

SEEDS (Solar Eruptive Event Detection System)

*Oscar Olmedo
Jie Zhang
Harry Wechsler
Art Poland
Kirk Borne*

*(George Mason University)
<http://solar.scs.gmu.edu/research/autocme>*

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Objectives

- ⊕ Develop an automatic system of solar transient events
 1. Detection (Current Work)
 2. Classification (Future Work)
 3. Association (Future Work) i.e. between EIT Dimming/CME/Flare

- ⊕ An automatic system is needed
 - ⊕ Explosive growth of data (STEREO and SDO)
 - ⊕ Human detection becomes costly or impossible
 - ⊕ Timely detection, for event catalog
 - ⊕ Timely detection, for space weather forecasting
 - ⊕ Objective detection, reducing human bias
 - ⊕ Accurate measurement
 - ⊕ Extracting more parameters
 - ⊕ Classification using high dimensional input data

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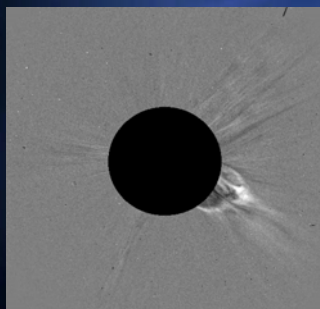
Methodology (For CME Detection)

- ⊕ Pre-Processing
 - ⊕ Apply filters
 - ⊕ Polar Transformation
- ⊕ Initial Detection
 - ⊕ Apply x-axis projection to find outstanding angles
 - ⊕ Parameter extraction (position angle, leading position, mean brightness, etc.)
- ⊕ Tracking and Detection
 - ⊕ Tracking CME's path
 - ⊕ Calculation of Velocity and Acceleration
- ⊕ Cataloguing

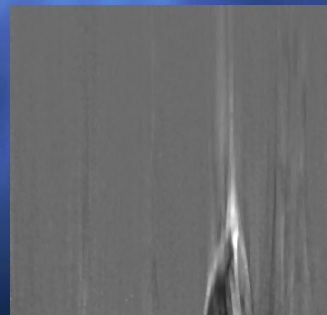
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Pre-Processing

- ⊕ Calibration
- ⊕ Median and Smoothing filters to reduce noise
- ⊕ Running Differencing to remove background
- ⊕ Polar Transformation for easy array manipulation



Running Difference image

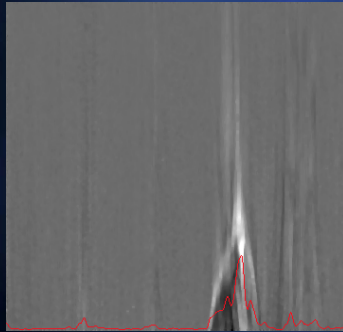


After polar Transformation

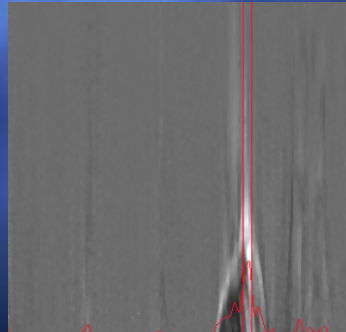
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Initial Detection

- ⊕ Detecting CME area or angles
 - ⊕ 1-D Projection
 - ⊕ Find CME core angles



1-D projection

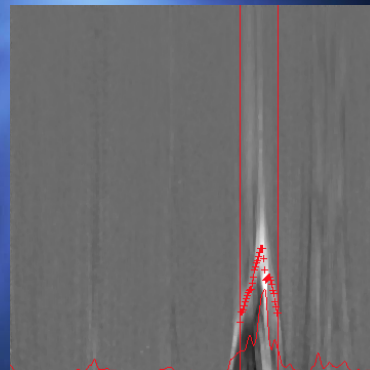
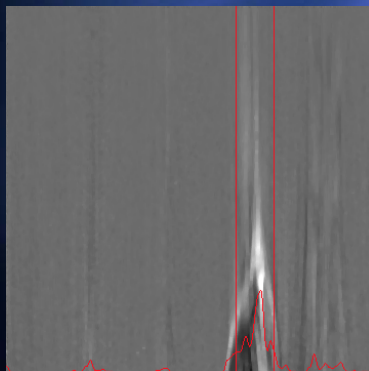


CME core angles found with thresholding

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Initial Detection

- Find CME full angles
 - ⊕ Region Growth
 - ⊕ Closing (Dilation + Erosion):
join features with narrow gaps
 - ⊕ Opening (Erosion + Dilation):
remove narrow features



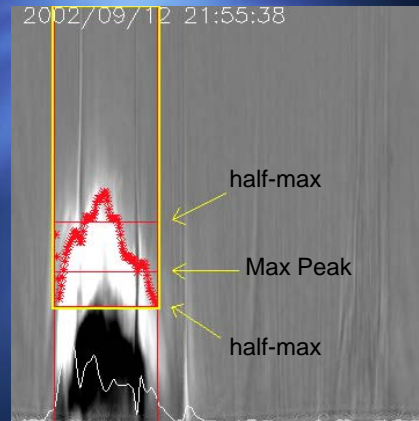
Find CME Outline with threshold segmentation

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CME Tracking

Setting tracking box (yellow):

- ⊕ Project image within found angles onto the y-axis and find the max peak, when tracking, this height must be greater than previous height detection.
- ⊕ Tracking box lower limit defined by finding the half-max below the max peak. Upper limit is the edge of the field of view.



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CME Trailing

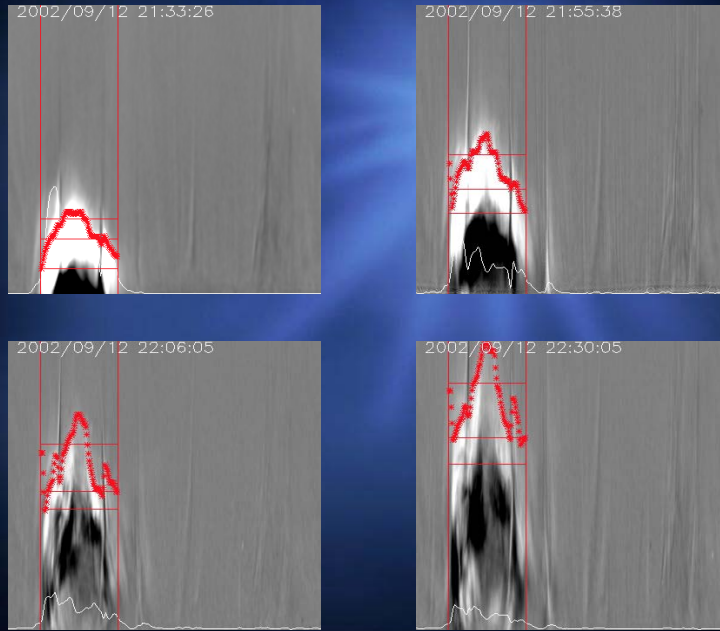
Setting trailing box (yellow):

- ⊕ Some filament cores are brighter than the Leading Edge (LE)
- ⊕ Remove the gusty outflow following a CME
- ⊕ Some outflow blobs are brighter than the LE
- ⊕ Initially re-setting all pixels in box to zero
- ⊕ Gradually return, in time, pixel value to normal level
 - $F_m = F_0 * (1 - \exp(-(T-T_0)/T_m))$
 - ⊕ Exponential modulation function is used
 - ⊕ Suppressing blobs
 - ⊕ But revealing a new CME

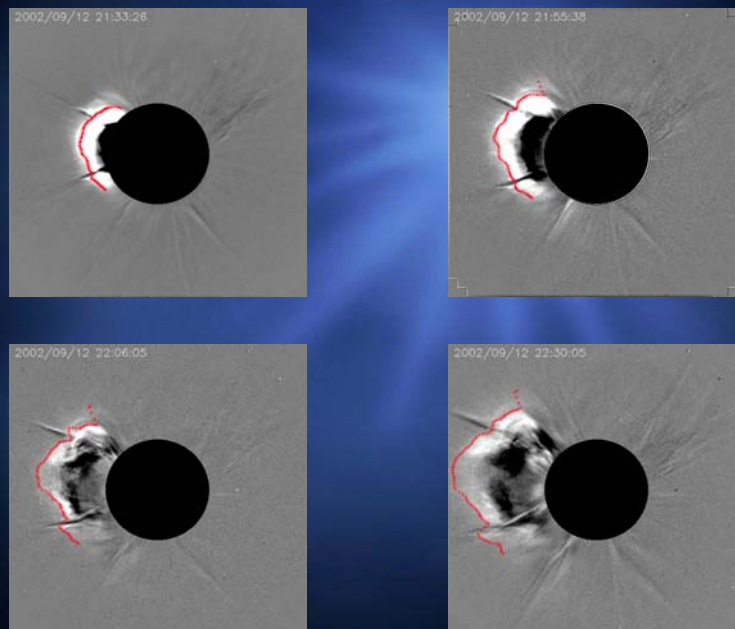


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Example Tracking Sequence



Example Sequence (Projection of edge onto Helio image)



Statistics

- ⊕ For one month data in 2002 September
 - ⊕ SEEDS --- 253 events
 - ⊕ CACTUS^[1] --- 409 event
 - ⊕ GSFC/NRL^[2] --- 162 events

- ⊕ Assuming GSFC/NRL catalog ground-truth
 - ⊕ SEEDS detects 76.5% (True Positive rate) CMEs (124/162)
 - ⊕ SEEDS misses 23.4% (False Negative rate) CMEs (38/162)

- ⊕ SEEDS finds 79.6% more events than GSFC/NRL (129/162)

- ⊕ GMU ground-truth (for the 253 events)
 - ⊕ CME: 58.9% (149/253)
 - ⊕ Blob/outflow: 23.7% (60/253)
 - ⊕ Noise Feature: 17.4% (44/253)

[1] E. Robbrecht & D. Berghmans, "Automated recognition of coronal mass ejections (CMEs) in near-real-time data", A&A 425, 1097 (2004)
<http://sidc.oma.be/cactus/publi/>

[2] CME catalog is generated and maintained at the CDAW Data Center by NASA and The Catholic University of America in cooperation with the Naval Research Laboratory. SOHO is a project of international cooperation between ESA and NASA.
http://cdaw.gsfc.nasa.gov/CME_list/

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The Next Stage

- ⊕ Formalize the C2 detection
- ⊕ Extend to C3 detection
- ⊕ Merging C2/C3 data
- ⊕ Detecting dimming/flaring features in EIT
- ⊕ Classification using data mining tools
- ⊕ Association using data mining tools:
 - ⊕ automatically identify solar surface source region of a CME
- ⊕ A full online catalog

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Discussion and the Future

- ⊕ Apply the tools with twin SECCHI instruments plus SOHO LASCO/EIT
- ⊕ Determine the true CME velocity and moving direction in the inner heliosphere in 3-D
- ⊕ Fully develop the image processing, data mining, 3-D tracking tools for automatic detection/tracking of solar eruptive events
- ⊕ Need real time tools for space weather forecasting

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