

Nov. 13, 2012

①

CH5. Heat Transfer by Convection
Energy $v \neq 0$

①

Radiation

$v=0$, high T

②

conduction

$v=0$, high P,

③

convection

$v \neq 0$, high $\left| \frac{dT}{dr} \right|$

All three ways of energy transfer are proportional to $\left| \frac{dT}{dr} \right|$

Fick's Law of Diffusion: $F = -D \frac{d\phi}{dr}$

$$\text{or } F = -D_T \frac{dT}{dr}$$

$$\textcircled{1} \text{ Radiation } F = -D \frac{dU}{dr} = -D_T \frac{dT}{dr} = -\frac{4ac}{3kp} T^3 \frac{dT}{dr}$$

$$D_{\text{rad}} = \frac{1}{3} c \lambda_{\text{ph}}$$

$$D_T = \frac{4ac}{3} \frac{1}{kp} T^3$$

$$U = aT^4$$

$$\textcircled{2} \text{ Conduction } F = -D \frac{dQ}{dr} = -D_T \frac{dT}{dr} = -\frac{1}{3} C_v v_e \lambda_e \frac{dT}{dr}$$

$$D_{\text{cond}} = \frac{1}{3} v_e \lambda_e, \quad Q = C_v T$$

$$D_T = \frac{1}{3} C_v v_e \lambda_e$$

$$\textcircled{3} \text{ Convection } F = -D \frac{dQ}{dr}$$

$$F = \rho w C_p \Delta T$$

ρ : density ; C_p : specific heat

w : average convection velocity

ΔT : temperature difference between the convection parcel and ambient

we need to find w and ΔT , given P, ρ, T [Q2]

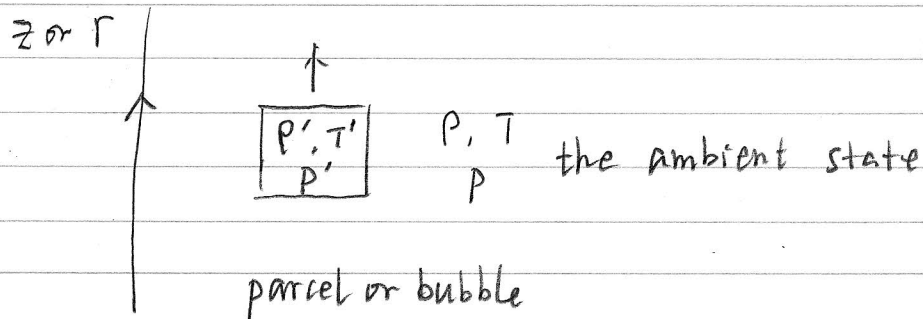
* In ZAMS, F_{rad} dominates

$$F_{\text{rad}} \sim \frac{dT}{dr}$$

However, $\left| \frac{dT}{dr} \right| > \left| \frac{dT}{dr} \right|_{\text{critical}} \Rightarrow$ convection occurs [Q1]

(2)

Criteria for Convection (CH 5.1.1)

* Convection is caused by high temperature gradient* the movement of convection parcel is driven by the buoyancy forceConsider a small disturbance of rise-up motion: w .

① the parcel is adiabatic (for now)

② the pressure internal is always in equilibrium

with external: $P' = P$, (since $w \ll c_s$

$$P = nkT = \frac{\rho}{\mu} NA kT \quad \text{sound speed}$$

$$\text{If } T' > T \Rightarrow \rho' < \rho$$

The buoyancy force (Archimedes' principle)

$$F = \rho g - \rho' g = (\rho - \rho') g$$

If internal ρ' smaller, positive F , continue to riseIf " " larger, negative F , return to equilibrium
(quenching the disturbance)

Thus, the criterion is

$$\left| \frac{dT'}{dz} \right| < \left| \frac{dT}{dz} \right| \quad \text{large temperature gradient in ambient}$$

$$\text{Since } \frac{dT}{dz} < 0, \quad \frac{dT'}{dz} > \frac{dT}{dz}$$

$$\text{Introduce } \beta: \quad \boxed{\beta = -\frac{dT}{dz}} \Rightarrow \beta' < \beta$$

(3)

β (or $-\frac{dT}{dz}$) is determined by the stellar model

β' (or $-\frac{dT'}{dz}$) = β_{ad} , often consider an adiabatic process

β_{ad} can be determined well, because of "ad" constraints

$$\beta = -\frac{dT}{dz} = -\frac{dT}{dp} \frac{dp}{dz} = -\frac{T}{p} \frac{dT}{T} \frac{dp}{p} \frac{dz}{dz}$$

$$\beta = -T \frac{d \ln T}{d \ln p} \frac{d \ln p}{dz}$$

Recall pressure scale height λ_p

$$\lambda_p = -\left(\frac{d \ln p}{dz}\right)^{-1} = \frac{p}{g_p} = \frac{V_s^2}{g_p}$$

$$p = p_0 e^{-\frac{z}{\lambda_p}}$$

$$\Gamma_1 = \left(\frac{\partial \ln p}{\partial \ln p}\right)_{ad} = \frac{p}{p} \left(\frac{\partial p}{\partial p}\right)_{ad}$$

$$\left(\frac{\partial p}{\partial p}\right)_{ad} = V_s^2. \quad V_s: \text{ sound speed}$$

Recall the definition of $\left(\frac{d \ln T}{d \ln p}\right)_{ad}$

$$\nabla = \frac{p_2 - 1}{p_2} = \left(\frac{d \ln T}{d \ln p}\right)_{ad} \quad \text{--- 3.94}$$

∇ is logarithm temperature gradient with respect to pressure.

$$\text{Thus, } \beta = \frac{T}{\lambda_p} \nabla.$$

" "

* ∇ , a convenient dimensionless number

$$\text{* For ideal gas, } p_2 = \frac{5}{3}, \quad \nabla_{ad} = \frac{\frac{5}{3} - 1}{\frac{5}{3}} = \frac{2}{5} = \underline{0.4}$$

* For mixture (radiation plus ideal gas).

$$\nabla_{ad} = \frac{2(4 - 3\beta)}{32 - 24\beta - 3\beta^2}, \quad \beta = \frac{p_g}{p_g + p_{rad}}$$

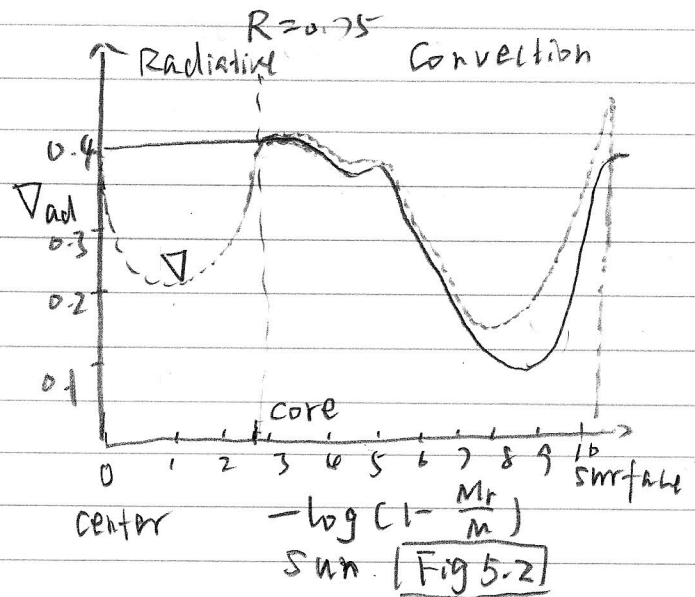
* For ionization, ∇_{ad} is a function of Y , $\rightarrow 0.1$

Thus, the criteria for convection

$$\left| \frac{dT}{dz} \right| > \left| \frac{dT'}{dz} \right|$$

$$\beta > \beta_{ad}$$

$$\nabla > \nabla_{ad}$$



x axis of $-\log(1 - \frac{M_r}{M})$ emphasize the surface

$$x = 1, 1 - \frac{M_r}{M} = 10^{-1} \Rightarrow \frac{M_r}{M} = 1 - 10^{-1}$$

or 10^{-1} (10%) mass outside

$1 - 10^{-1}$ (90%) mass inside

$$x = 3, \frac{M_r}{M} = 1 - 10^{-3}; 10^{-3} (0.1\%) \text{ mass outside}$$

99.9% mass inside

$x = 2.5$: separate the convection zone ($R \approx 0.75$)
convection zone has a mass $< 1\%$

Mixing Length Theory (CH 5.1) — MLT

How to find F_{cond} ?How to find w and ΔT ?

Use MLT to simplify the solution.

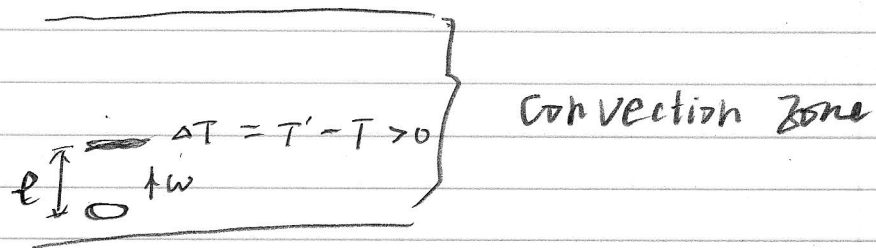
MLT: Assume the convective parcel travels a characteristic distance " l ", the mixing length, before the parcel breaks up, lose identity and merge with the ambient fluid.

" l " is usually a fraction of " λ_p "

To find w and ΔT , we need to look into

① the energy equation (CH 5.1-2)

② the equation of motion (CH 5.1-3)



(5)

Energy Equation (Radiative Leakage CH 5.1.2)

Since $v \neq 0$, $\frac{DA}{Dt} = \frac{\partial A}{\partial t} + \vec{w} \cdot \nabla A$: A any variable

1D $\frac{DT}{Dt} = \frac{dT}{dt} + w \frac{dT}{dz} = A = T$

↑
Total derivative
Lagrangian form

↑
Local derivative
Euler form

↖
Advective
(convective)
component.

Thus, for ambient: $\frac{dT'}{dt} = 0$ (no local change)

$$\frac{DT}{Dt} = w \frac{dT}{dz} = -\beta w$$

For the parcel:

$$\frac{DT'}{Dt} = \frac{dT'}{dt} + w \frac{dT'}{dz} \quad \text{--- (5.18)}$$

For adiabatic convection process only,

$$\frac{DT'}{Dt} = -\beta_{ad}$$

$$\Rightarrow \frac{DT'}{Dt} = -\beta_{ad} w$$

$$\boxed{\frac{D(\Delta T)}{Dt} = (\beta - \beta_{ad}) w}$$

5.21

$$\Delta T = T' - T$$

The energy equation in adiabatic case

x non-adiabatic: $\boxed{\frac{DT'}{Dt} \neq 0}$ radiation from parcel to the surrounding is considered.

The parcel cools down due to radiation because of $T' > T$; this is in addition to adiabatic expansion cooling.

* Radiative Leakage

$$\frac{dQ'}{dt} = -\nabla \cdot F_{rad} \quad \text{--- (5.11)}$$

$dQ' = C_p \rho dT'$, C_p : due to pressure equilibrium

$$F_{rad} = -K \nabla T$$

$$K \equiv \frac{4a c T^3}{3 \kappa \rho}$$

$$\nabla \cdot (\nabla T) = \nabla^2 T \quad \text{Laplace}$$

$$\Rightarrow \left(\frac{\partial T'}{\partial t} \right) = \frac{4a c T^3}{3 \kappa \rho^2 C_p} \nabla^2 T \quad \text{--- (5.15)}$$

Introduce the thermal diffusivity ν_T

$$\boxed{\nu_T = \frac{4a c T^3}{3 \kappa \rho^2 C_p}} \quad \text{--- (5.16)}$$

$$\left(\frac{\partial T'}{\partial t} \right) = \nu_T \nabla^2 T \quad \text{--- (5.17)}$$

To find $\nabla^2 T$, we resort to mixing length

$$\nabla^2 T = \frac{T - T'}{l^2}$$

Add $\left(\frac{\partial T'}{\partial t} \right)$ term to Eq. (5.18)

$$\Rightarrow \frac{D \Delta T}{Dt} = (\beta - \beta_{rad}) w - \frac{\nu_T}{l^2} \Delta T \quad \text{--- (5.21)}$$

Assuming $\beta, \beta_{rad}, \nu_T, l^2$ are known, find ΔT and w , need one more differential equation