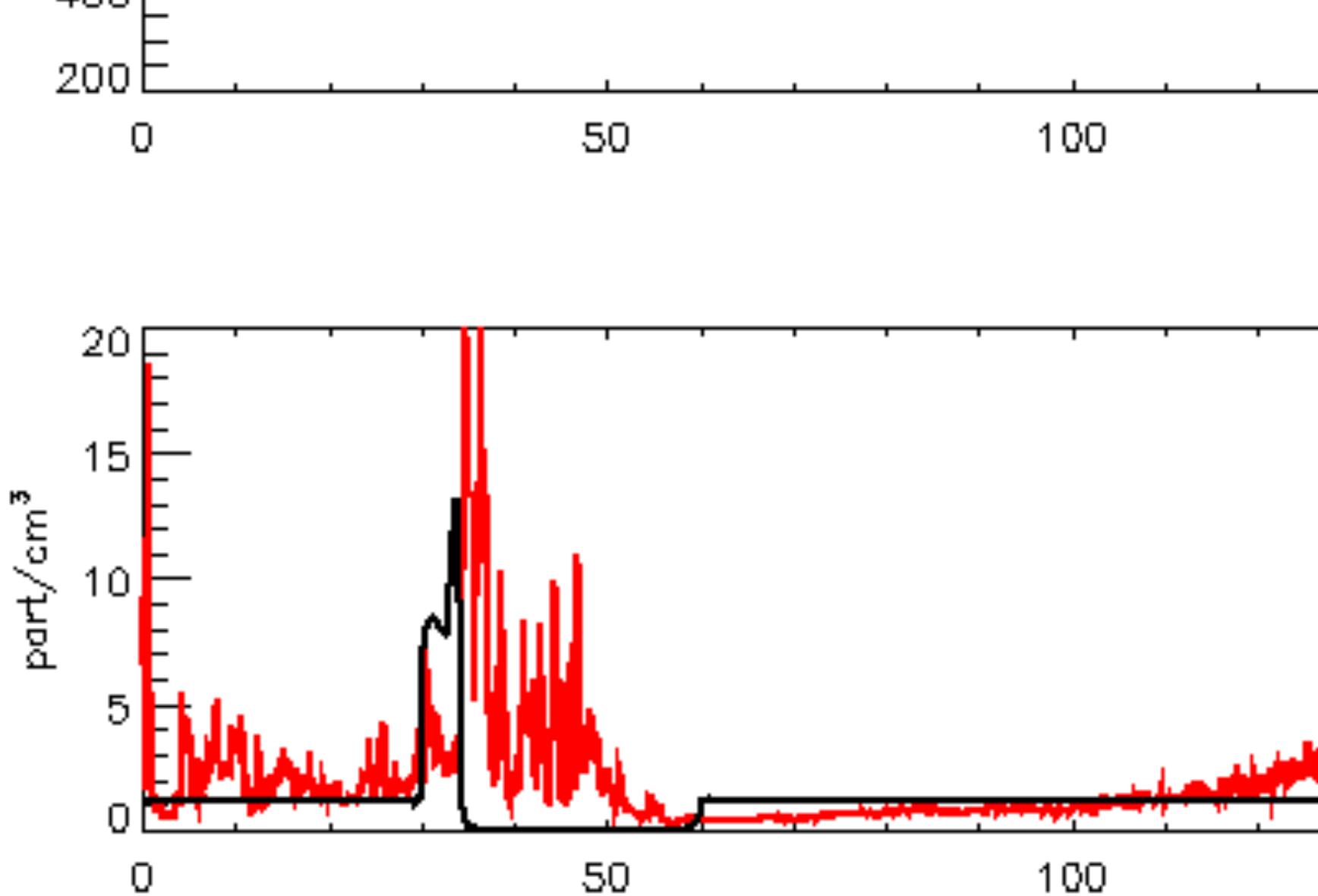


ENLIL



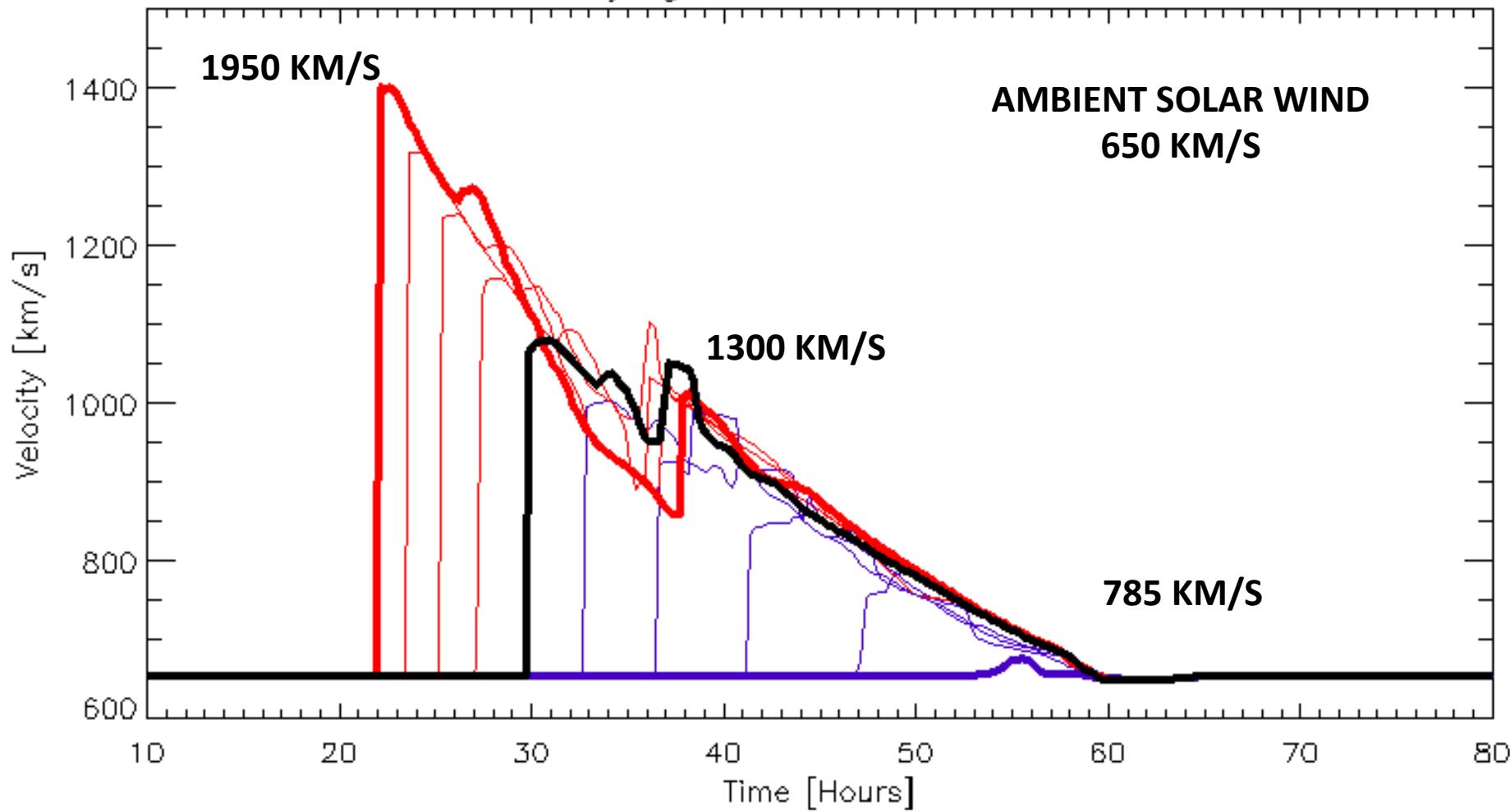
YGUAZÚ



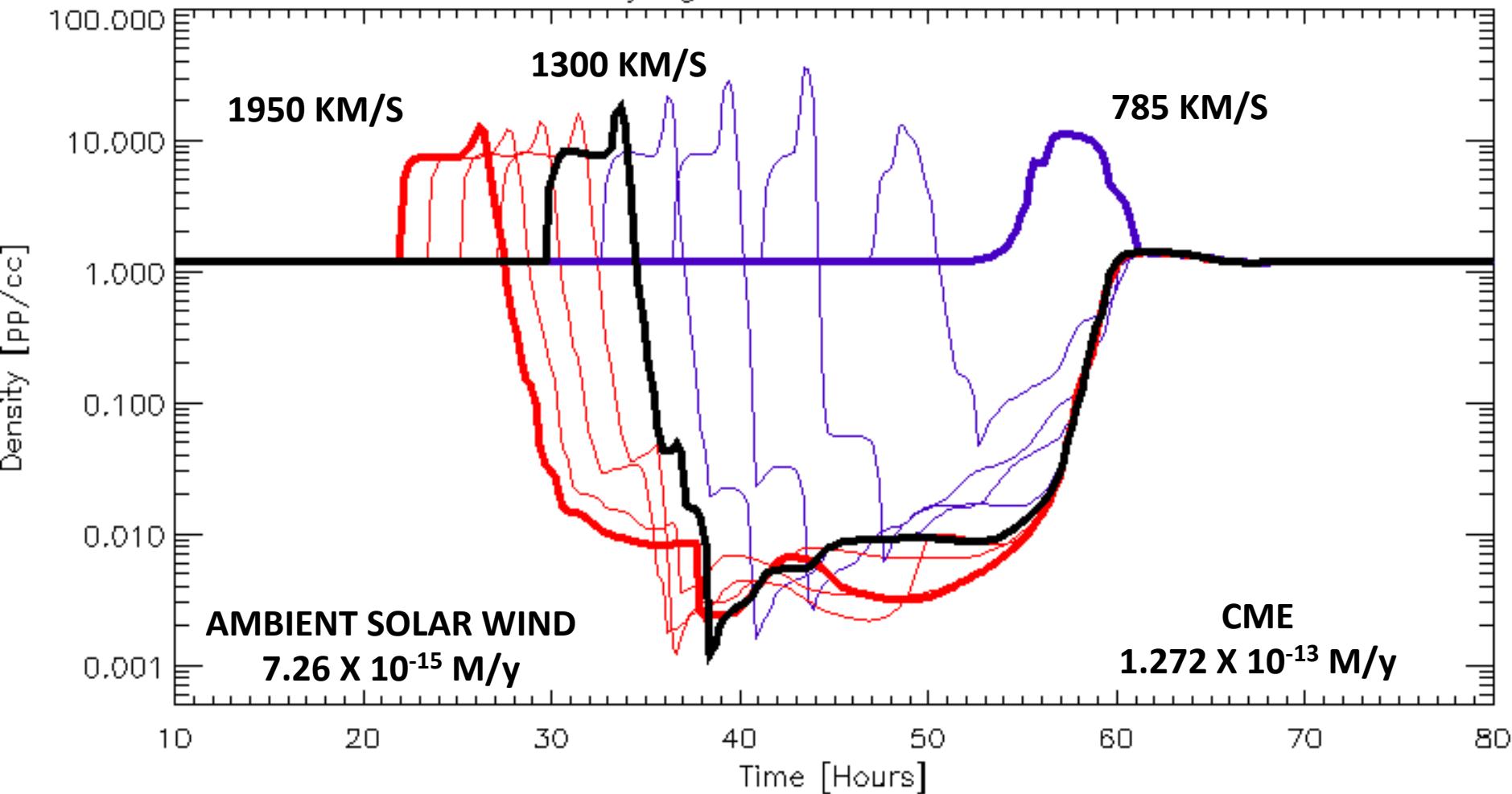


CHANGING THE VELOCITY OF THE CME

Varying CME INITIAL VELOCITY

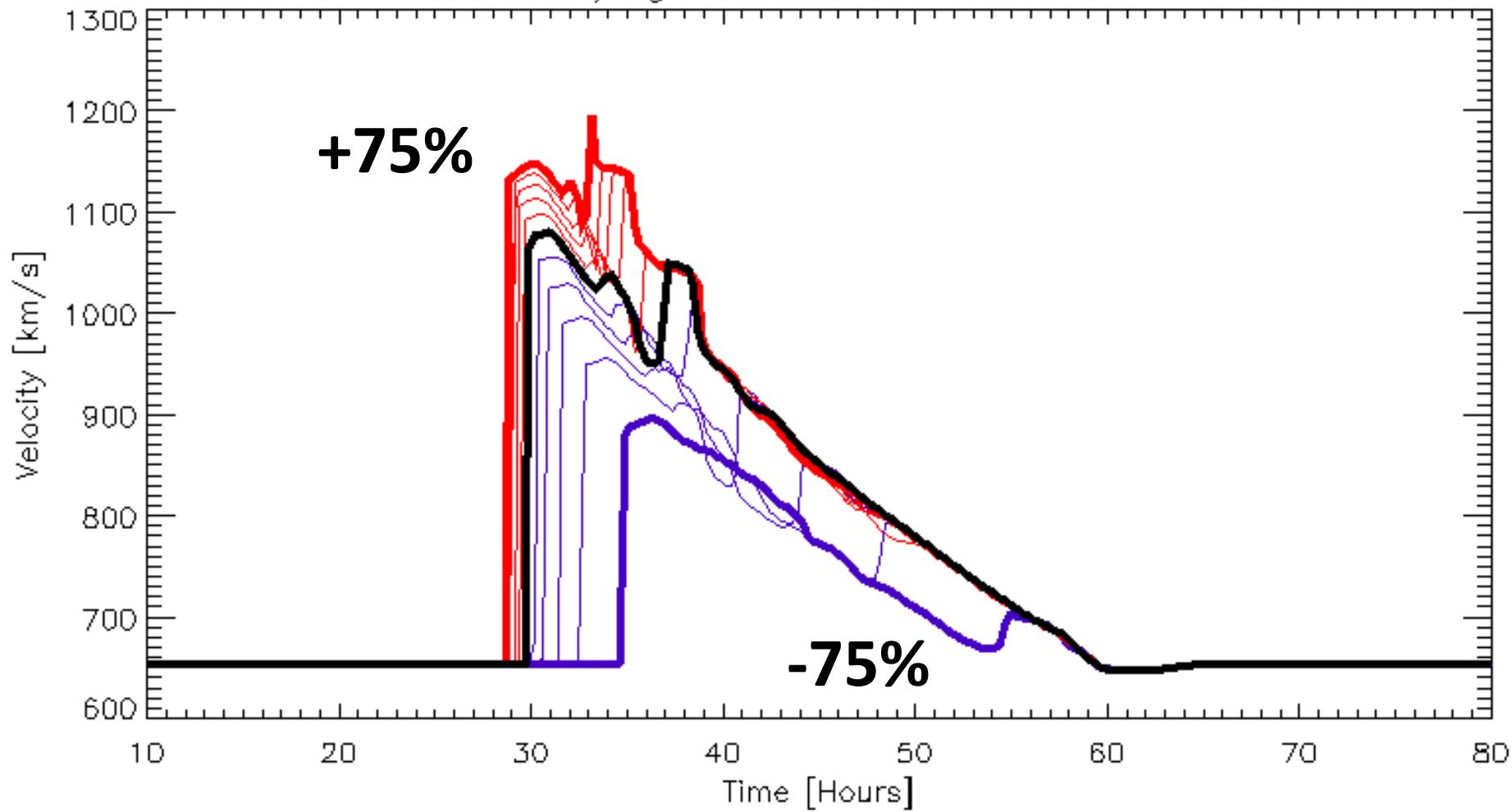


Varying CME INITIAL VELOCITY

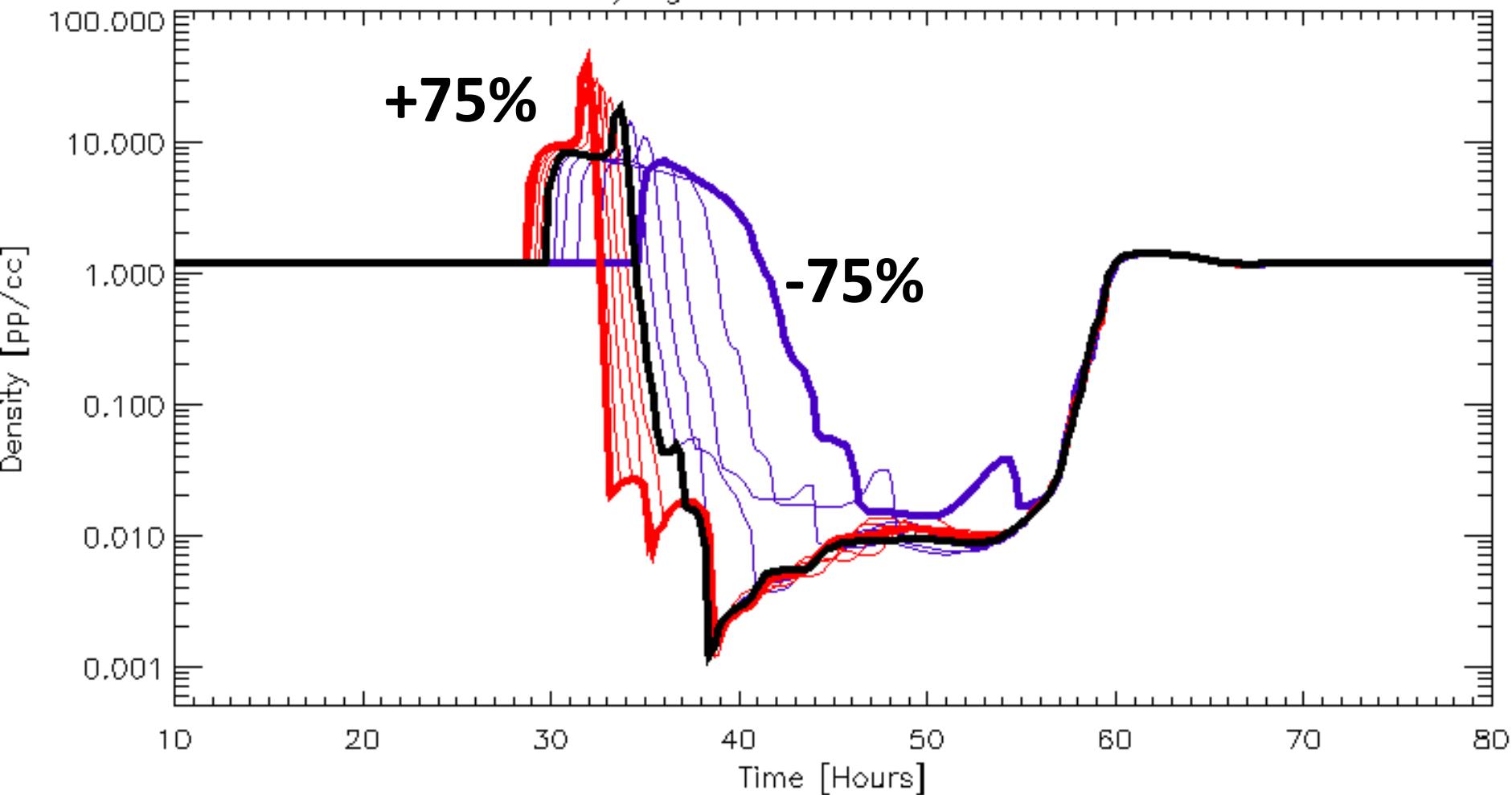


CHANGING THE MASS LOSS RATE OF THE CME

Varying CME MASS-LOSS RATE

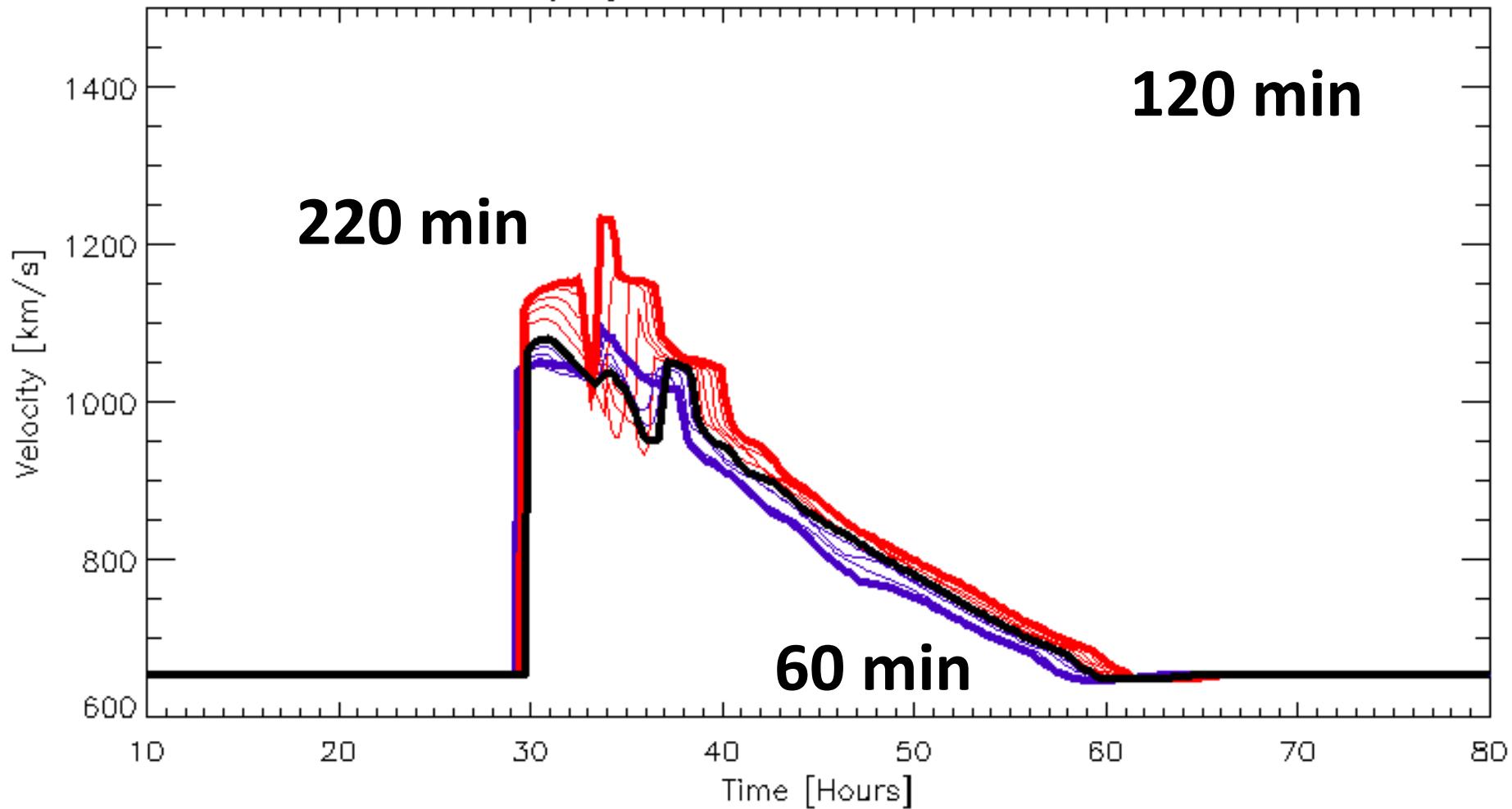


Varying CME MASS-LOSS RATE

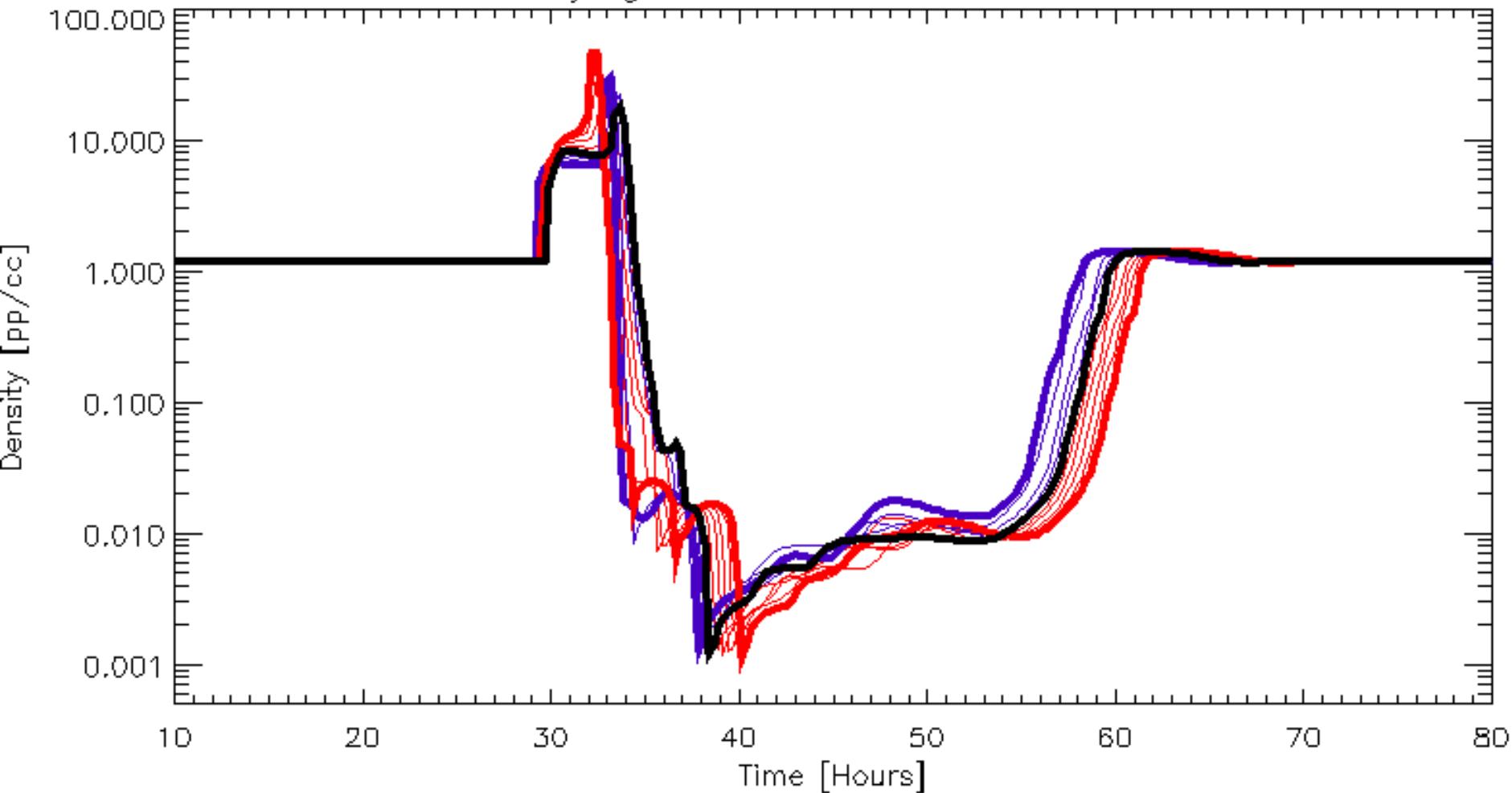


CHANGING THE INJECTION TIME OF THE CME

Varying CME INITIAL INJECTION TIME

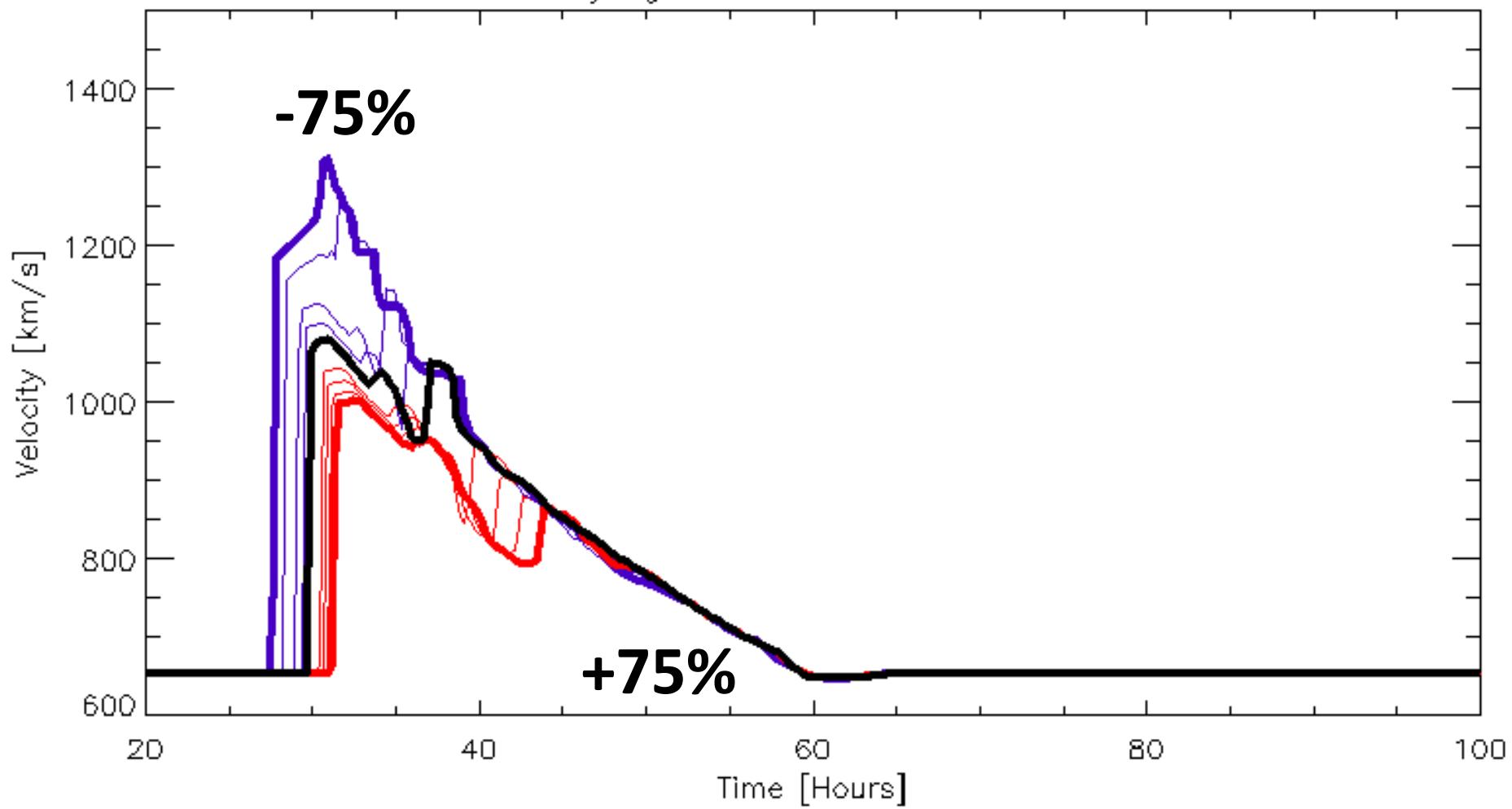


Varying CME INITIAL INJECTION TIME

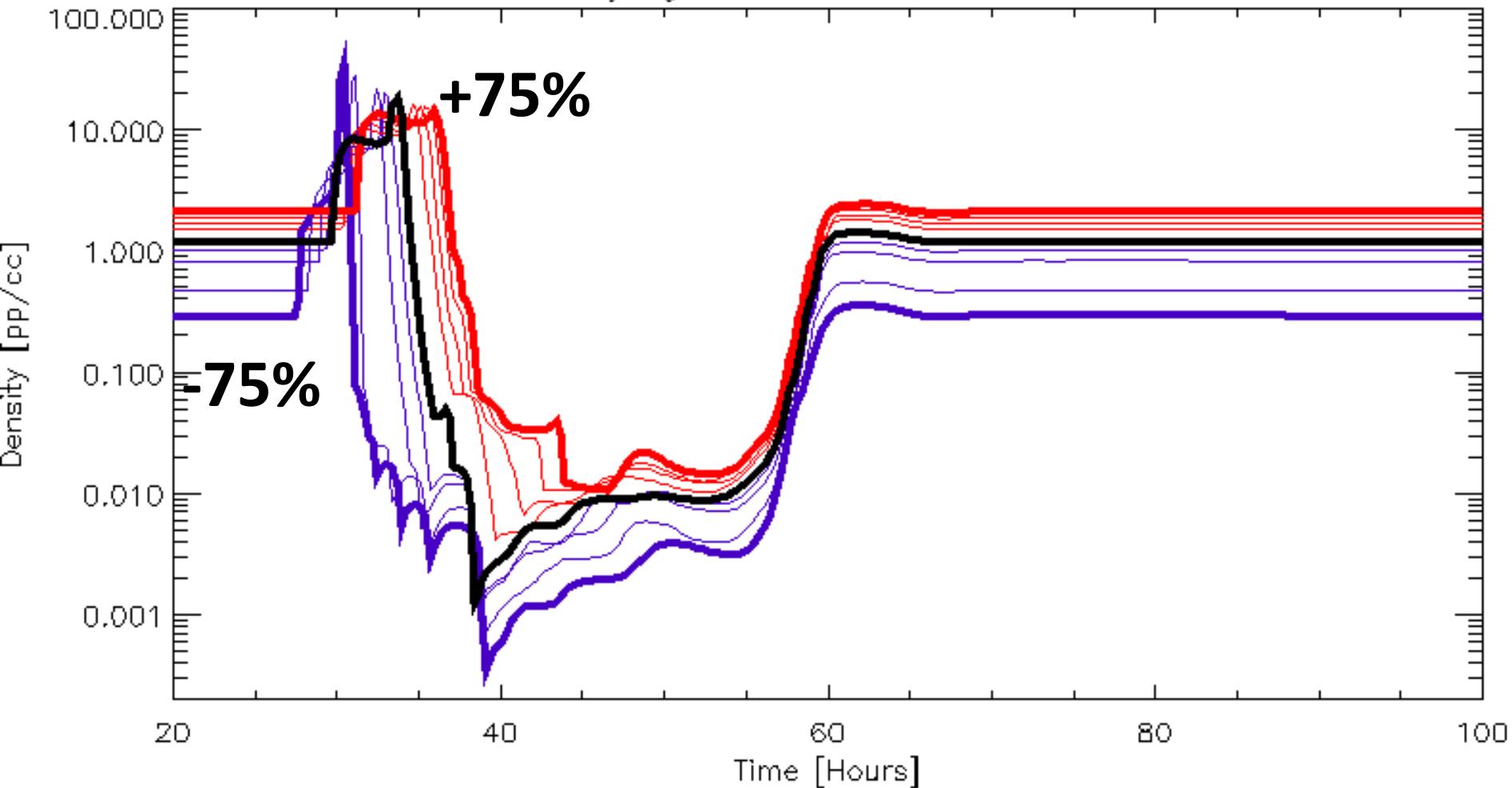


CHANGING THE
MASS LOSS RATE OF THE SW

Varying ASW MASS-LOSS RATE

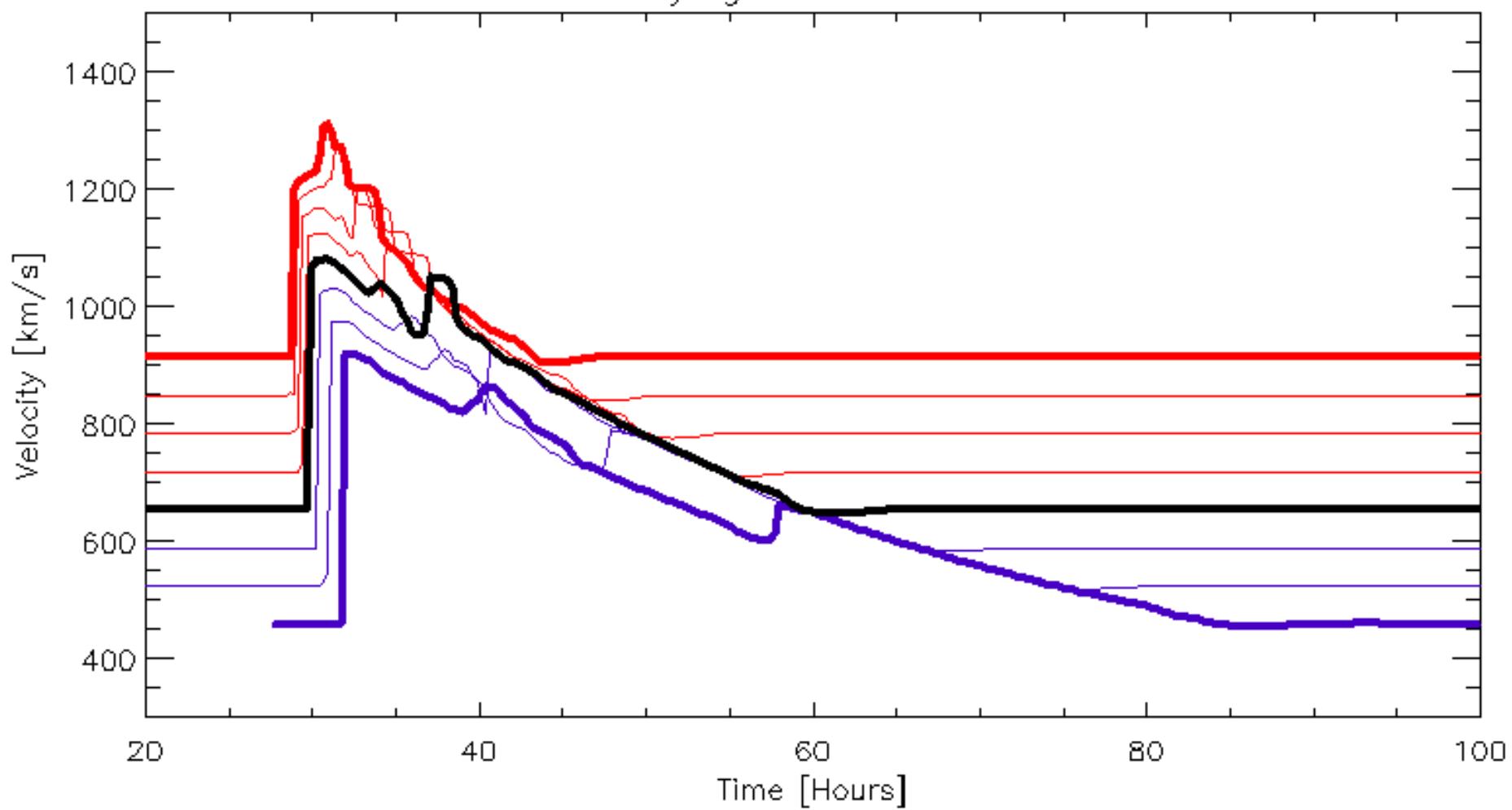


Varying ASW MASS-LOSS RATE

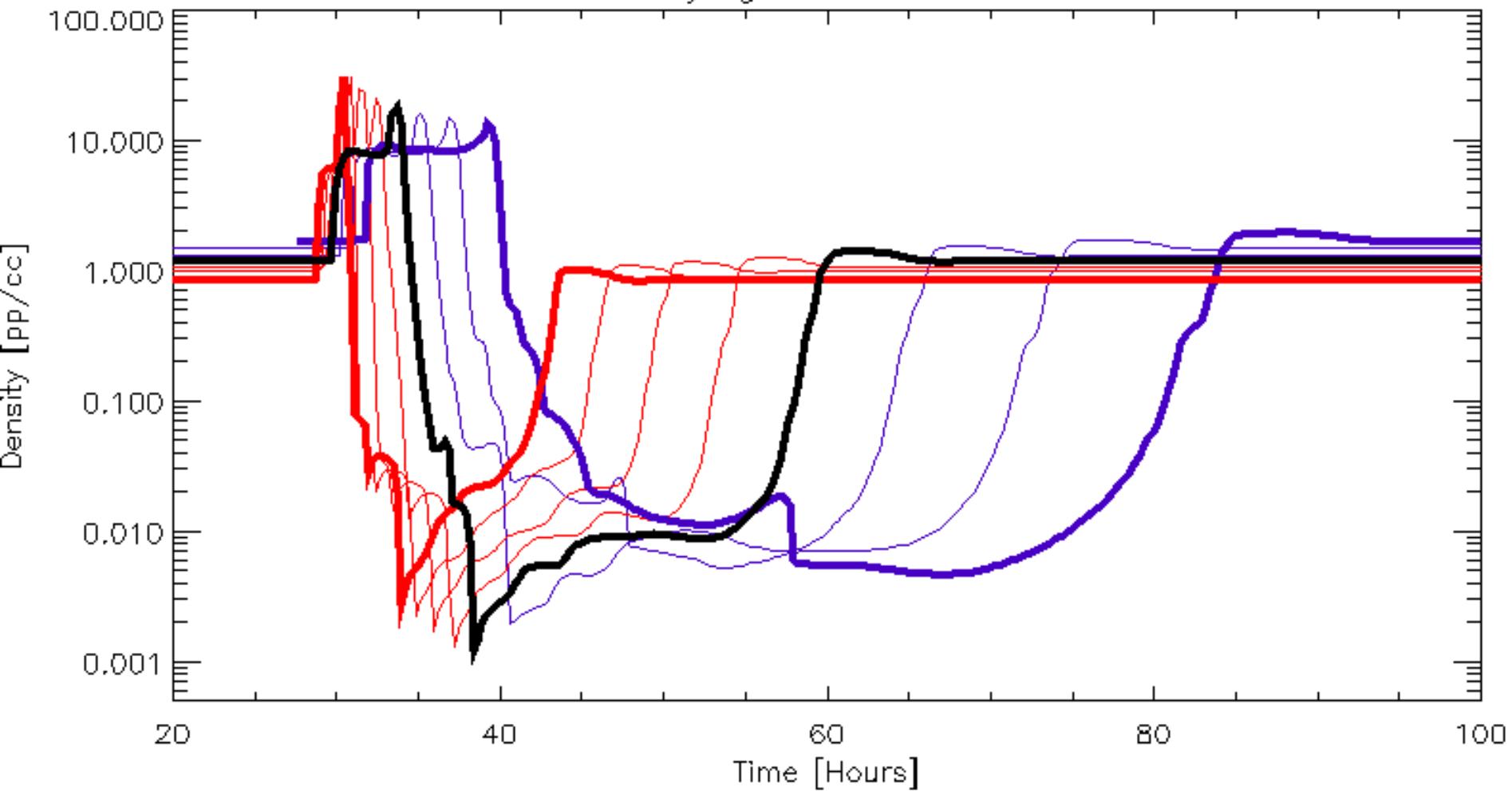


CHANGING THE
VELOCITY OF THE SW

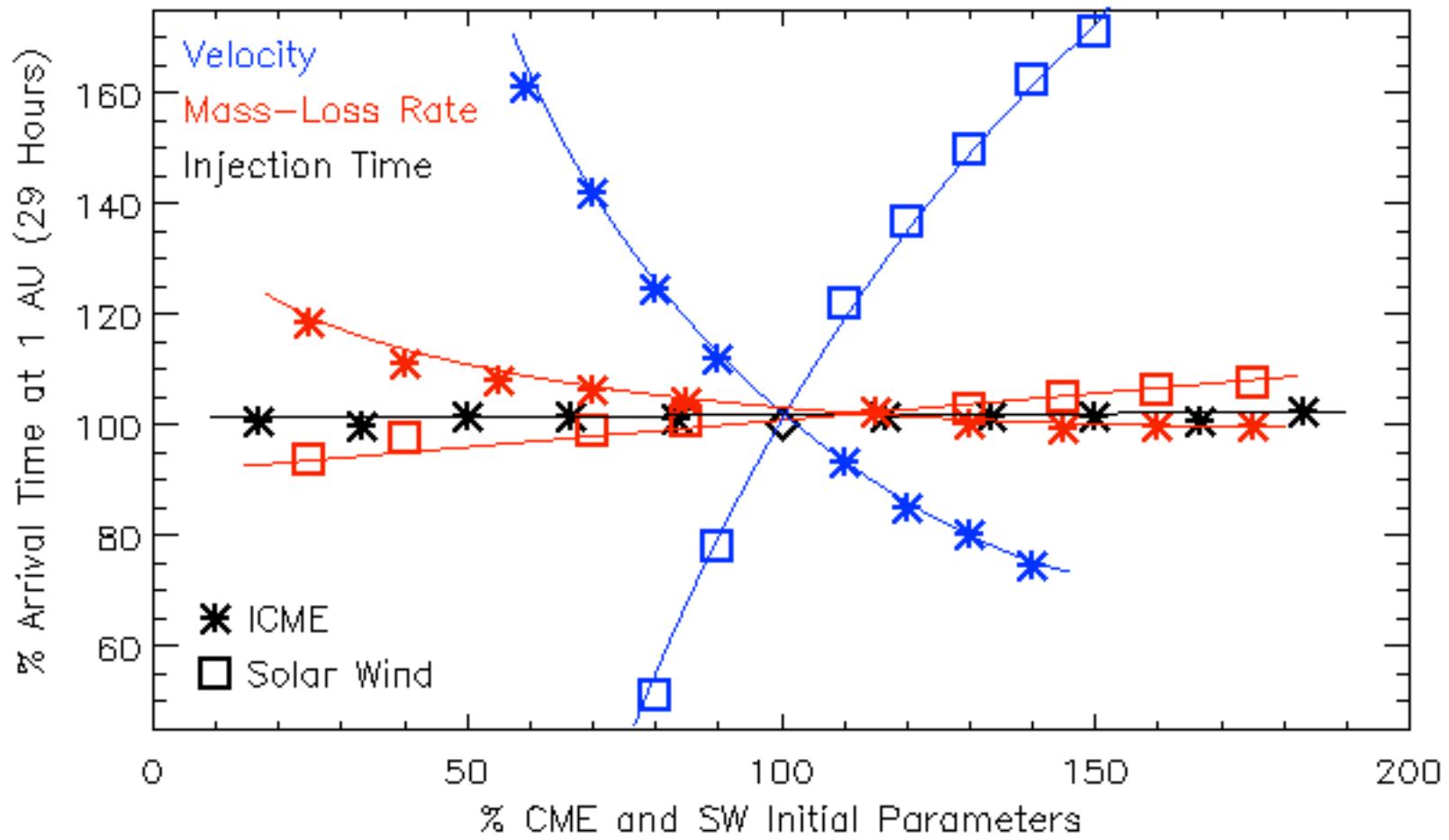
Varying ASW VELOCITY

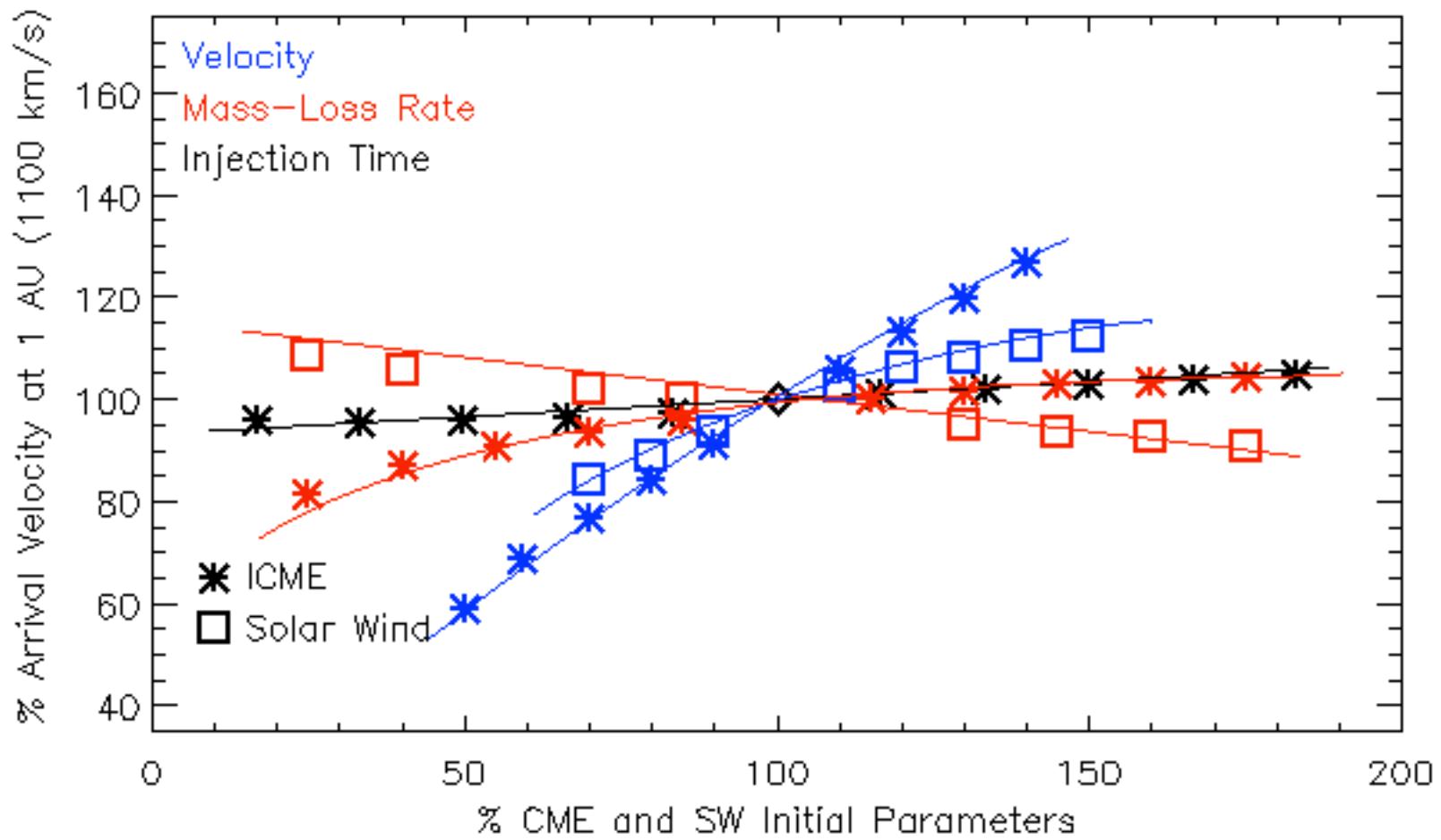


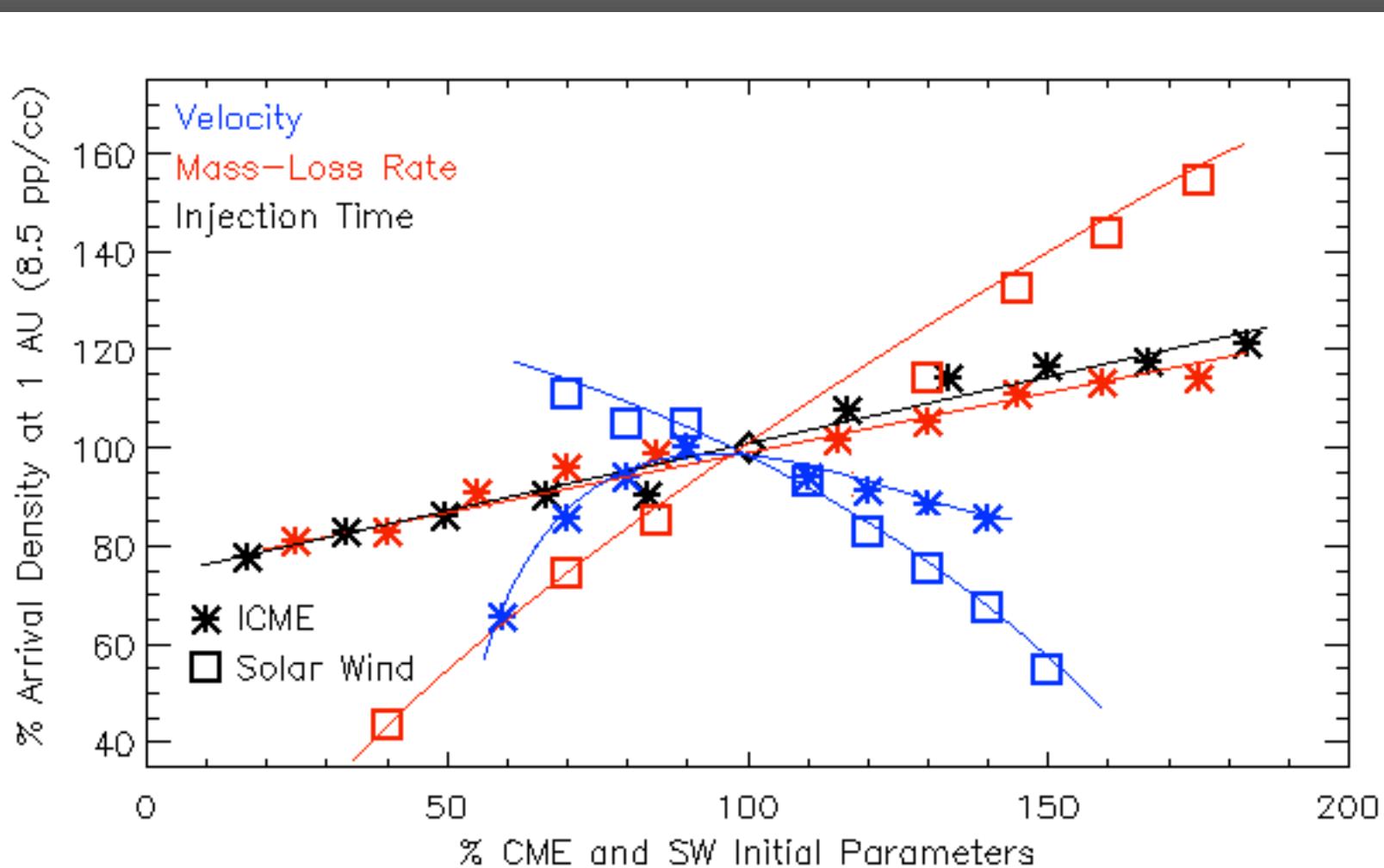
Varying ASW VELOCITY

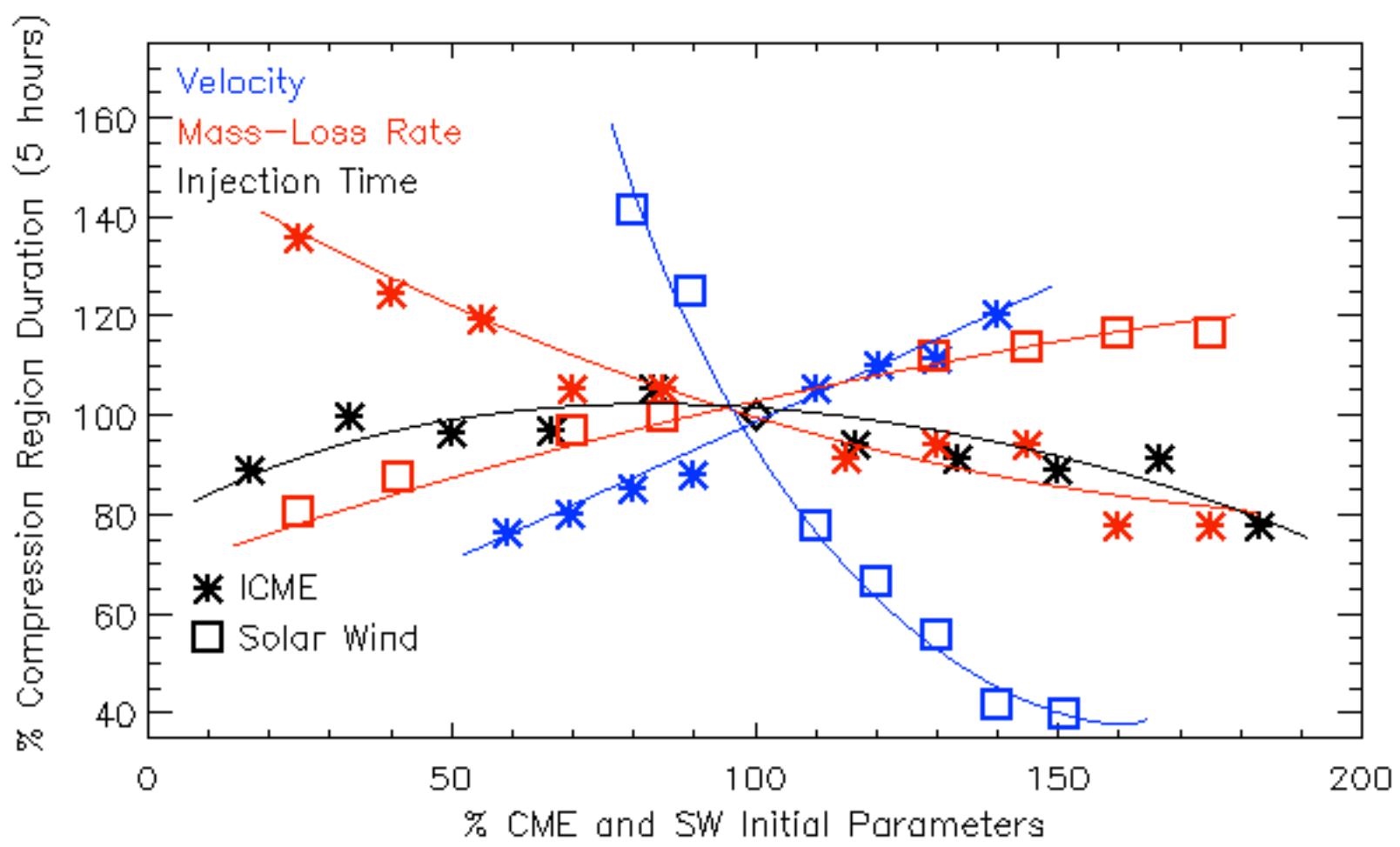


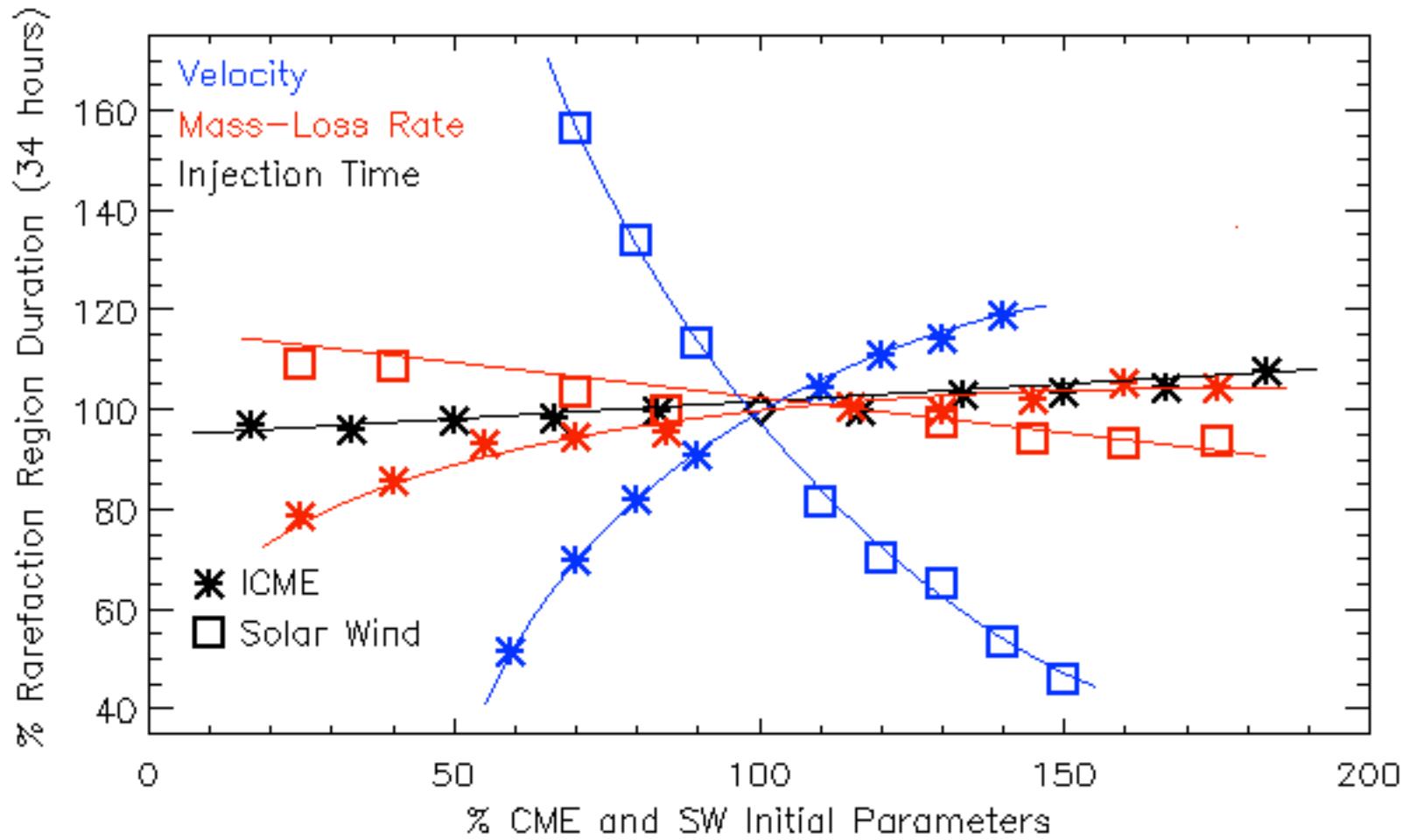
- **Arrival Time**
 - **Arrival Velocity**
 - **Arrival Density**
-
- **Duration of the compression region**
 - **Duration of the rarefaction zone**











BACK TO THE CME – CME INTERACTION

ADDITIONALLY WE CAN VARY THE AMBIENT SOLAR WIND
BETWEEN THE TWO CMES.

FOR THE EVENT OF MAY 2010

AMBIENT SOLAR WIND

320 km/s	340 km/s
$2 \times 10^{-14} \text{ M/y}$	$1.7 \times 10^{-14} \text{ M/y}$

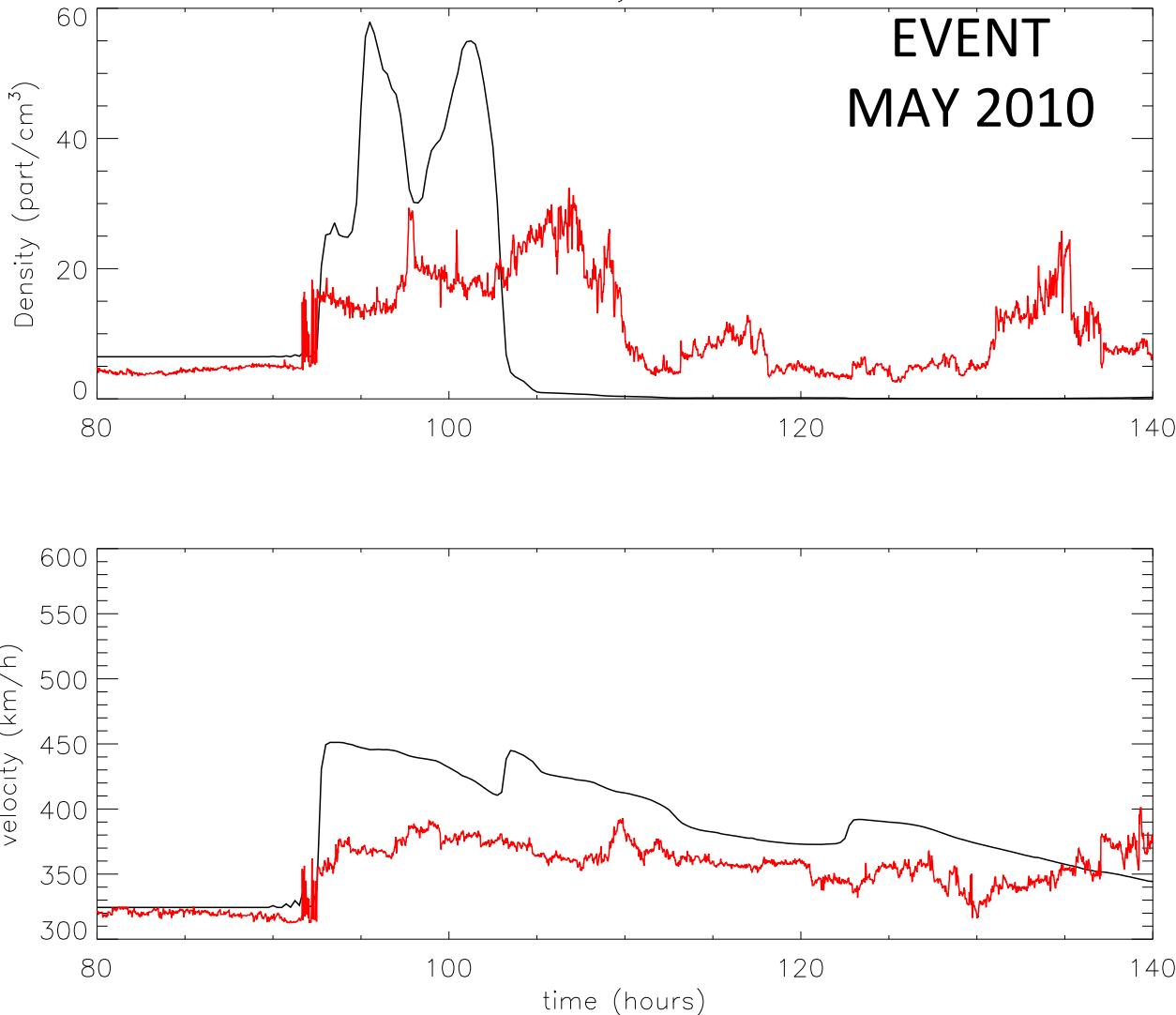
1st CME

400km/s	650 km/s
$1.5 \times 10^{16} \text{ g}$	$1.0 \times 10^{16} \text{ g}$

2nd CME

23-May-2010

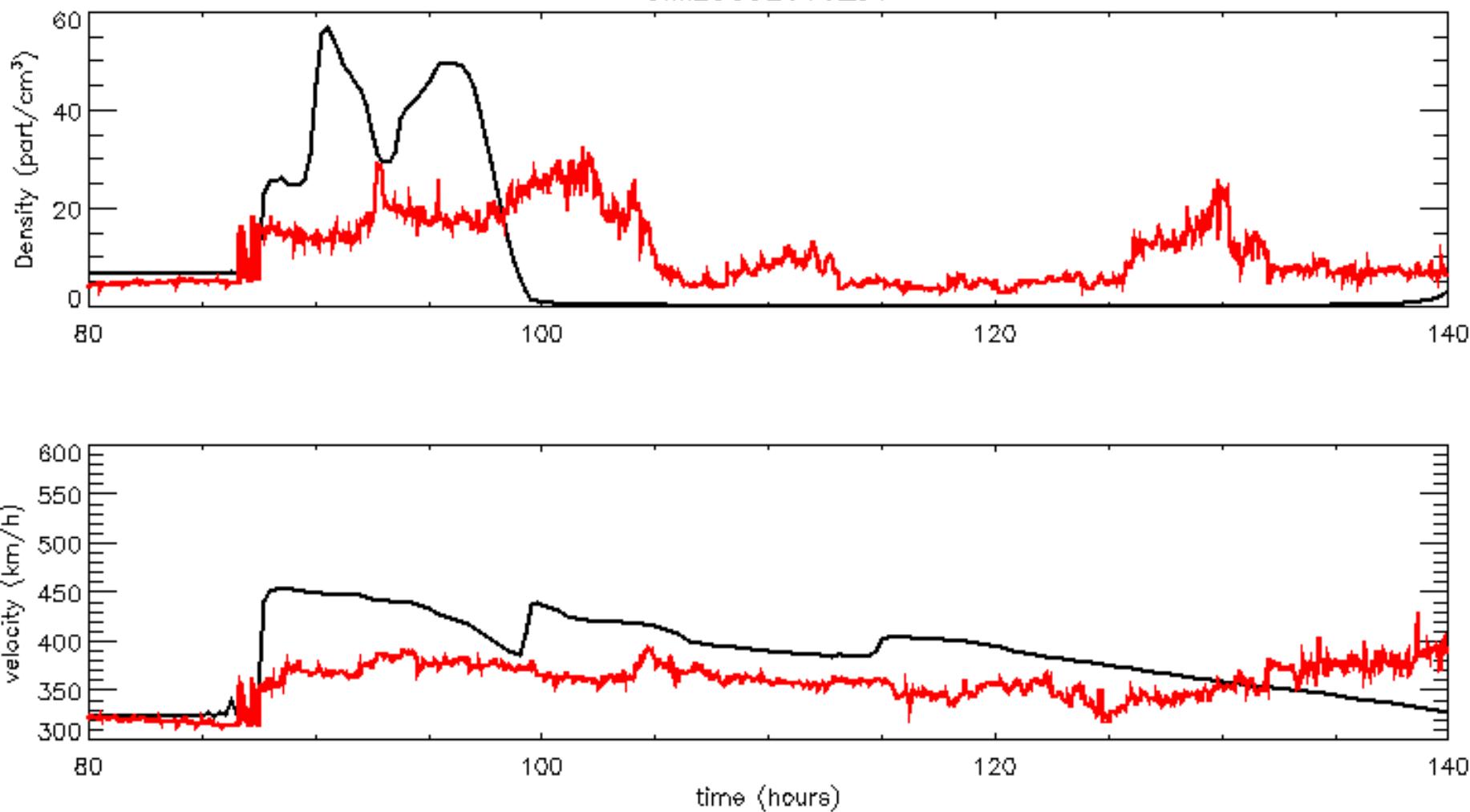
EVENT
MAY 2010



AMBIENT SOLAR WIND
320 km/s
 2×10^{-14} M/y

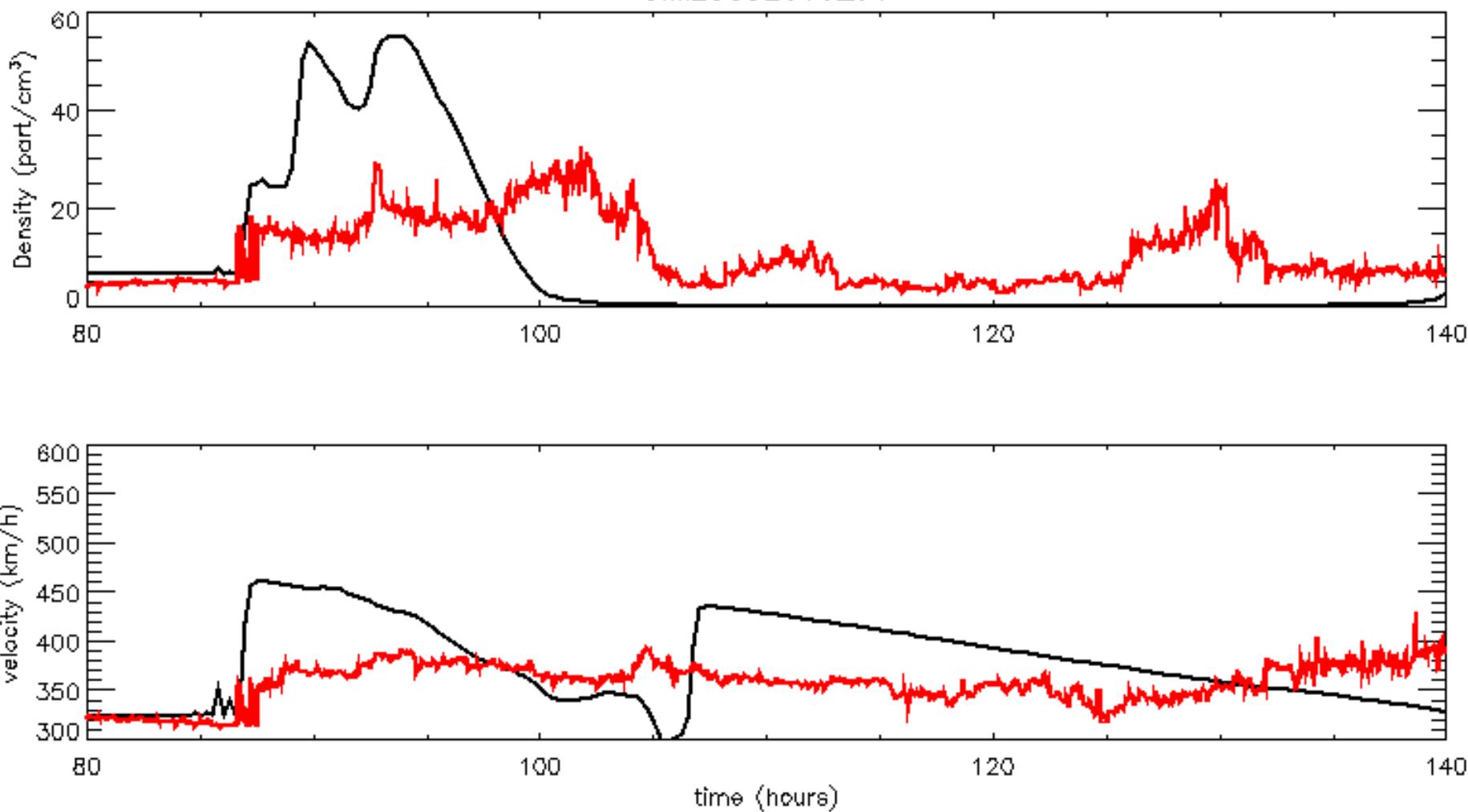


SIM23052010_56



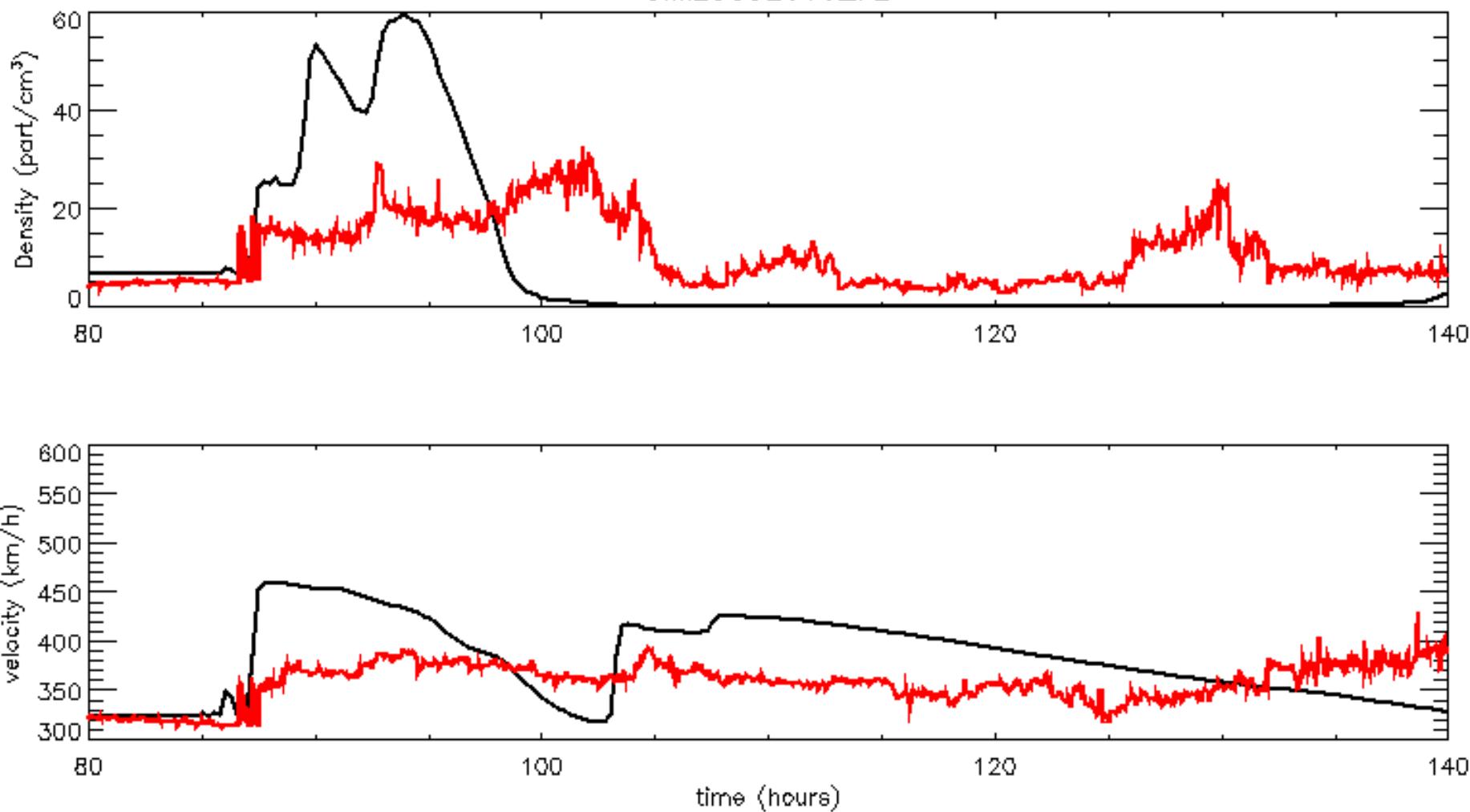
AMBIENT SOLAR WIND
340 km/s
 2×10^{-14} M/y

SIM23052010_64



AMBIENT SOLAR WIND
320 km/s
 1.7×10^{-14} M/y

SIM23052010_72



AMBIENT SOLAR WIND
340 km/s
 1.7×10^{-14} M/y

